

Mendeleev's Table (145 years ago, 63 elements)

Handwritten notes in Cyrillic:
 Порядок элементов...
 D. Mendeleev.

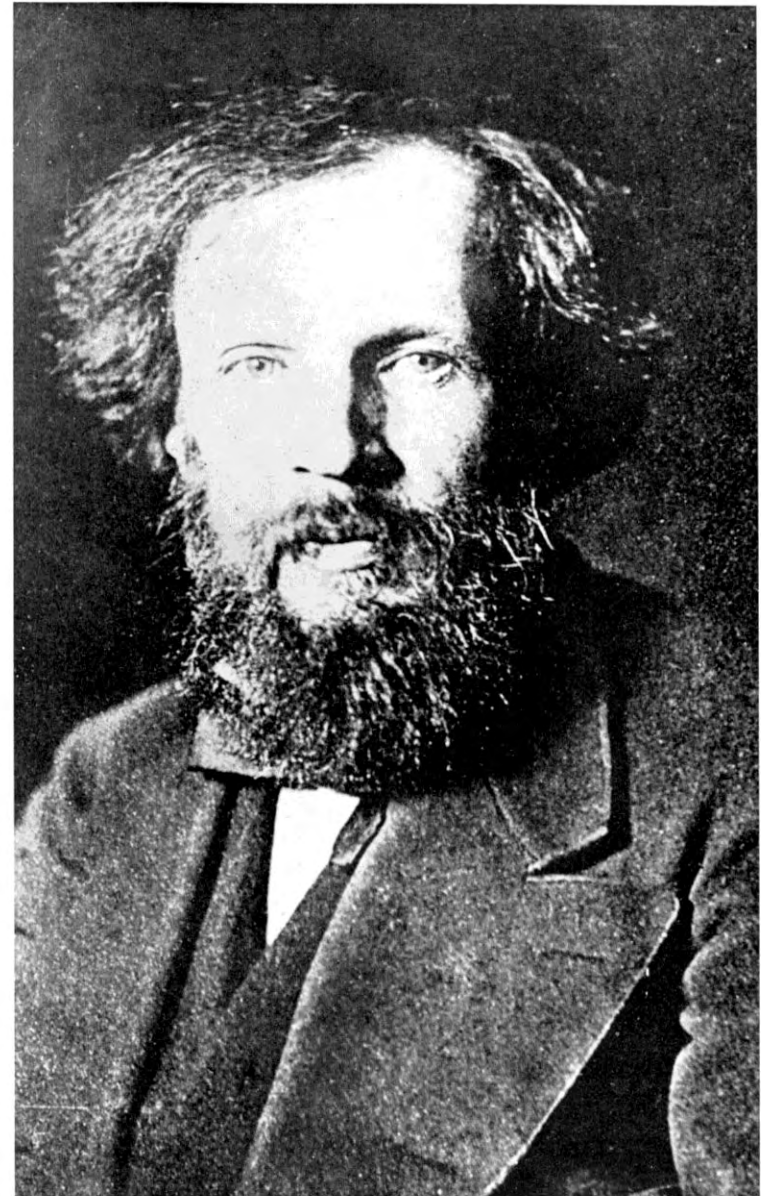
	Li=7	Be=9	B=11	C=12	N=14	O=16	F=19	Na=23	Mg=24	Al=27	Si=28	P=31	S=32	Cl=35.5	K=39	Ca=40	Sc=45	Ti=48	V=51	Cr=52	Mn=55	Fe=56	Ni=59	Cu=63.4	Zn=65.2	Ga=69.7	Ge=72	As=75	Se=78	Br=80	Kr=83.4	Rb=85.4	Sr=87.6	Zr=91.2	Nb=92.9	Mo=95.9	Ru=101.4	Rh=102.9	Pd=106.4	Ag=107.87	Cd=112.4	In=114.8	Sn=118.7	Sb=121.75	Te=127.6	I=126.9	Xe=131.3	Ba=137.33	La=138.905	Ce=140.12	Pr=140.908	Nd=144.24	Pm=144.9126	Sm=150.35	Eu=151.964	Gd=157.25	Tb=158.92532	Dy=162.50	Ho=164.93032	Er=167.259	Tm=168.93002	Yb=173.054	Lu=174.967	Hf=178.49	Ta=180.94788	W=183.84	Re=186.207	Os=190.23	Ir=192.222	Pt=195.084	Au=196.966569	Hg=200.59	Tl=204.387	Pb=207.2	Bi=208.9804	Po=209	At=210	Rn=222
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Vertical handwritten note on the left:
 Мандельцев
 1869

Essai d'une système des éléments d'après leurs poids atomiques et fonctions chimiques par D. Mendeleeff.

18 II 69.

Bottom handwritten note:
 Андрей Иванович Менделѣевъ



Mendeleev's Table

период	ряд	группы элементов															
		a I б	a II б	a III б	a IV б	a V б	a VI б	a VII б	a VIII б								
1	I	Водород H 1,00794 Hydrogen 1s ¹														Гелий He 4,0026 Helium 1s ²	
2	II	Литий Li 6,941 Lithium 2s ¹	Бериллий Be 9,012182 Beryllium 2s ²	Бор B 10,811 Boron 2p ¹	Углерод C 12,011 Carbon 2p ²	Азот N 14,00674 Nitrogen 2p ³	Кислород O 15,9994 Oxygen 2p ⁴	Фтор F 18,9984032 Fluorine 2p ⁵	Неон Ne 20,1797 Neon 2p ⁶								
3	III	Натрий Na 22,989768 Sodium 3s ¹	Магний Mg 24,3050 Magnesium 3s ²	Алюминий Al 26,981539 Aluminum 3p ¹	Кремний Si 28,0855 Silicon 3p ²	Фосфор P 30,973762 Phosphorus 3p ³	Сера S 32,066 Sulfur 3p ⁴	Хлор Cl 35,4527 Chlorine 3p ⁵	Аргон Ar 39,948 Argon 3p ⁶								
4	IV	Калий K 39,0983 Potassium 4s ¹	Кальций Ca 40,078 Calcium 4s ²	21 Скандий Sc 44,955910 Scandium 3d ¹ 4s ²	22 Титан Ti 47,88 Titanium 3d ² 4s ²	23 Ванадий V 50,9415 Vanadium 3d ³ 4s ²	24 Хром Cr 51,9961 Chromium 3d ⁵ 4s ¹	25 Марганец Mn 54,93805 Manganese 3d ⁵ 4s ²	26 Железо Fe 55,847 Iron 3d ⁶ 4s ²	27 Кобальт Co 58,93320 Cobalt 3d ⁷ 4s ²	28 Никель Ni 58,6934 Nickel 3d ⁸ 4s ²						
	V	29 Медь Cu 63,546 Copper 3d ¹⁰ 4s ¹	30 Цинк Zn 65,39 Zinc 3d ¹⁰ 4s ²	Галлий Ga 69,723 Gallium 4p ¹	Германий Ge 72,61 Germanium 4p ²	Мышьяк As 74,92159 Arsenic 4p ³	Селен Se 78,96 Selenium 4p ⁴	Бром Br 79,904 Bromine 4p ⁵	Криптон Kr 83,80 Krypton 4p ⁶								
5	VI	Рубидий Rb 85,4678 Rubidium 5s ¹	Стронций Sr 87,62 Strontium 5s ²	39 Иттрий Y 88,90585 Yttrium 4d ¹ 5s ²	40 Цирконий Zr 91,224 Zirconium 4d ² 5s ²	41 Ниобий Nb 92,90638 Niobium 4d ⁴ 5s ¹	42 Молибден Mo 95,94 Molybdenum 4d ⁵ 5s ¹	43 Технеций Tc [98] Technetium 4d ⁵ 5s ²	44 Рутений Ru 101,07 Ruthenium 4d ⁶ 5s ¹	45 Родий Rh 102,90550 Rhodium 4d ⁷ 5s ¹	46 Палладий Pd 106,42 Palladium 4d ¹⁰						
	VII	47 Серебро Ag 107,8682 Silver 4d ¹⁰ 5s ¹	48 Кадмий Cd 112,411 Cadmium 4d ¹⁰ 5s ²	Индий In 114,818 Indium 5p ¹	Олово Sn 118,710 Tin 5p ²	Сурьма Sb 121,757 Antimony 5p ³	Теллур Te 127,60 Tellurium 5p ⁴	Иод I 126,90447 Iodine 5p ⁵	Ксенон Xe 131,29 Xenon 5p ⁶								
6	VIII	Цезий Cs 132,90543 Cesium 6s ¹	Барий Ba 137,327 Barium 6s ²	57 Лантан La 138,9055 Lanthanum 5d ¹ 6s ²	72 Гафний Hf 178,49 Hafnium 5d ² 6s ²	73 Тантал Ta 180,9479 Tantalum 5d ³ 6s ²	74 Вольфрам W 183,84 Tungsten 5d ⁴ 6s ²	75 Рений Re 186,207 Rhenium 5d ⁵ 6s ²	76 Осмий Os 190,23 Osmium 5d ⁶ 6s ²	77 Иридий Ir 192,22 Iridium 5d ⁷ 6s ²	78 Платина Pt 195,08 Platinum 5d ⁹ 6s ¹						
	IX	79 Золото Au 196,96654 Gold 5d ¹⁰ 6s ¹	80 Ртуть Hg 200,59 Mercury 5d ¹⁰ 6s ²	Таллий Tl 204,3833 Thallium 6p ¹	Свинец Pb 207,2 Lead 6p ²	Висмут Bi 208,98037 Bismuth 6p ³	Полоний Po [209] Polonium 6p ⁴	Астат At [210] Astatine 6p ⁵	Радон Rn [222] Radon 6p ⁶								
7	X	Франций Fr [223] Francium 7s ¹	Радий Ra 226,025 Radium 7s ²	89 Актиний Ac [227] Actinium 6d ¹ 7s ²	104 Резерфордий Rf [261] Rutherfordium	105 Дубний Db [262] Dubnium	106 Сиборгий Sg [266] Seaborgium	107 Борий Bh [267] Bohrium	108 Хасий Hs [269] Hassium	109 Мейтнерий Mt [268] Meitnerium	110 Дармштадтий Ds [269] Darmstadtium						
	XI	111 Рентгений Rg [269] Roentgenium	112	113	Fl 114	115	Lv 116	117	118								



- s-элементы
- p-элементы
- d-элементы
- f-элементы

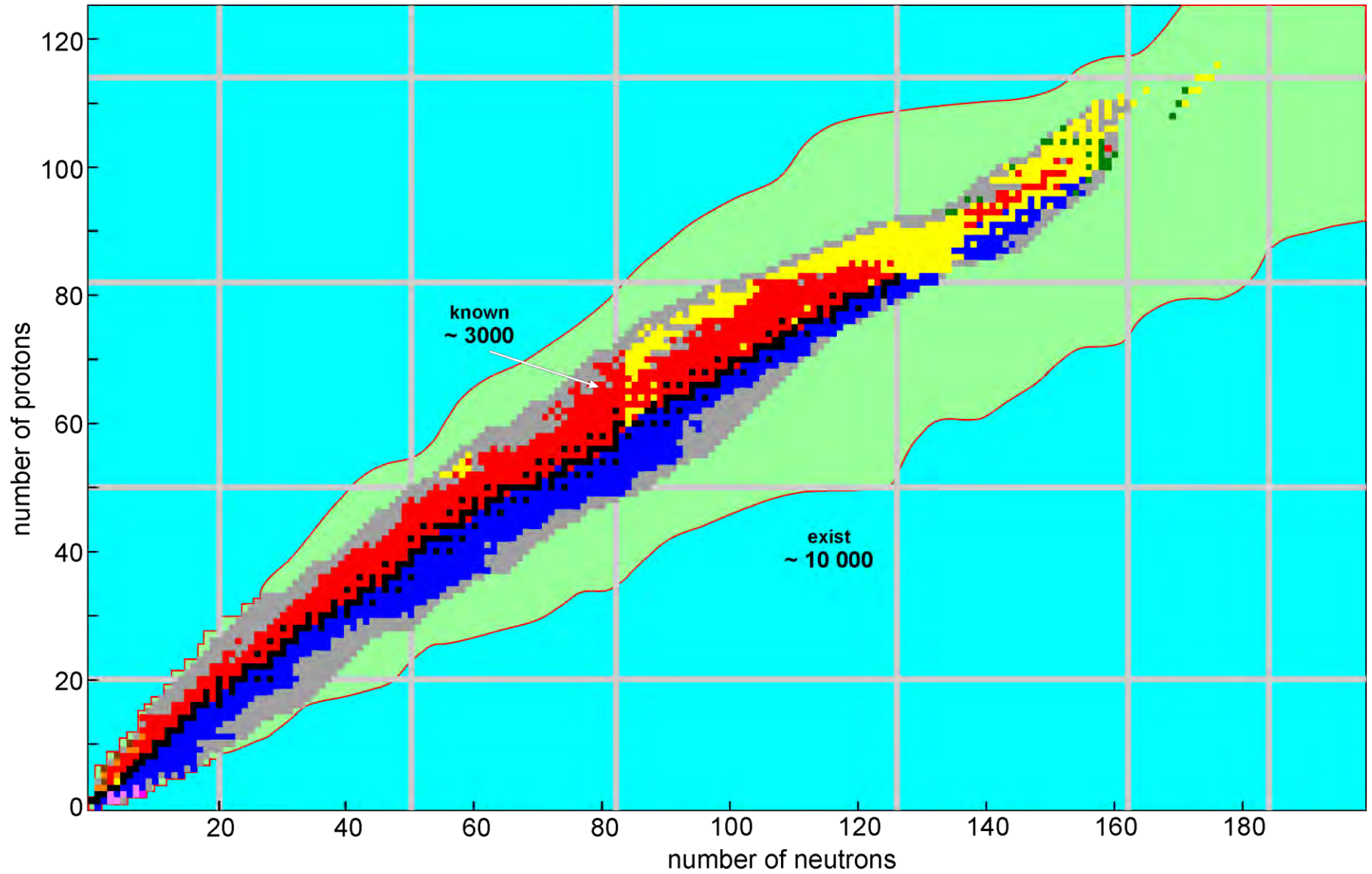
Лантаноиды Lanthanides

Церий Ce 4f ¹ 5d ¹ 140,115 Cerium	Празеодим Pr 4f ³ 140,90765 Praseodymium	Неодим Nd 4f ⁴ 144,24 Neodymium	Прометий Pm 4f ⁵ [145] Promethium	Самарий Sm 4f ⁶ 150,36 Samarium	Европий Eu 4f ⁷ 151,965 Europium	Гадолиний Gd 4f ⁷ 5d ¹ 157,25 Gadolinium	Тербий Tb 4f ⁹ 158,92534 Terbium	Диспрозий Dy 4f ¹⁰ 162,50 Dysprosium	Гольмий Ho 4f ¹¹ 164,93032 Holmium	Эрбий Er 4f ¹² 167,26 Erbium	Тулий Tm 4f ¹³ 168,93421 Thulium	Иттербий Yb 4f ¹⁴ 173,04 Ytterbium	Лютеций Lu 4f ¹⁴ 5d ¹ 174,967 Lutetium
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Актиноиды Actinides

Торий Th 7s ² 6d ² 232,0381 Thorium	Протактиний Pa 5f ² 6d ¹ 231,03588 Protactinium	Уран U 5f ³ 6d ¹ 238,0289 Uranium	Нептуний Np 5f ⁴ 6d ¹ [237] Neptunium	Плутоний Pu 5f ⁶ [244] Plutonium	Америций Am 5f ⁷ [243] Americium	Кюрий Cm 5f ⁷ 6d ¹ [247] Curium	Берклий Bk 5f ⁹ [247] Berkelium	Калифорний Cf 5f ¹⁰ [251] Californium	Эйнштейний Es 5f ¹¹ [252] Einsteinium	Фермий Fm 5f ¹² [257] Fermium	Менделевий Md 5f ¹³ [258] Mendelevium	Нобелий No 5f ¹⁴ [259] Nobelium	Лоуренсий Lr 5f ¹⁴ 6d ¹ [262] Lawrencium
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Nuclear Map

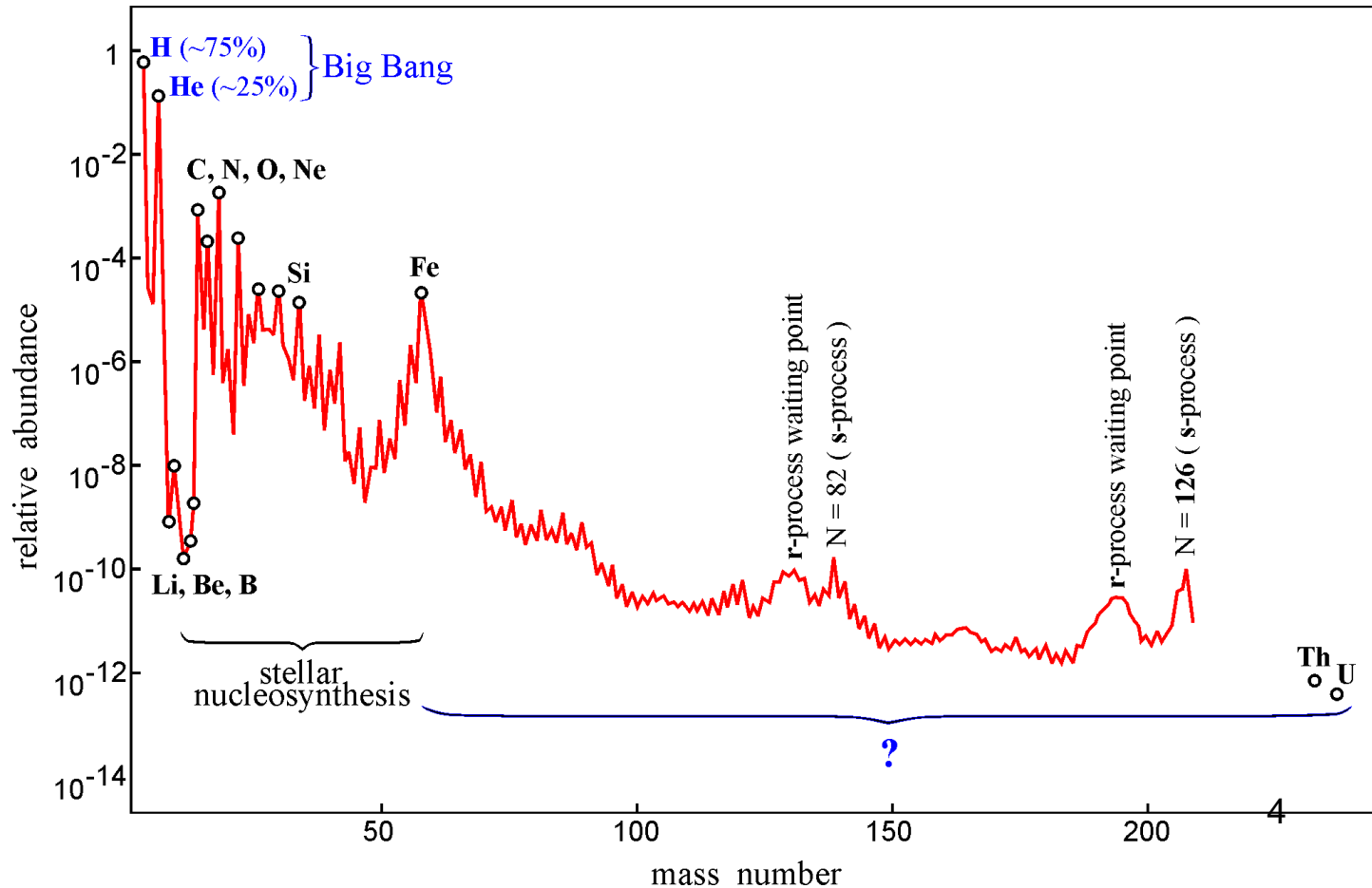


Abundance of Elements in the Universe

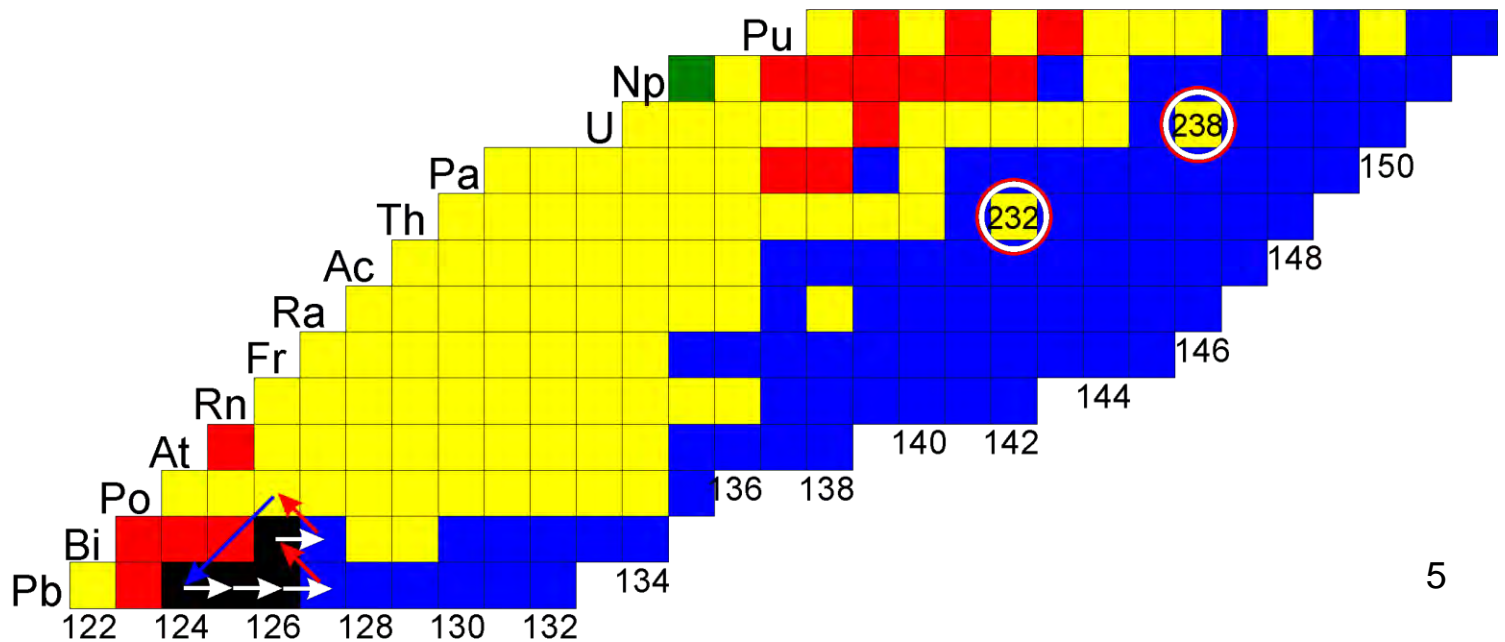
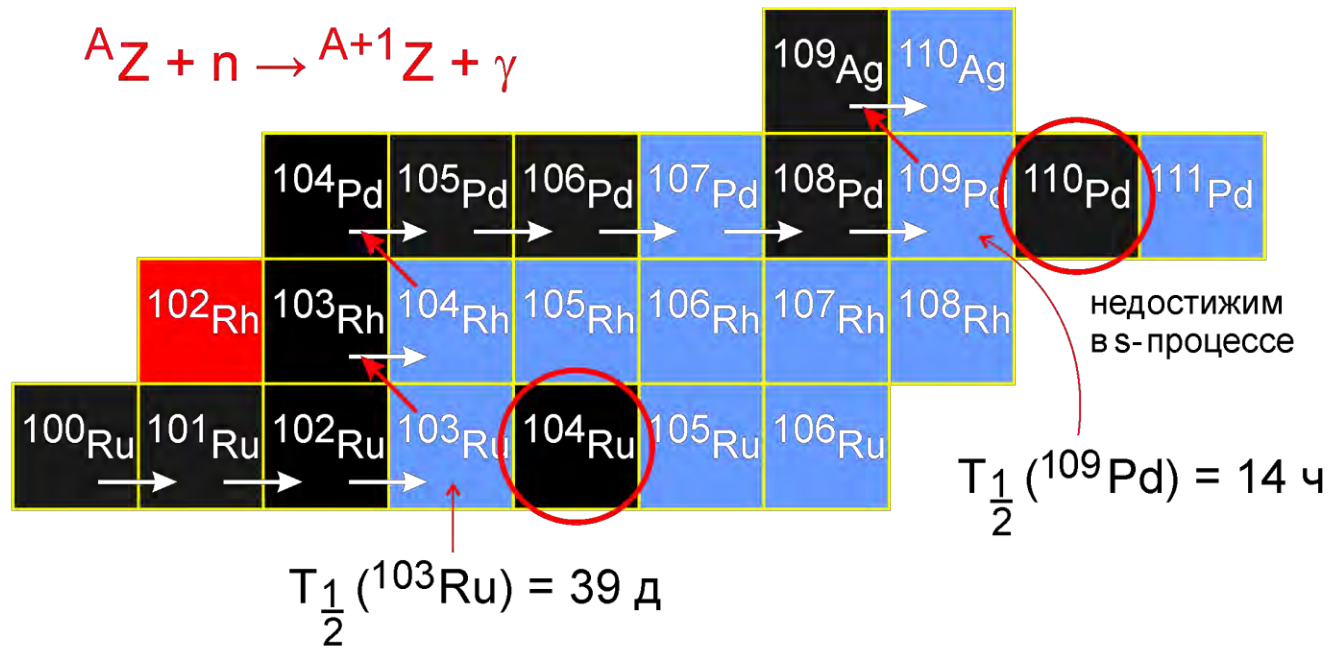
The 11 Greatest Unanswered Questions of Physics
(National Research Council, NAS, USA, 2002):

1. What is dark matter?
2. What is dark energy?
- 3. How were the heavy elements from iron to uranium made?**
4. Do neutrinos have mass?

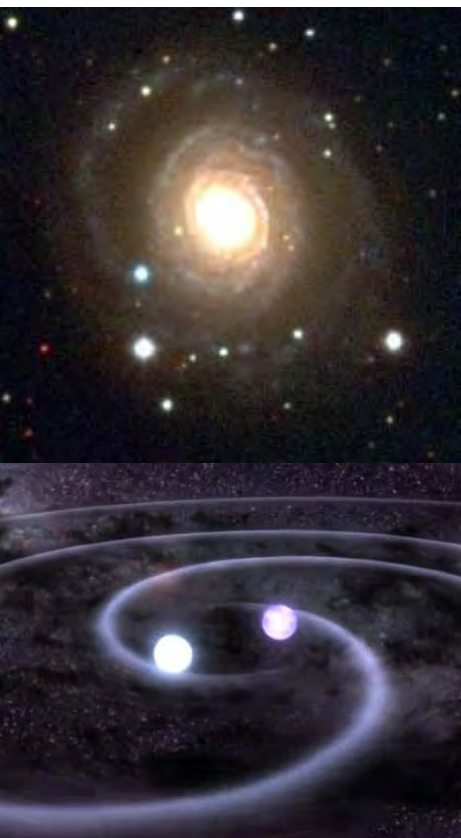
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Limitation of s process

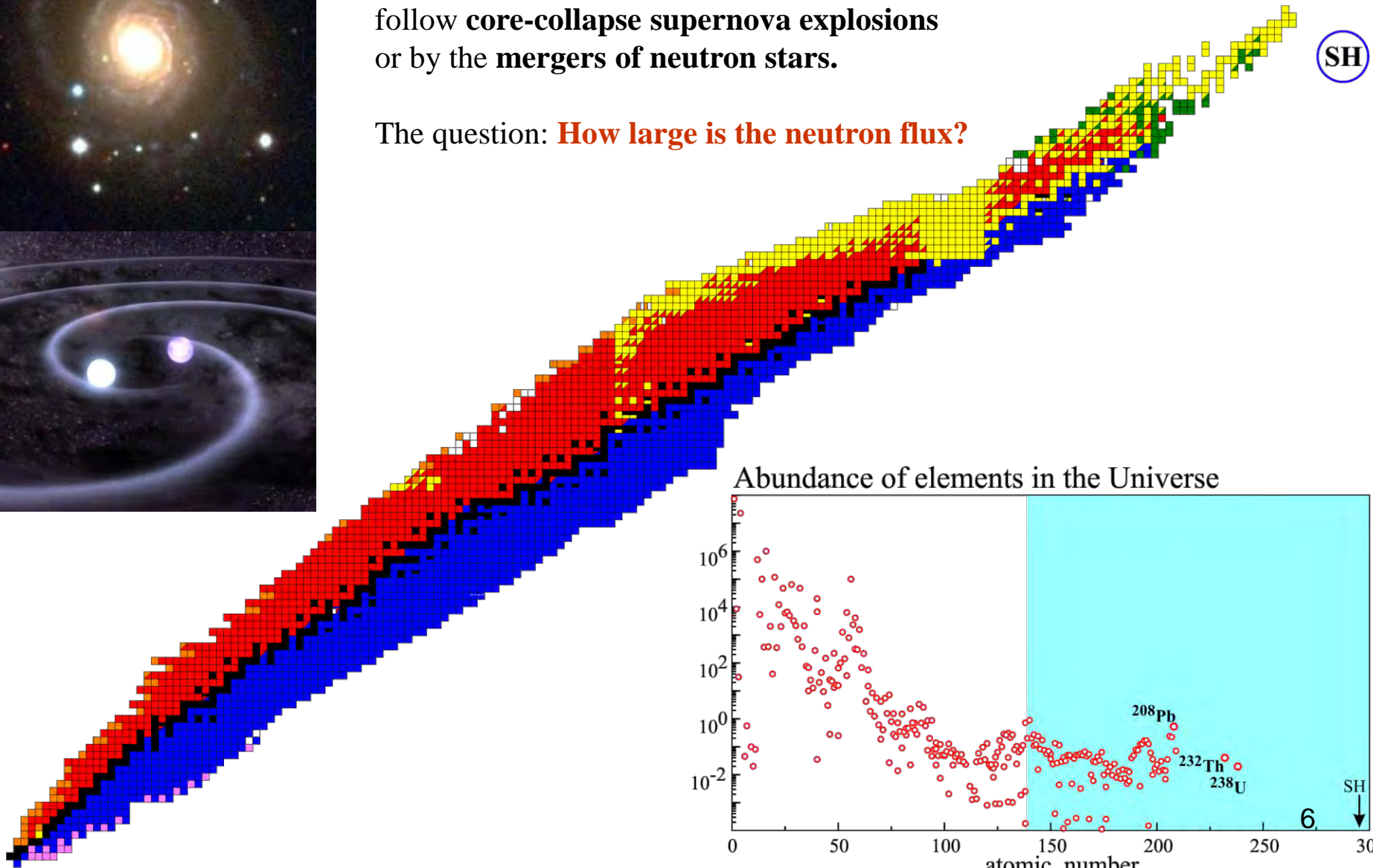


Formation of SH elements in astrophysical r-process

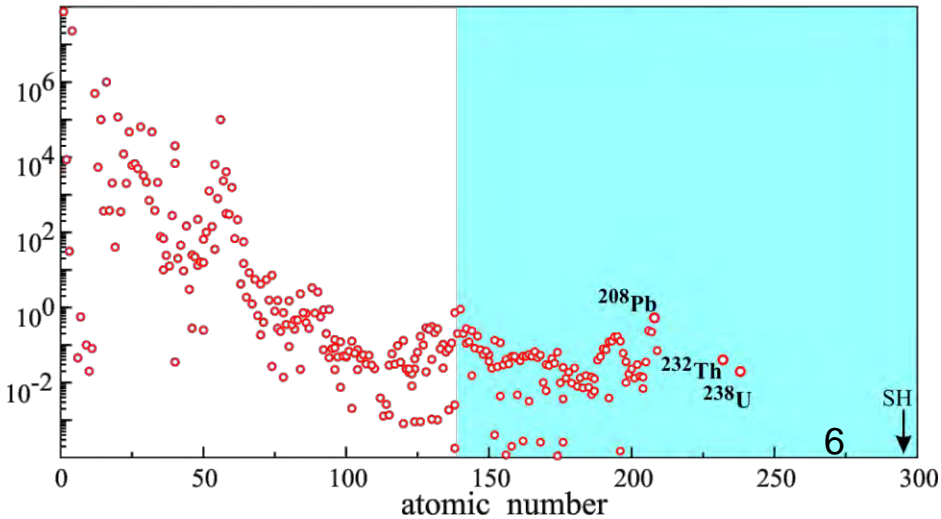


Strong neutron fluxes are expected to be generated by neutrino-driven proto-neutron star winds which follow **core-collapse supernova explosions** or by the **mergers of neutron stars**.

The question: **How large is the neutron flux?**



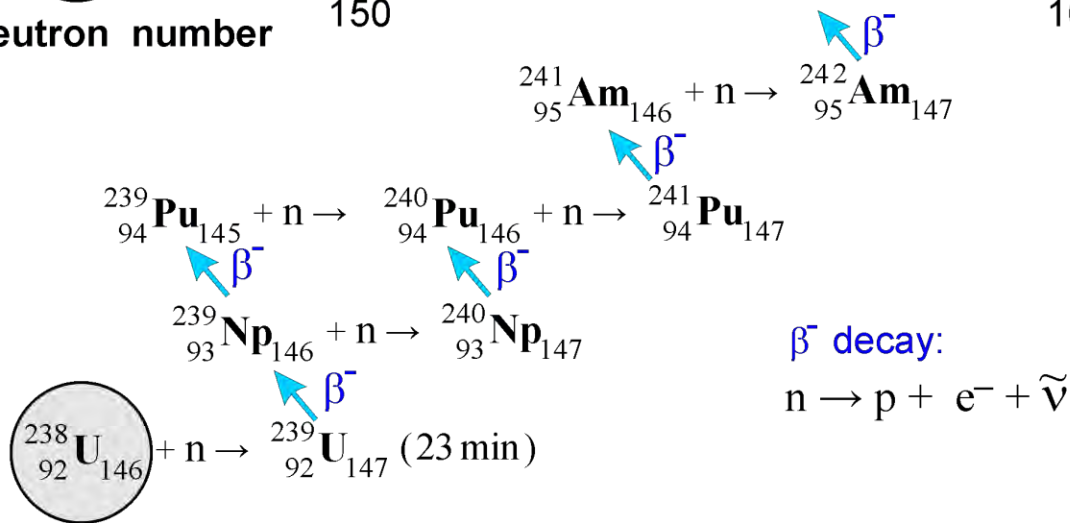
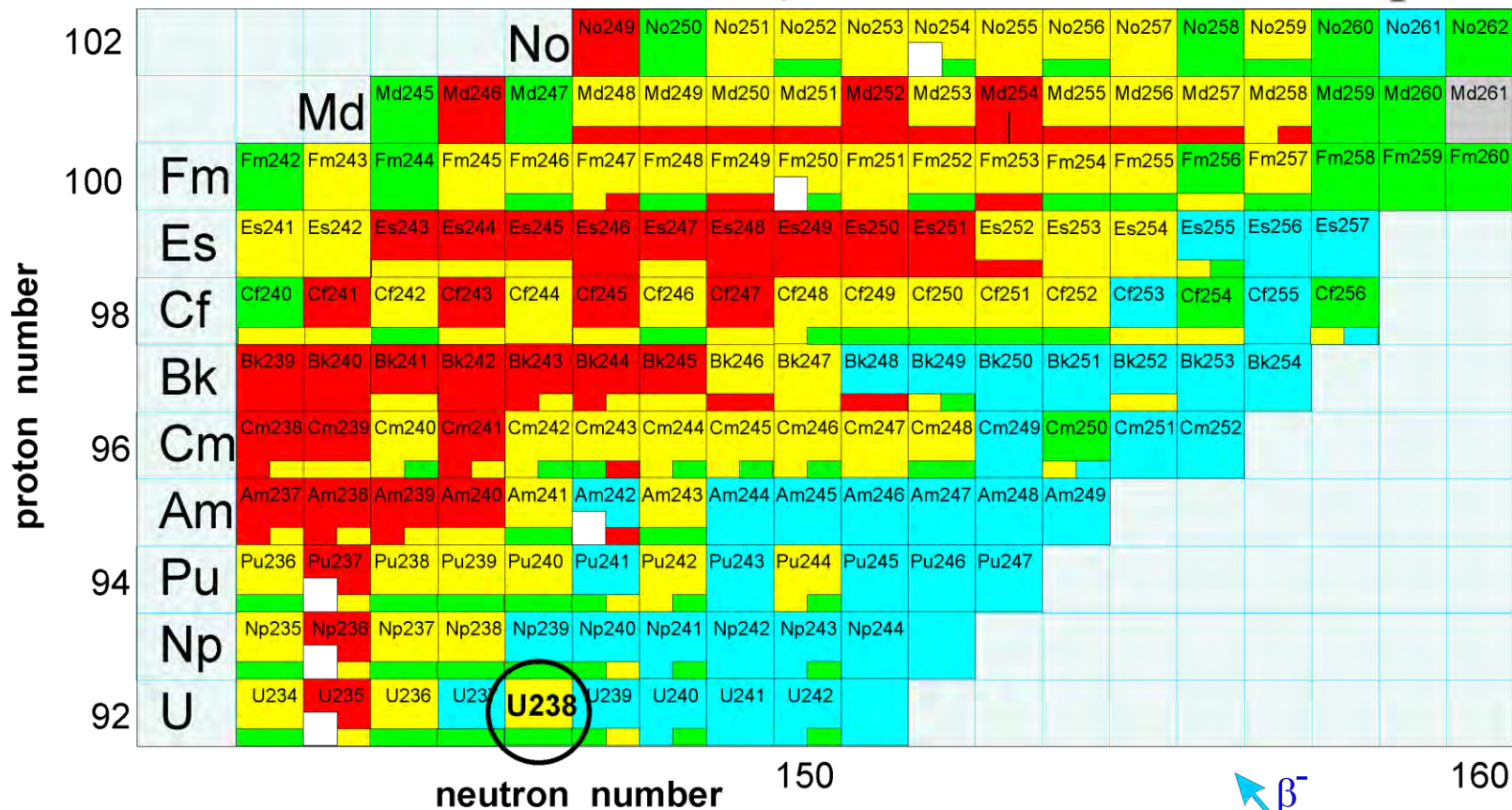
Abundance of elements in the Universe



Search for SHE in cosmic rays
1971, Dubna, P. Fowler: Tracks of SHE !?



Hand-made chemical elements (reactors and nuclear explosion)



Nucleosynthesis by neutron capture

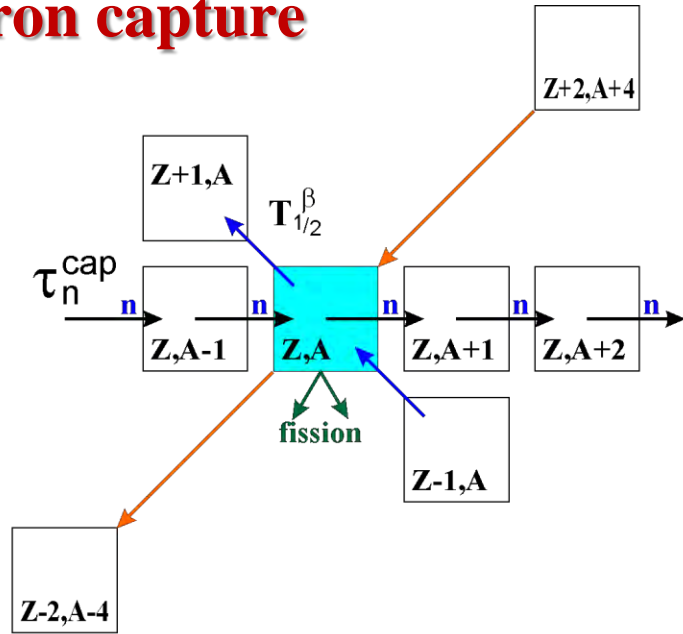
n_0 is the neutron flux
time of neutron capture

$$\tau_n^{cap} = \frac{1}{n_0 \times \sigma(n, \gamma)}$$

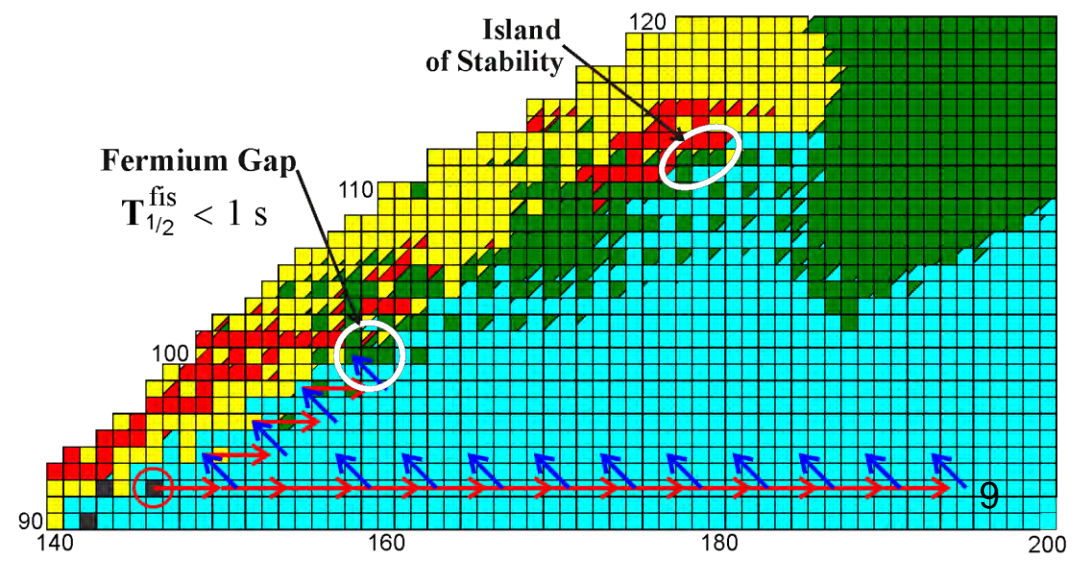
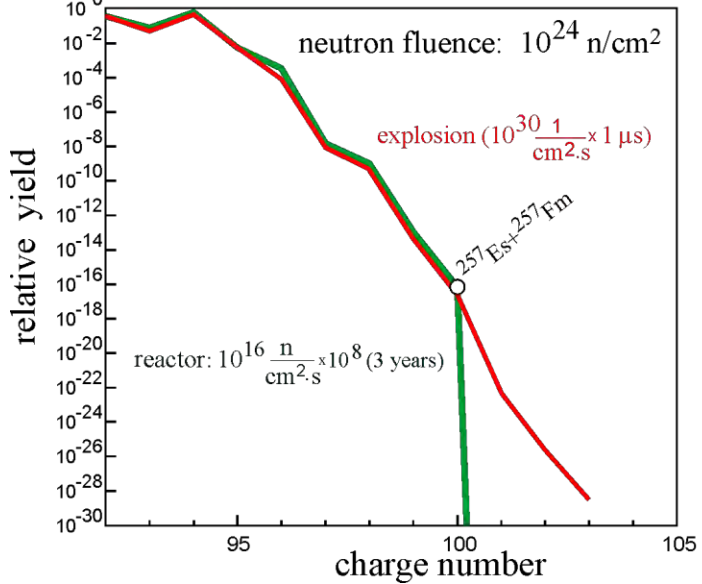
$(Z, A) \rightarrow (Z, A+1)$ if $T_{1/2} > \tau_n^{cap}$

nuclear reactor: $\tau_n^{cap} \sim 1$ year

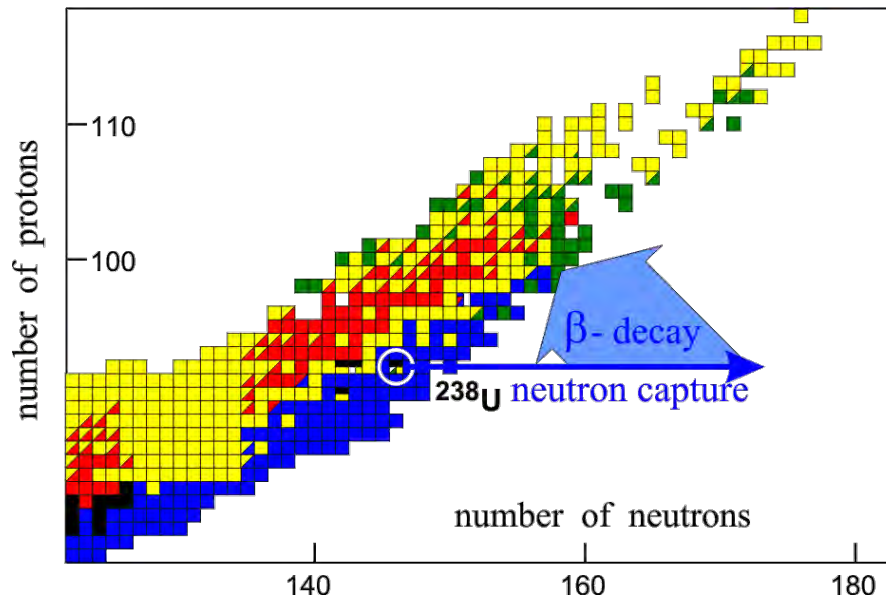
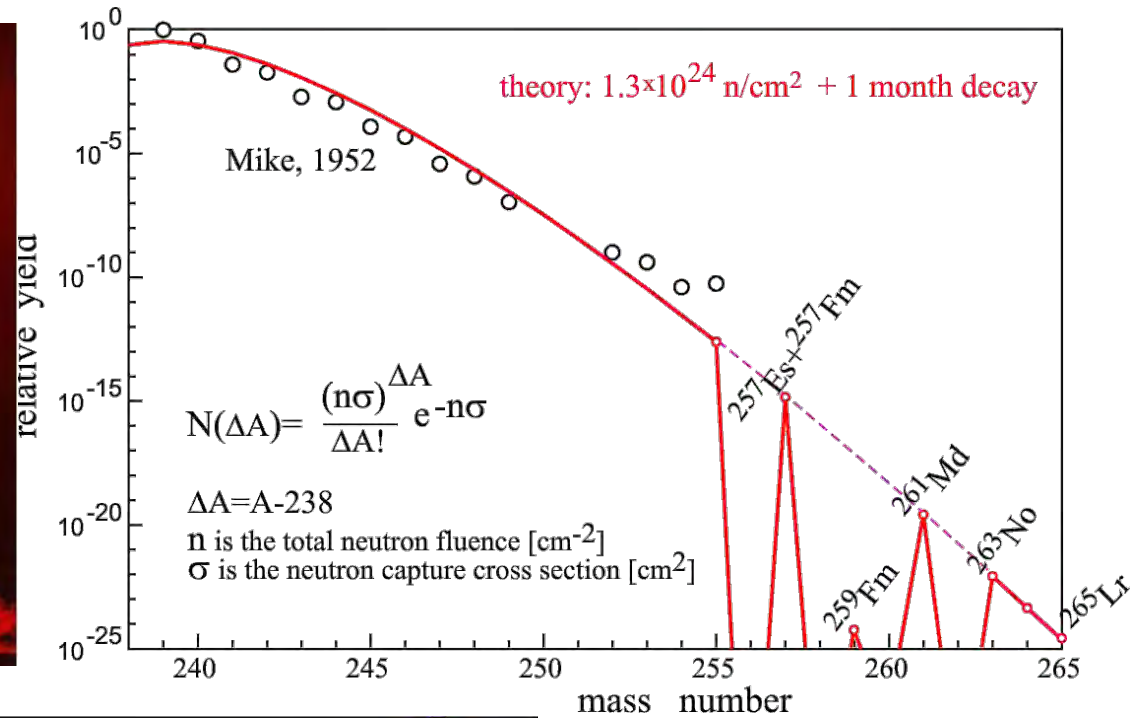
nuclear explosion: $\tau_n^{cap} \sim 1 \mu s$



$$\frac{dN_{ZA}}{dt} = N_{ZA-1} n_0 \sigma_{ZA-1}^{n\gamma} - N_{ZA} n_0 \sigma_{ZA}^{n\gamma} - N_{ZA} \frac{\ln 2}{T_{ZA}^{\beta}} - N_{ZA} \frac{\ln 2}{T_{ZA}^{\alpha}} - N_{ZA} \frac{\ln 2}{T_{ZA}^{fis}} + N_{Z-1A} \frac{\ln 2}{T_{Z-1A}^{\beta}} + N_{Z+2A+4} \frac{\ln 2}{T_{Z+2A+4}^{\alpha}}$$



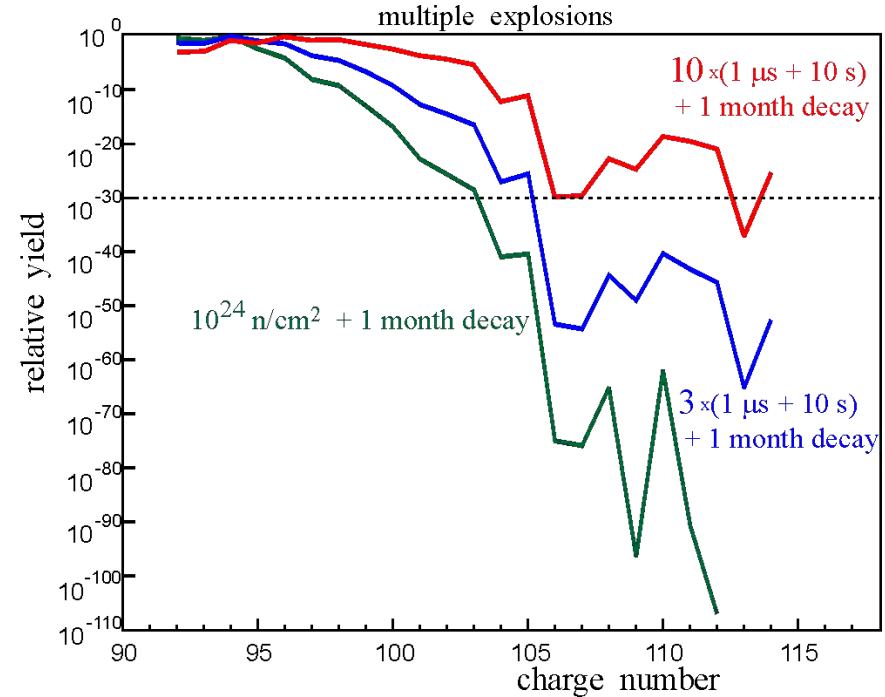
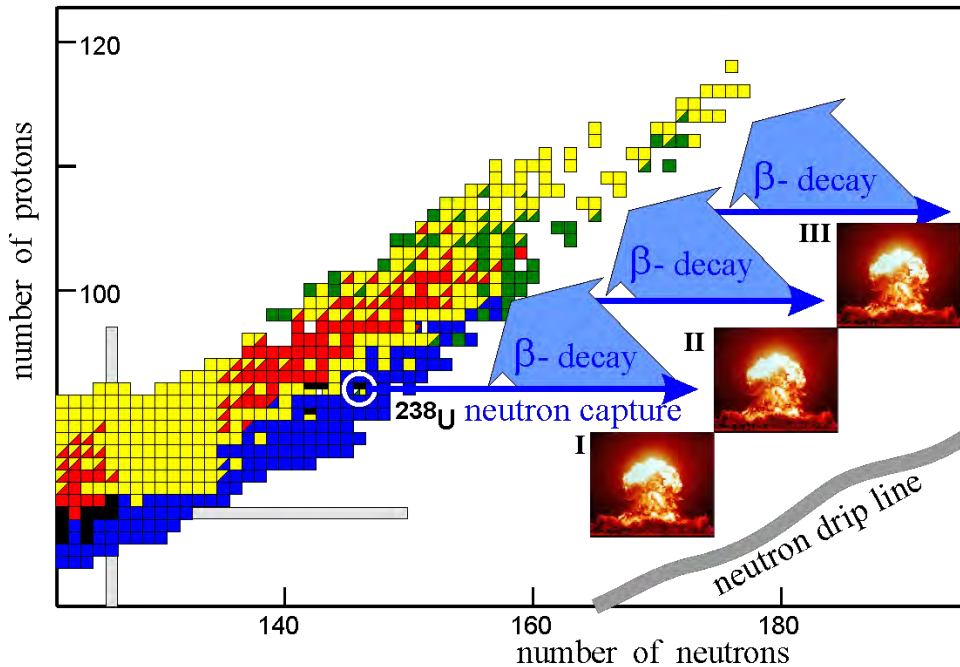
Rapid neutron capture in nuclear explosions



Multiple nuclear explosions

(proposed first by H.W. Meldner, PRL 28,1972)

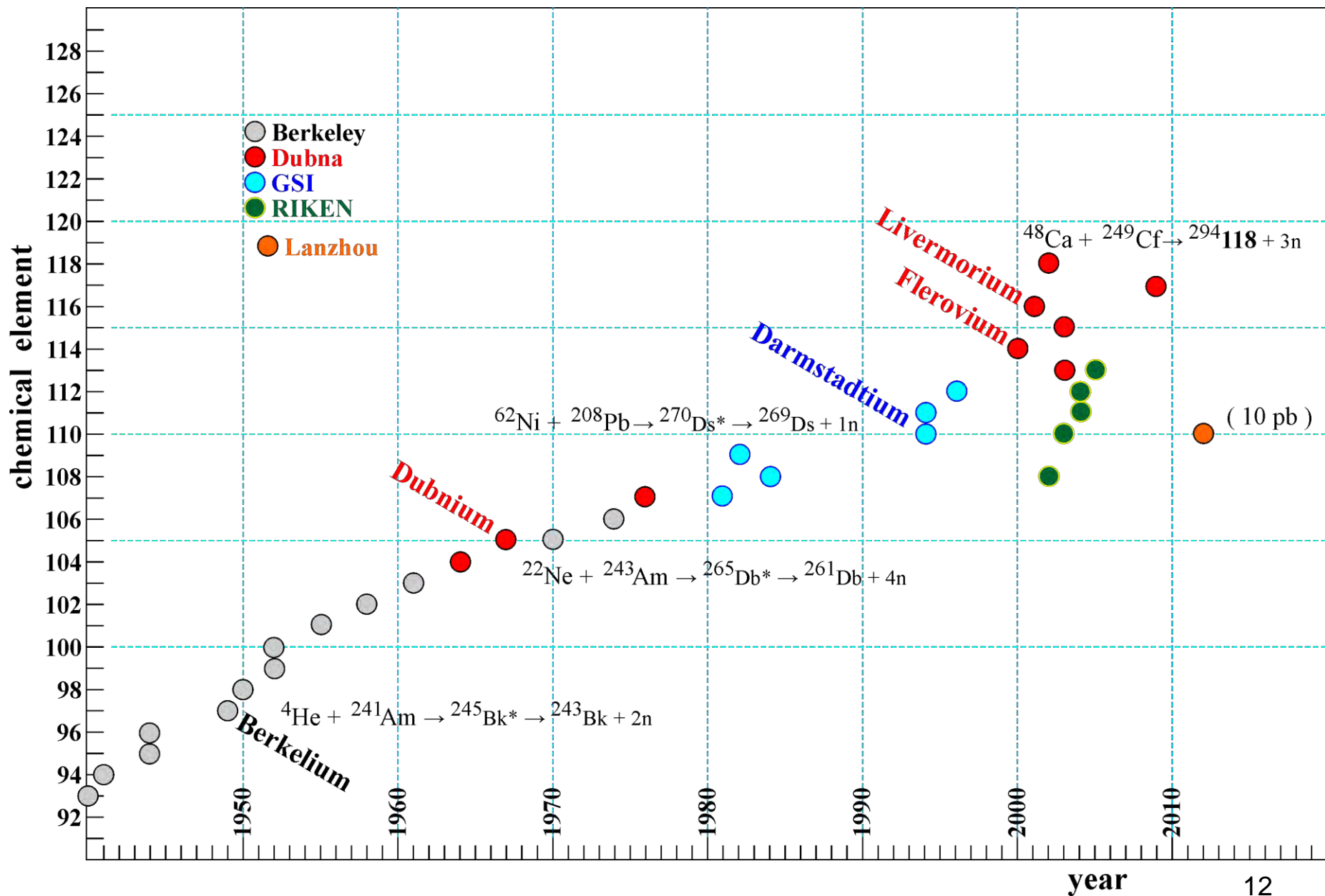
Edward Teller: Technically it is quite possible



Probability for formation of element 112
increases by **90 orders** of magnitude !

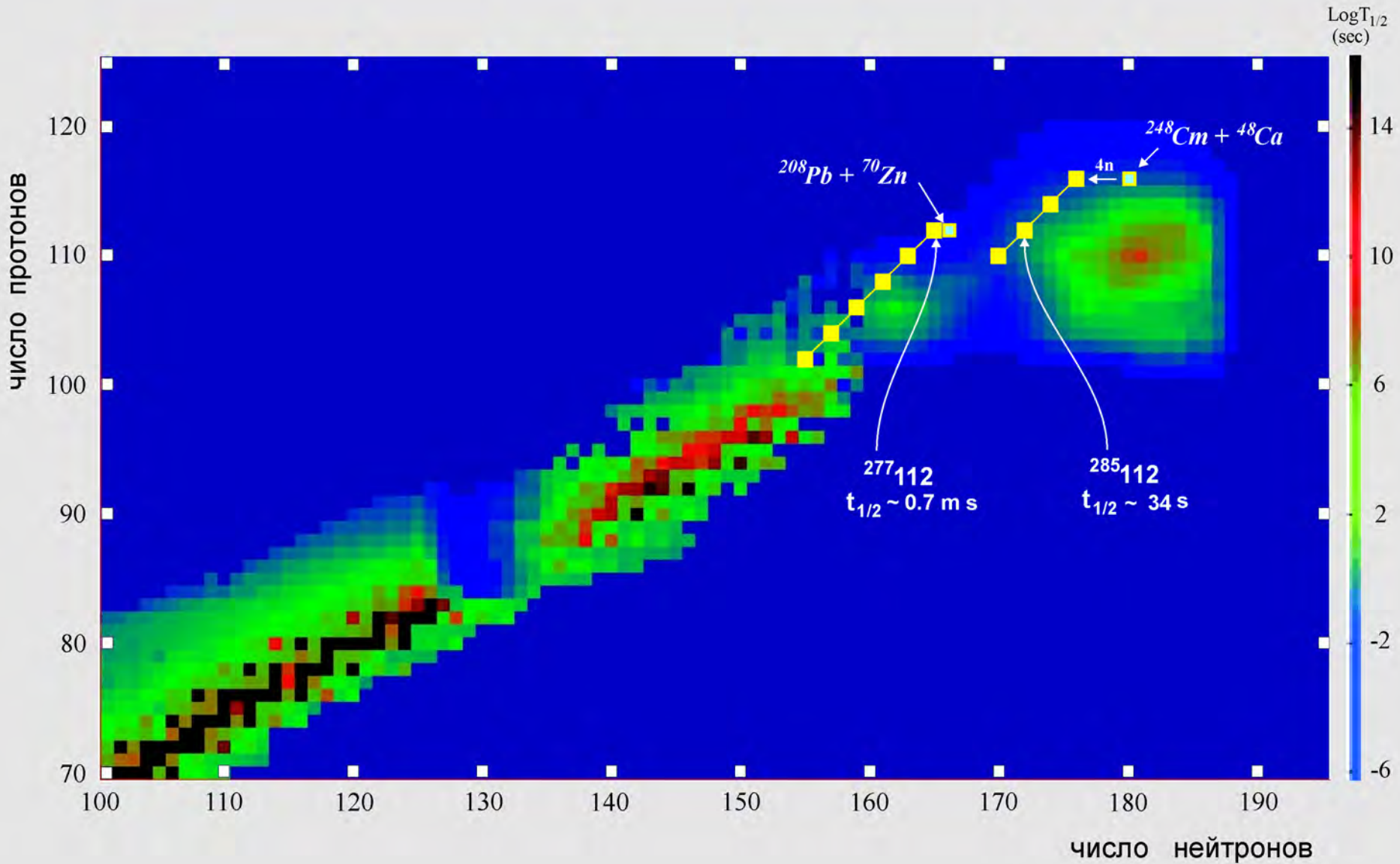
Pulsed nuclear reactors of the next generation ?

Hand-made elements (history)



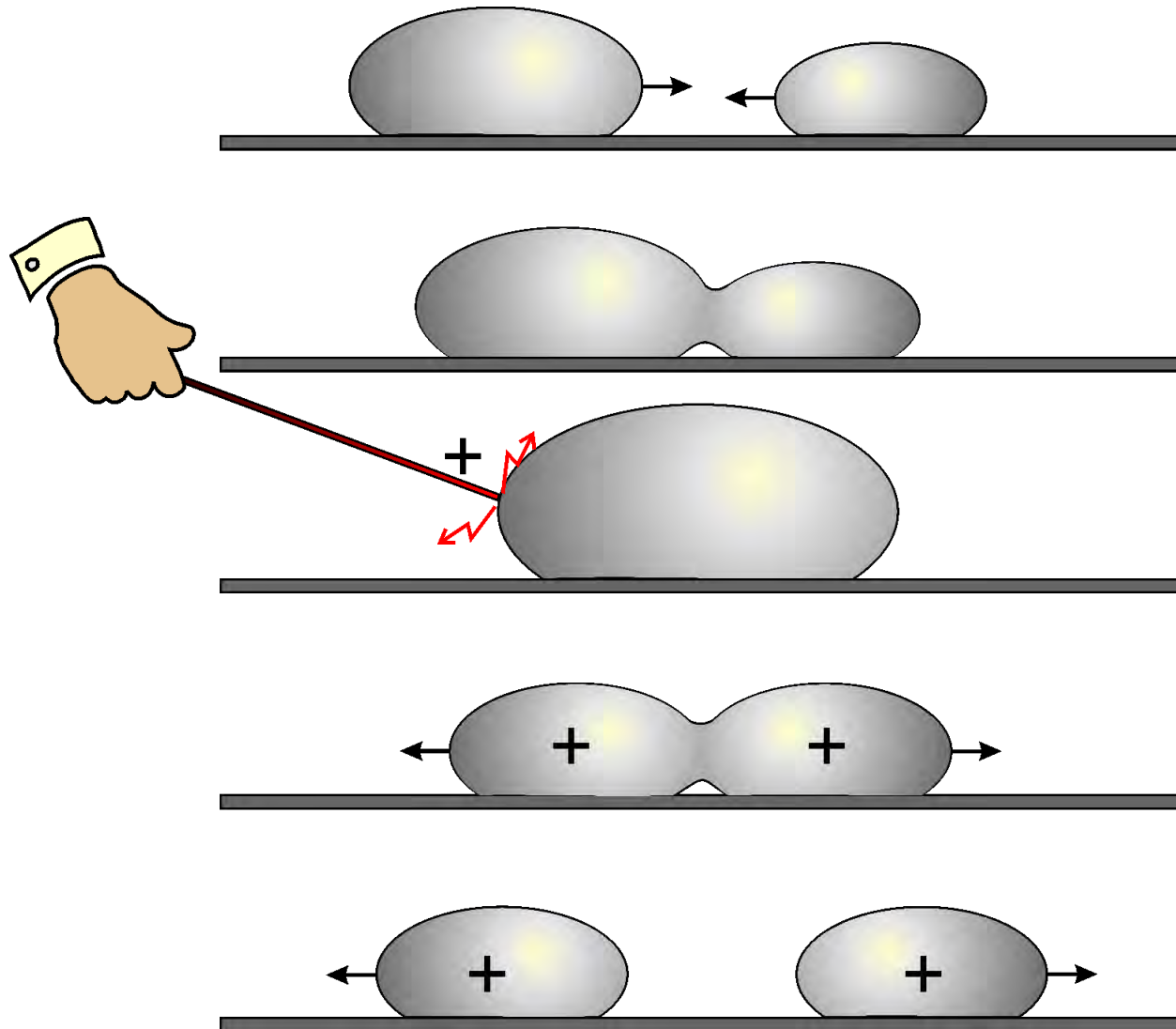
2012: a new player in the field: IMP, Lanzhou, China

Superheavy Elements (Island of Stability)



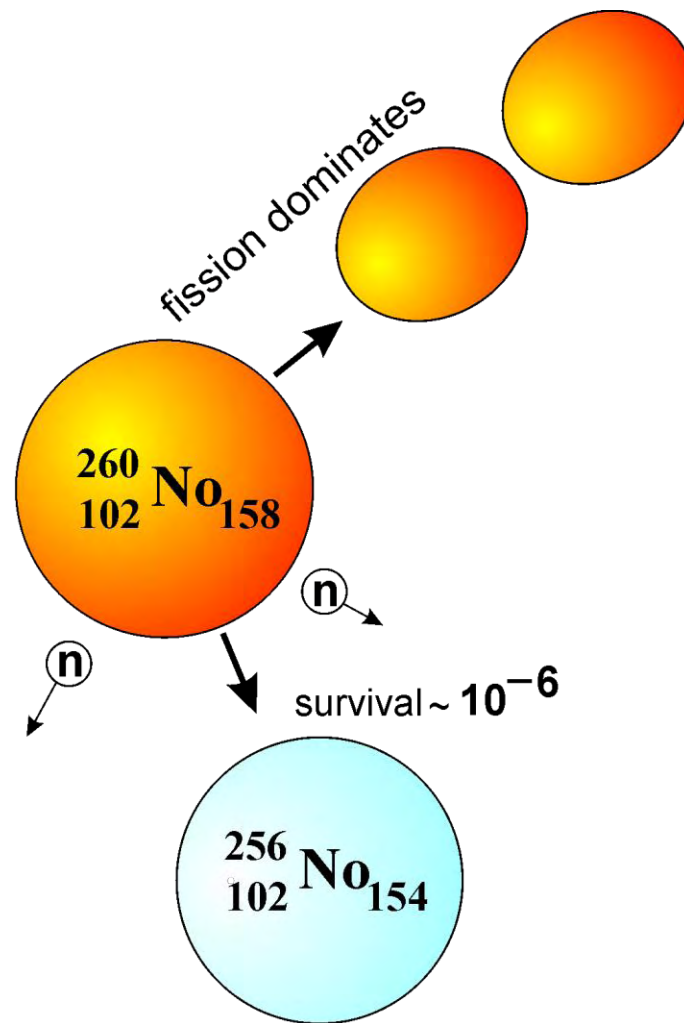
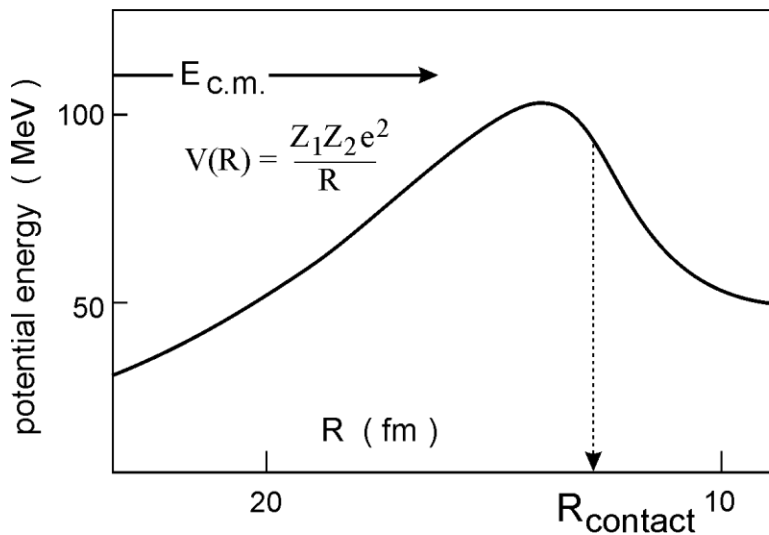
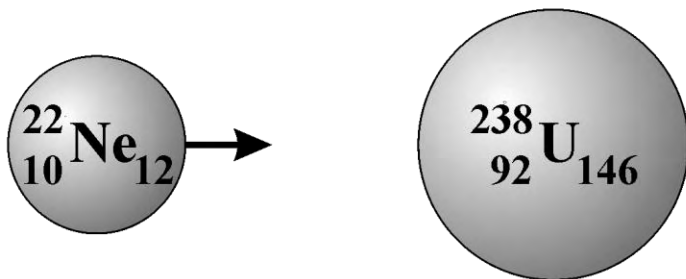
Why don't nuclei fuse ?

(example: two drops of mercury)



Synthesis of new elements in fusion reactions

acceleration up to 6 MeV/nucleon to overcome the Coulomb barrier



Physics problems

(1) How to choose appropriate fusion reaction ?

«cold» synthesis: $^{208}\text{Pb} + ^{64}\text{Ni}, ^{70}\text{Zn}, \dots \rightarrow ^{272}110, ^{278}112, \dots$ (GSI, Germany)

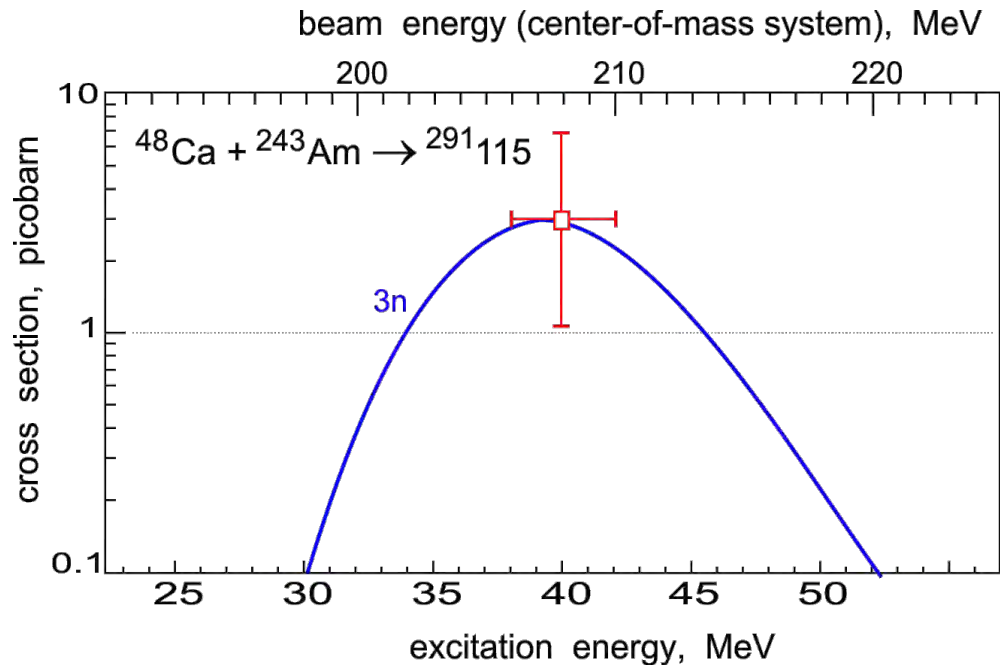
«hot» synthesis: $^{238}\text{U}, ^{244}\text{Pu}, ^{248}\text{Cm}, ^{249}\text{Cf} + ^{48}\text{Ca} \rightarrow ^{286}112, ^{292}114, ^{296}116, ^{297}118$ (Dubna)

Symmetric combinations: $^{148}\text{Nd} + ^{154}\text{Sm} \rightarrow ^{302}122$?

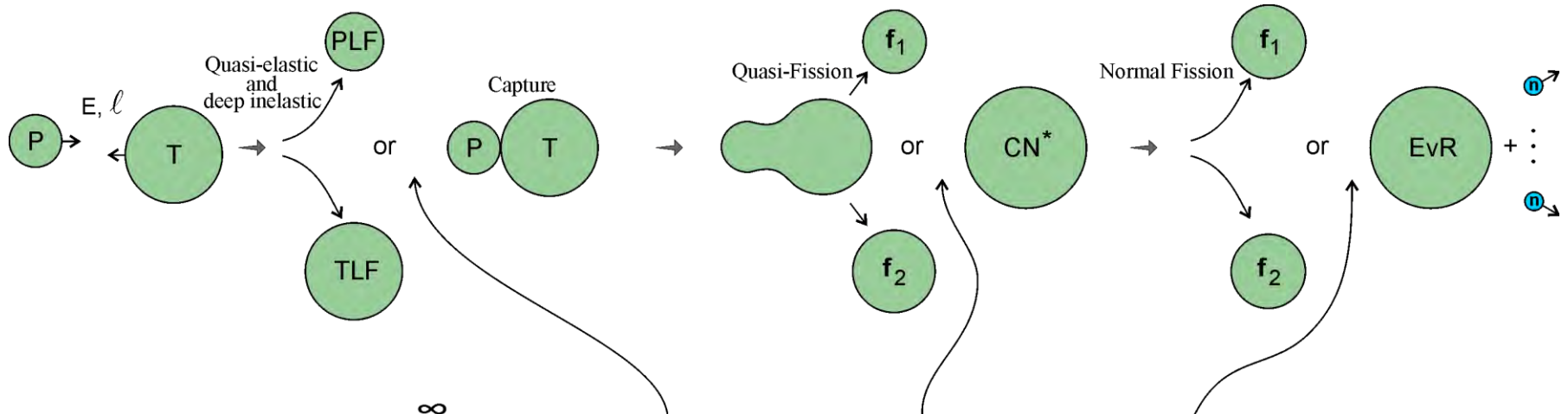
Radioactive ion beams of ^{132}Sn , etc. ?

(2) How to choose colliding energy ?

(3) How to estimate cross sections ?



Synthesis of SHE in fusion reactions (conventional view)



$$\sigma_{ER}^{xn}(E) = \frac{\pi}{k^2} \sum_{l=0}^{\infty} (2l+1) \cdot P_{\text{cont}}(l, E) \cdot P_{\text{CN}}(l, E^*) \cdot P_{\text{xn}}(l, E^*)$$

$$\sigma_{\text{capture}}(E) \stackrel{\text{exp}}{=} \sum (f_{1,2}^{\text{all}} + \text{EvR}), \text{ if } f_1 f_2 \neq \text{PLF, TLF}$$

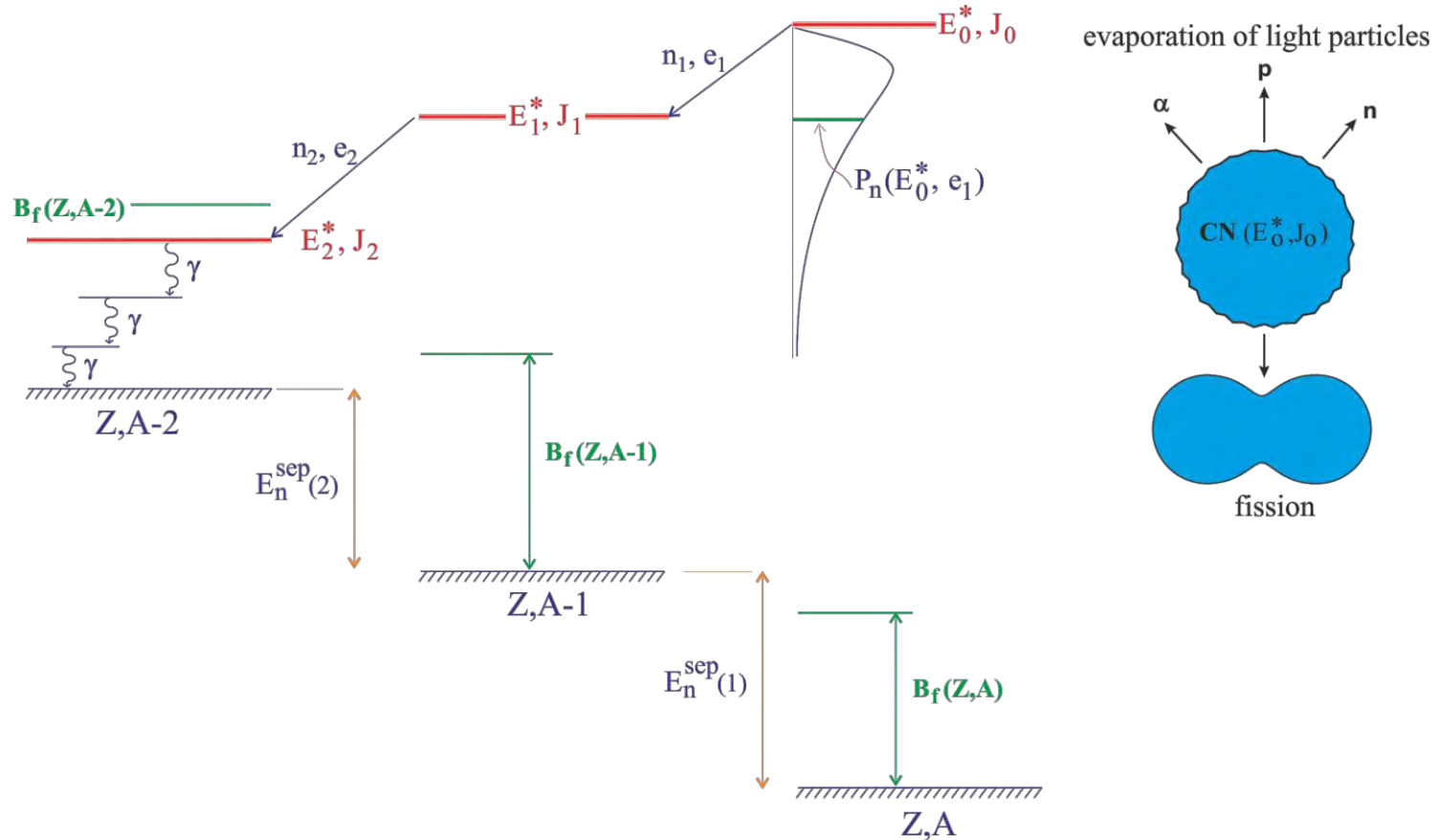
$$\sigma_{\text{CN}} \equiv \sigma_{\text{fusion}}(E) \stackrel{\text{exp}}{=} \sum (f_{1,2}^{\text{NF}} + \text{EvR}), \text{ if } f_{1,2}^{\text{NF}} \neq f_{1,2}^{\text{QF}}$$

P_{cont} : Penetration probability of the multi-dimensional Coulomb barrier $V_C^B(r, \beta_1, \beta_2, \varphi_1, \varphi_2)$ - ?

P_{CN} : Probability of CN formation (evolution in the space of shape parameters)

P_{xn} : Survival probability of excited CN (Statistical Model: $\Gamma_n, \Gamma_f, E_n^{\text{sep}}, B_{\text{fis}}$)

Decay and survival of excited compound nucleus



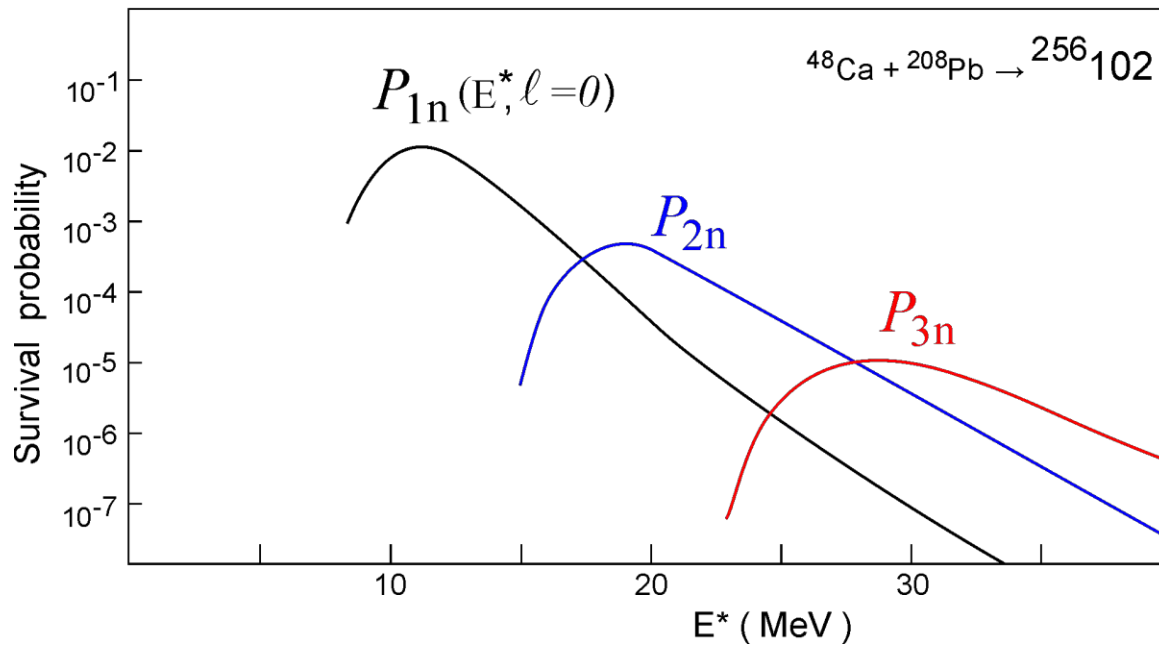
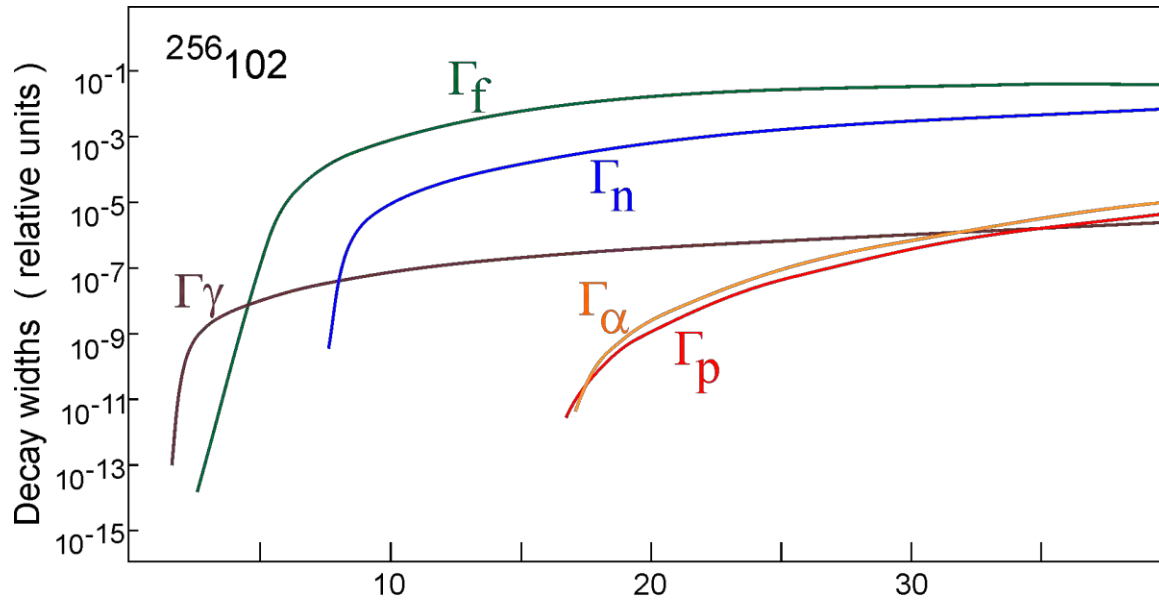
Survival probability: $\text{CN} (E_0^*, J_0) \rightarrow \text{EvR}(\text{g.s.}) + xn + N\gamma$

$$P_{xn} = \int_0^{E_0^* - E_n^{\text{sep}}(1)} \frac{\Gamma_n}{\Gamma_{\text{tot}}} (E_0^*, J_0) P_n(E_0^*, e_1) de_1 \int_0^{E_1^* - E_n^{\text{sep}}(2)} \frac{\Gamma_n}{\Gamma_{\text{tot}}} (E_1^*, J_1) P_n(E_1^*, e_2) de_2 \dots \int_0^{E_{x-1}^* - E_n^{\text{sep}}(x)} \frac{\Gamma_n}{\Gamma_{\text{tot}}} (E_{x-1}^*, J_{x-1}) P_n(E_{x-1}^*, e_x) G_{N\gamma} (E_x^*, J_x \rightarrow \text{g.s.}) de_x$$

Cross section for formation of evaporation residues:

$$\sigma_{\text{EvR}}^{xn} (E) = \frac{\pi}{k^2} \sum_{\ell} (2\ell + 1) P(E, \ell) \cdot P_{\text{CN}} (E^*, \ell) \cdot P_{xn} (E^*, \ell)$$

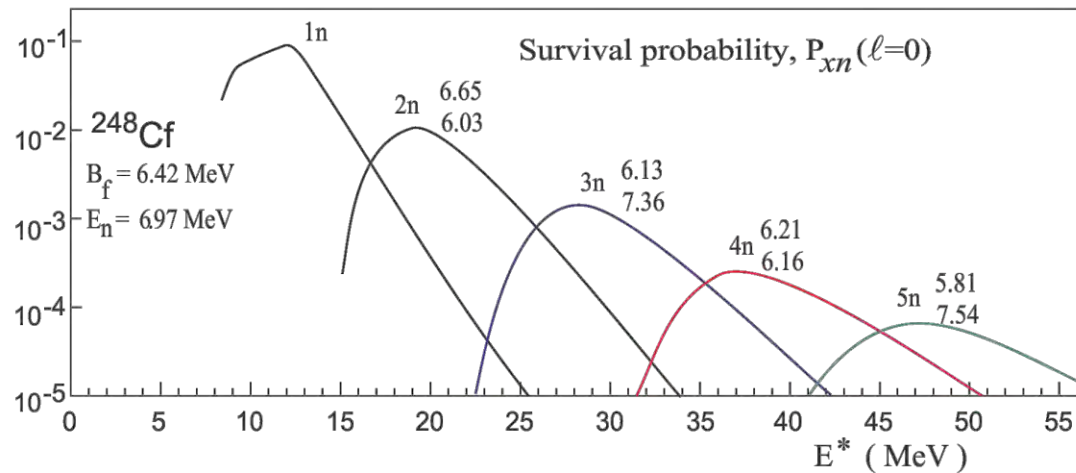
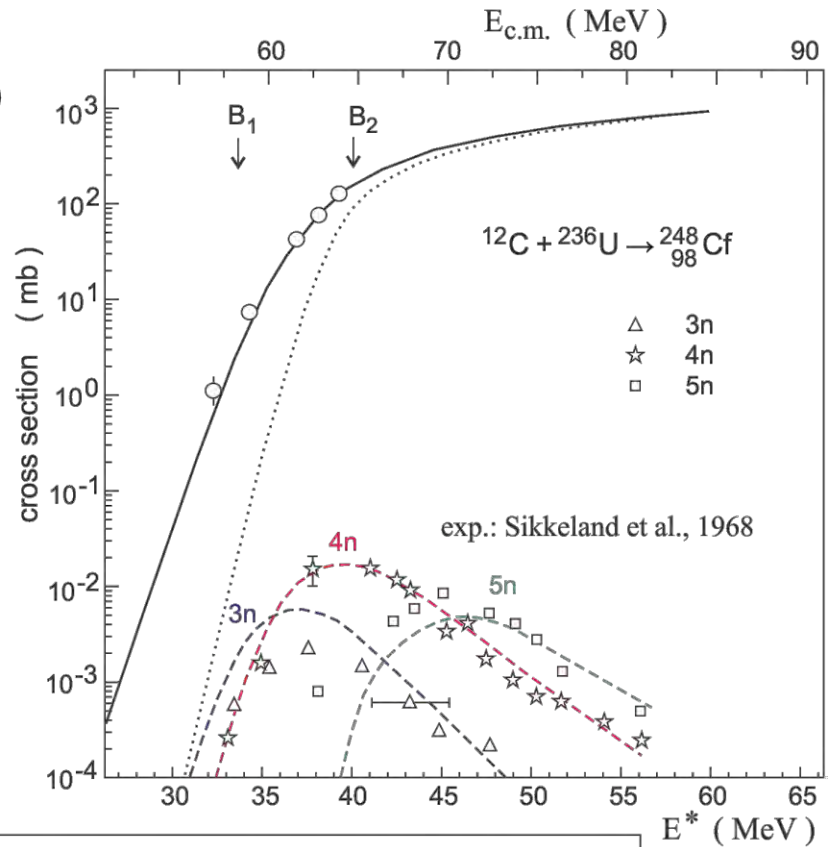
Decay widths and survival probability



Fusion reactions with light projectiles

$$\sigma_{\text{EvR}}^{xn}(E) = \frac{\pi \hbar^2}{2\mu E} \sum_{\ell} (2\ell+1) P(\ell, E) \cdot P_{\text{CN}}(\ell, E^*) \cdot P_{xn}(\ell, E^*)$$

standard CC model
= 1
statistical model



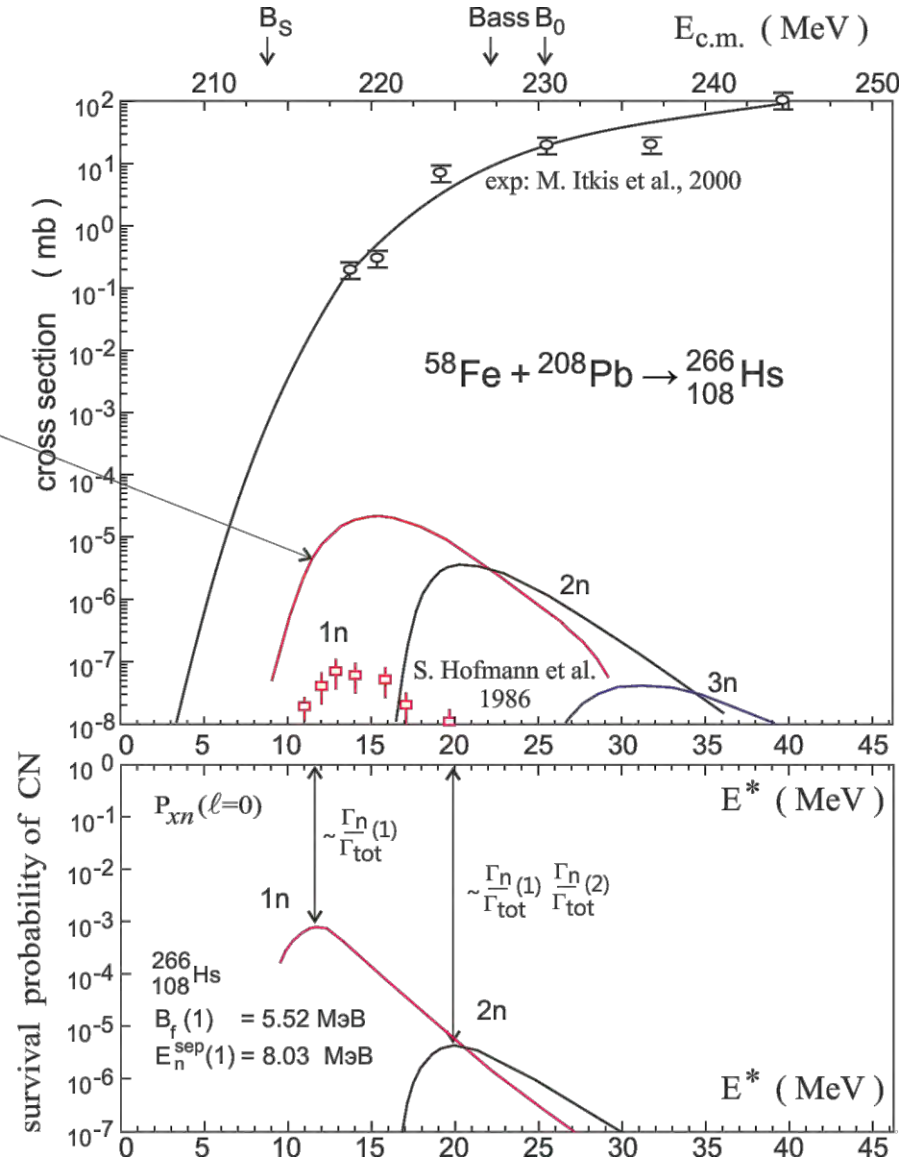
Formation of compound nucleus after contact, $P_{CN} = ?$

$$\sigma_{EvR}^{xn}(E) = \frac{\pi \hbar^2}{2\mu E} \sum_l (2l+1) P(l, E) \cdot P_{CN}(l, E^*) \cdot P_{xn}(l, E^*)$$

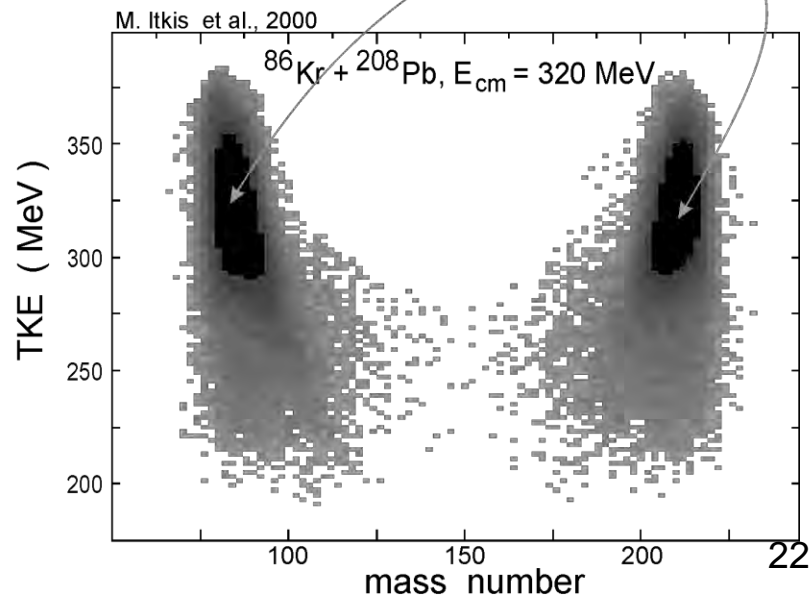
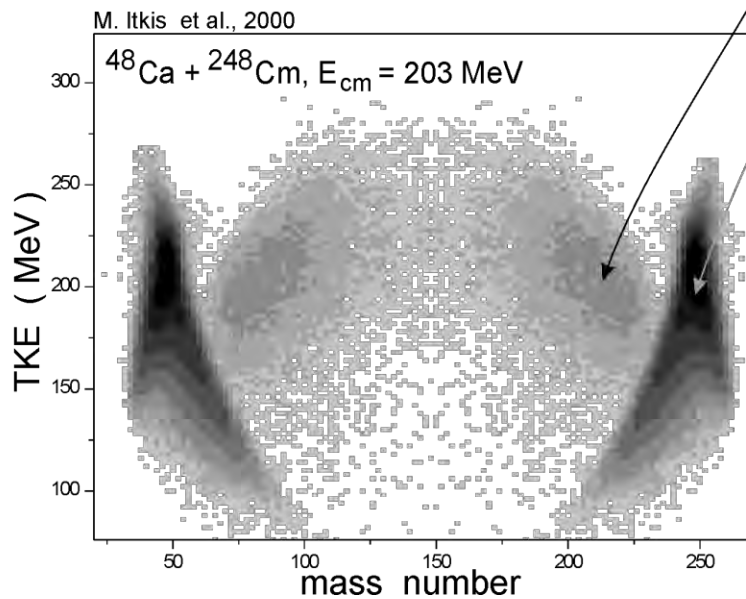
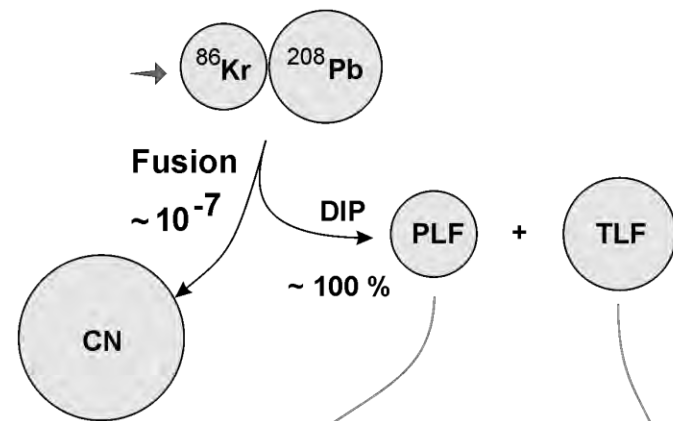
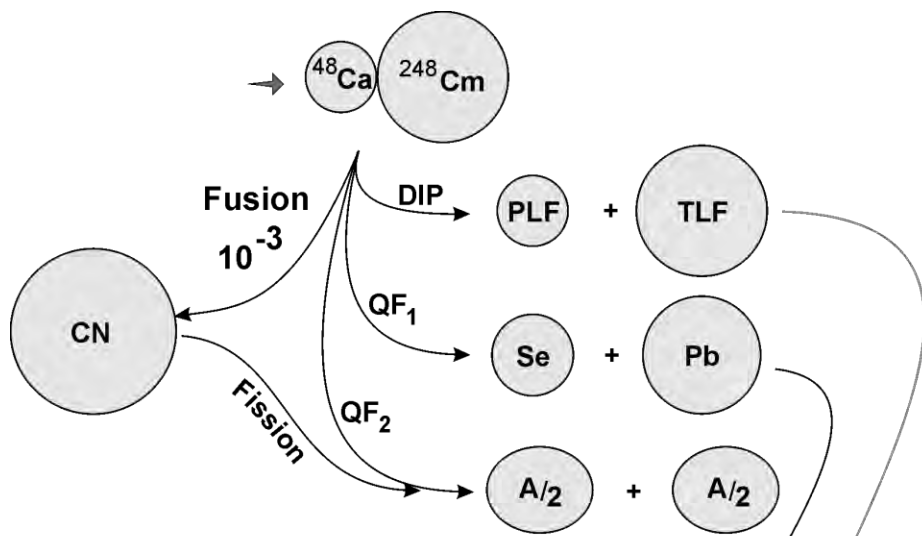
↙
↘

standard CC model
statistical model

$P_{CN}(l, E^*) = 1$



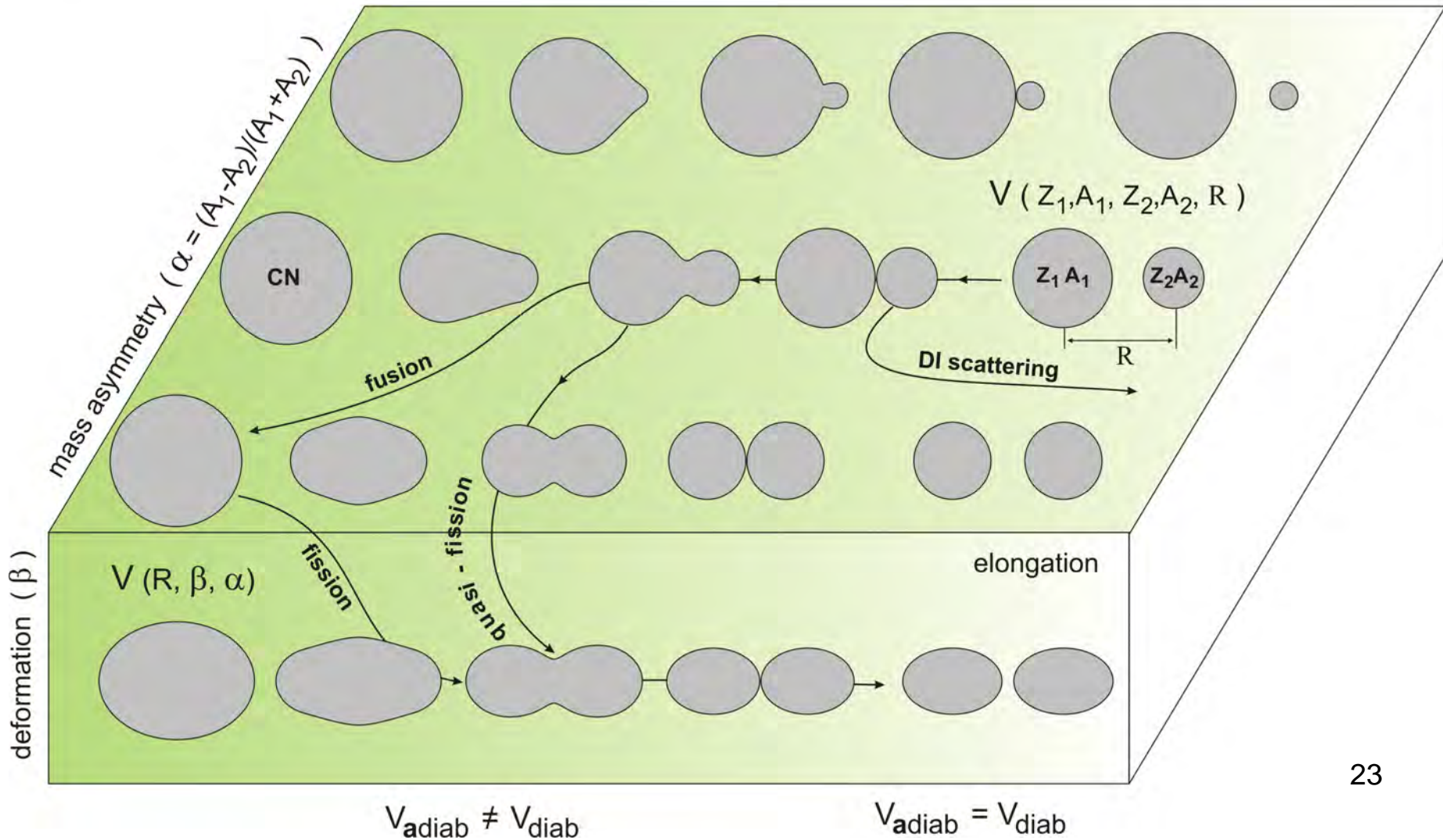
Competition of deep inelastic scattering, quasi-fission and fusion



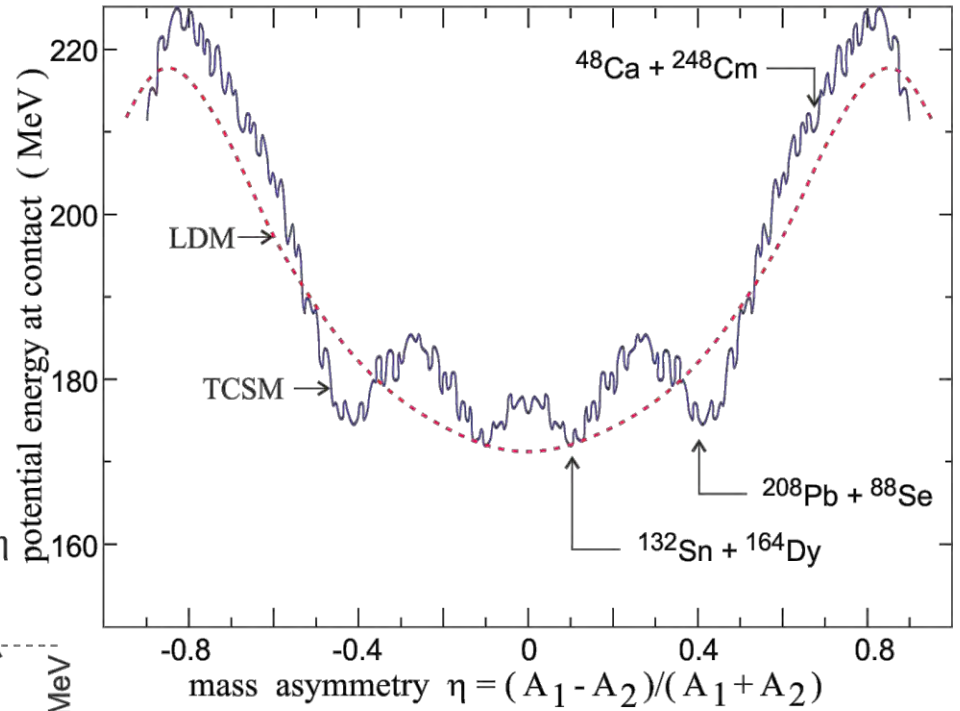
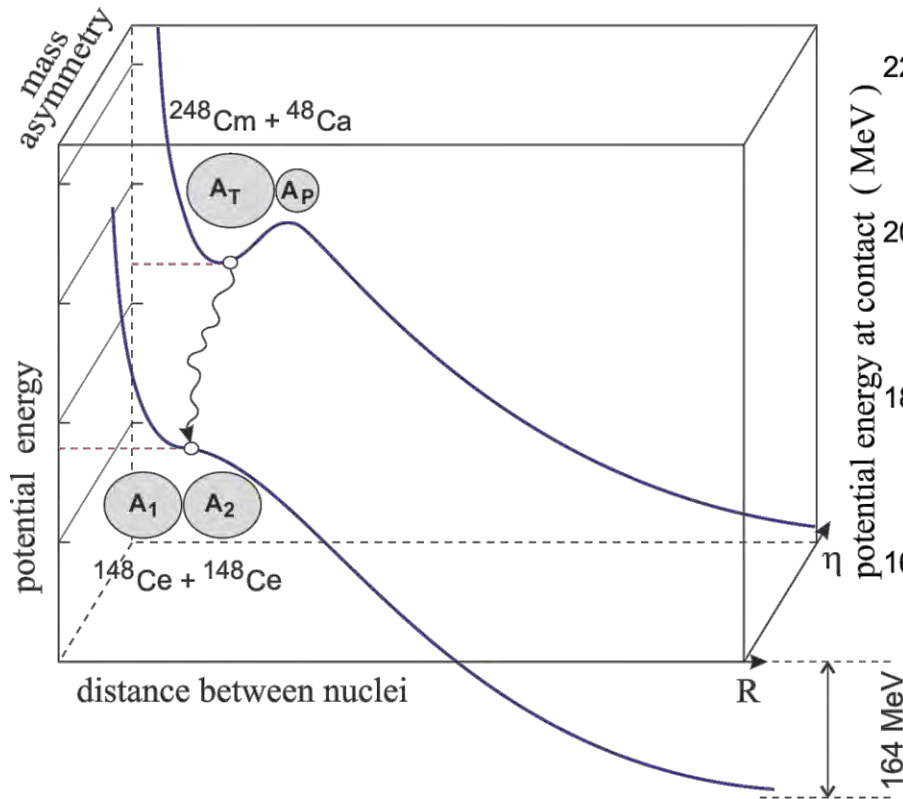
Variables ?

- ? - principal degrees of freedom: $\{q_1, q_2, \dots\}$,
- ? - potential energy surface: $V(q_1, q_2, \dots)$,
- ? - dynamic equations of motion: $dq_i/dt = \dots$

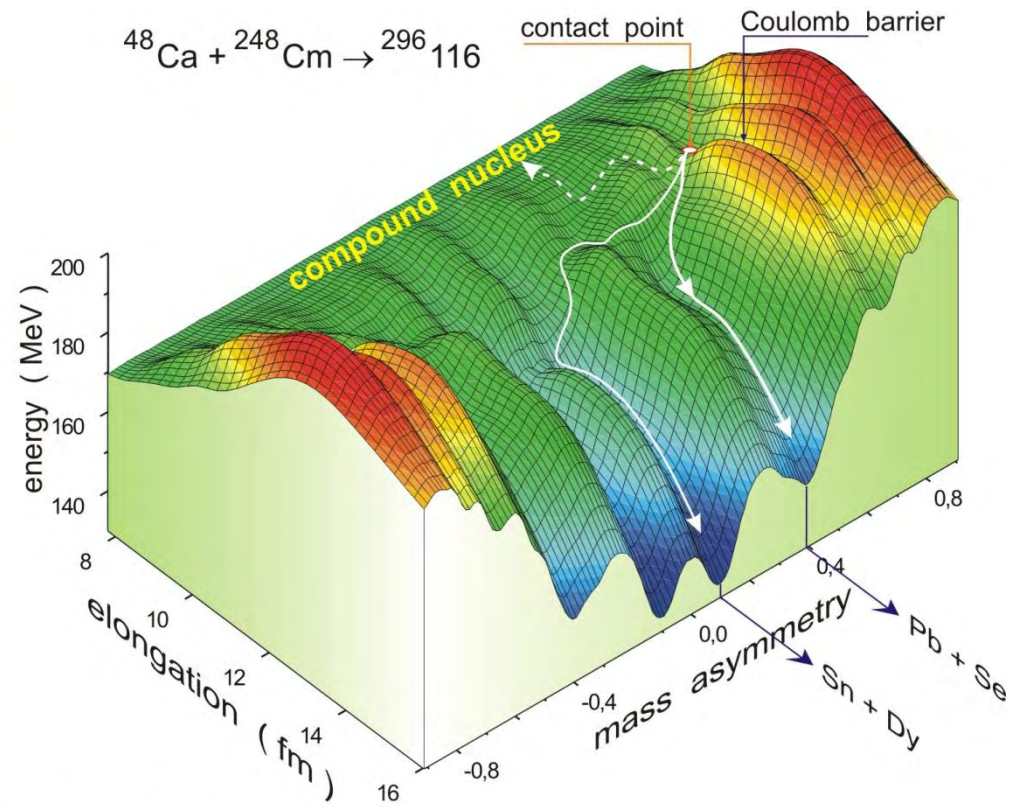
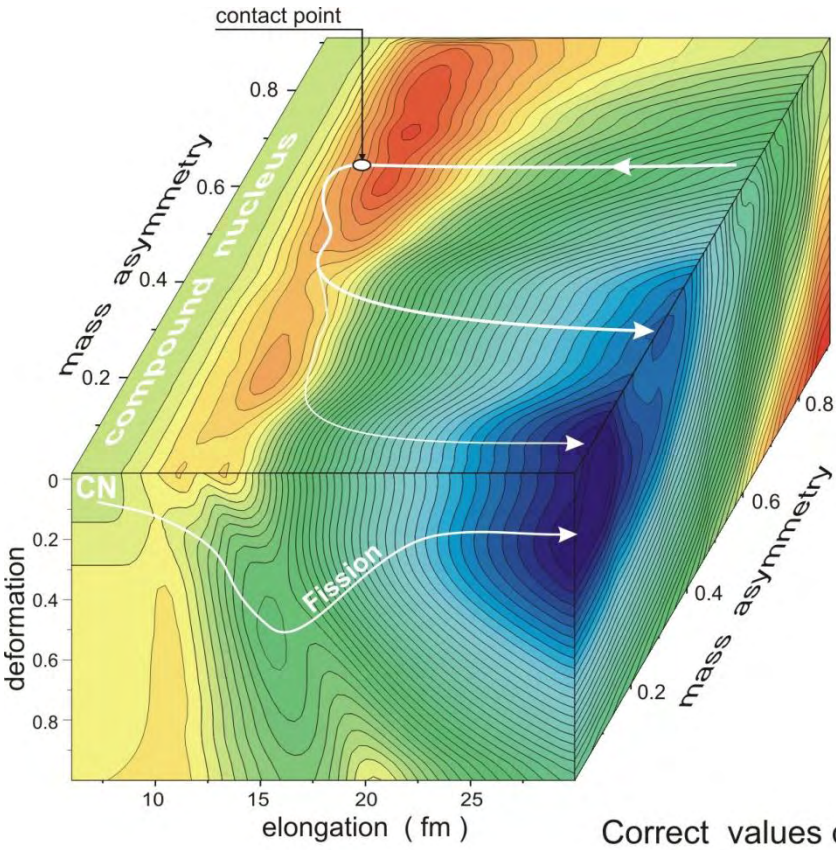
Common (unified) for all the processes:
 Deep Inelastic, Quasi-Fission, Fusion-Fission !!!



Potential Energy: Dependence on mass rearrangement



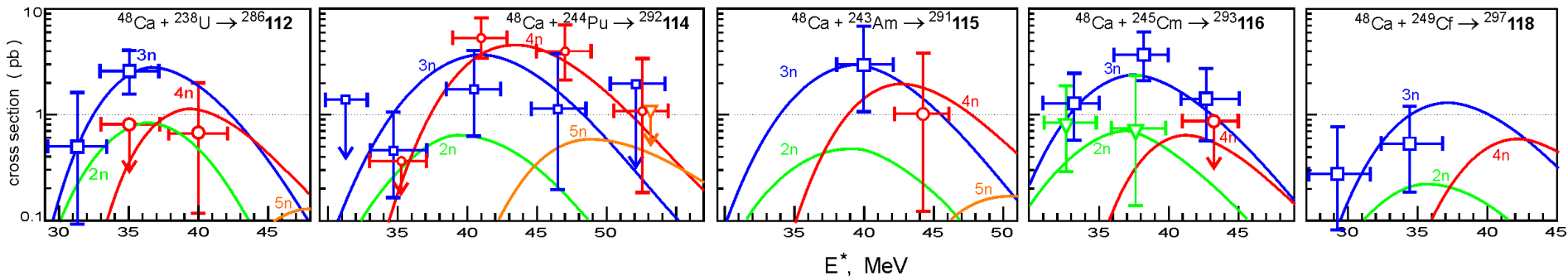
Potential Energy: Fusion, Fission and Quasi-Fission



Correct values of the Coulomb barriers and the height of the fission barrier of CN

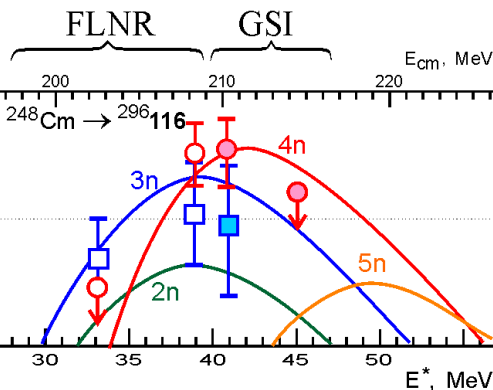
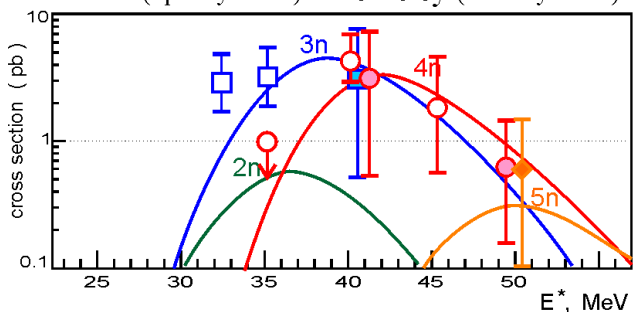
Predictive power of the theory for the hot fusion reactions

predictions of 2002

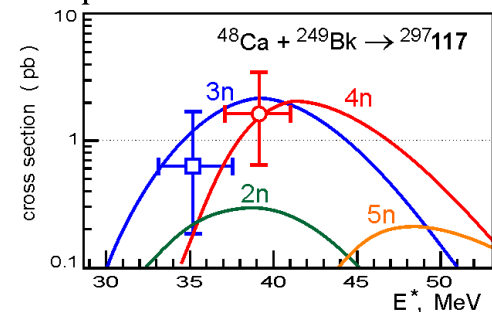


$^{48}\text{Ca} + ^{242}\text{Pu} \rightarrow ^{290}\text{114}$

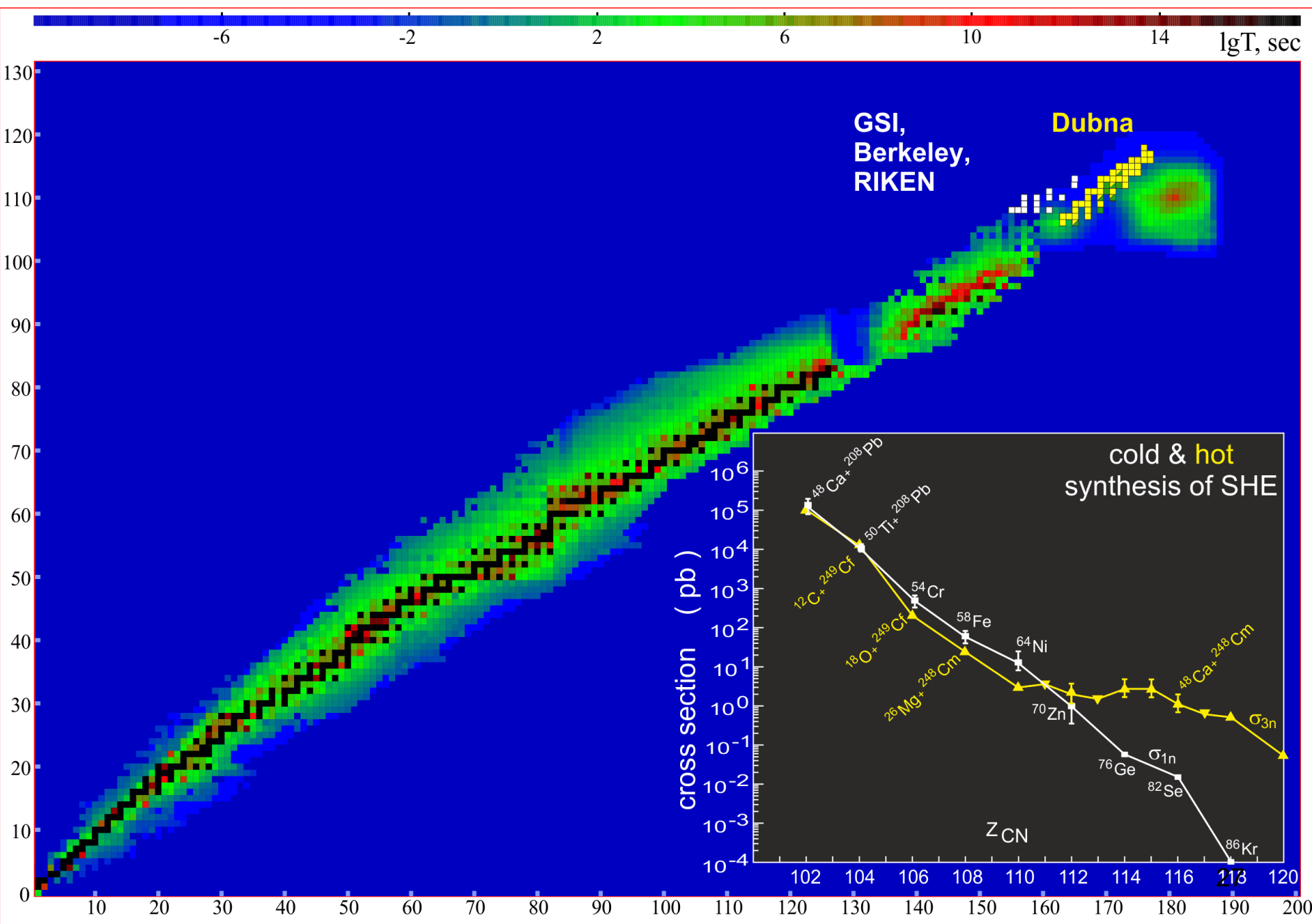
FLNR (open symbols) Berkeley (filled symbols)



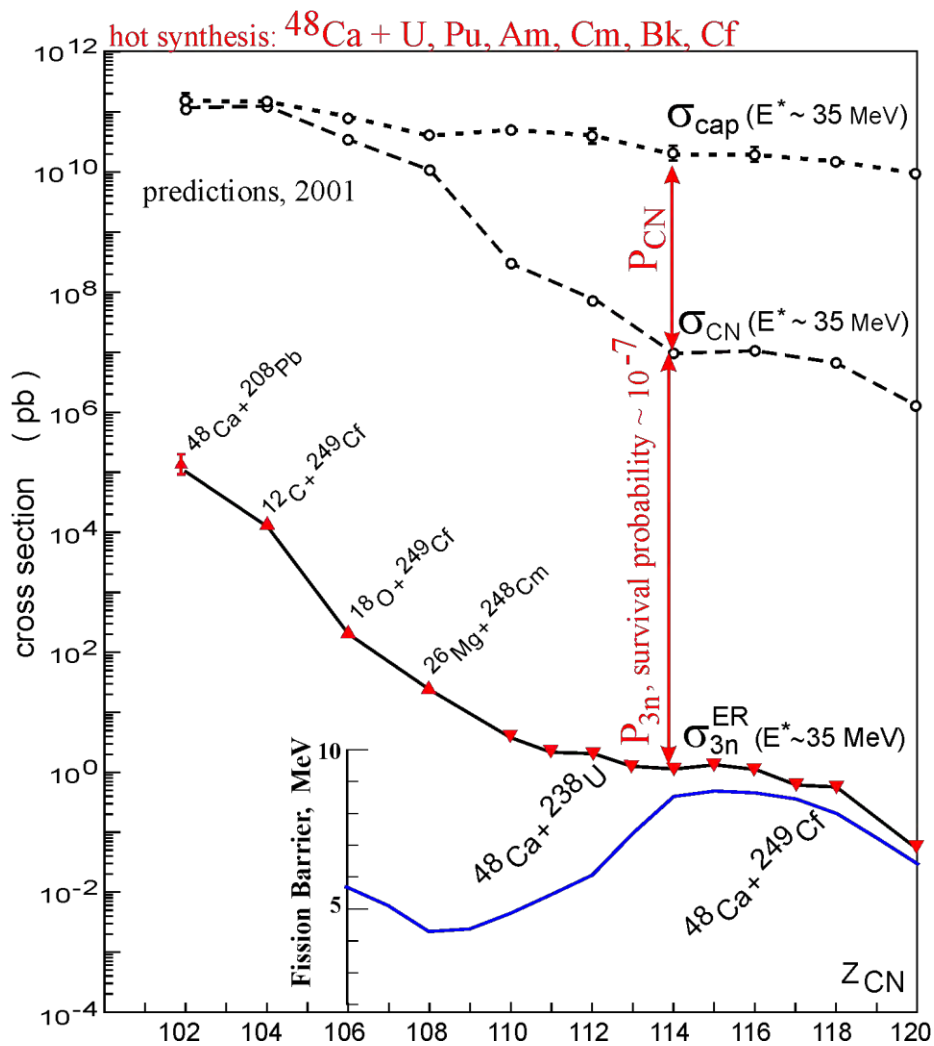
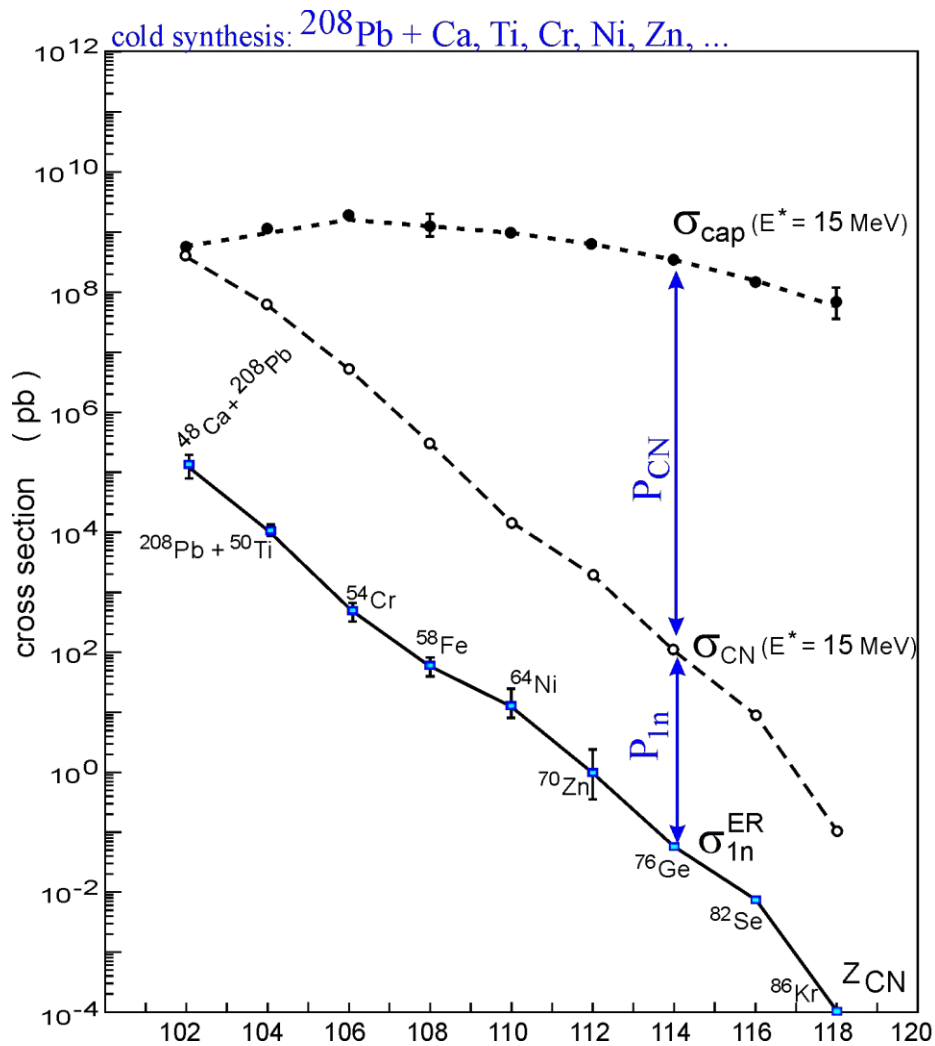
predictions of 2008



Synthesis of superheavy elements (cold and hot fusion)

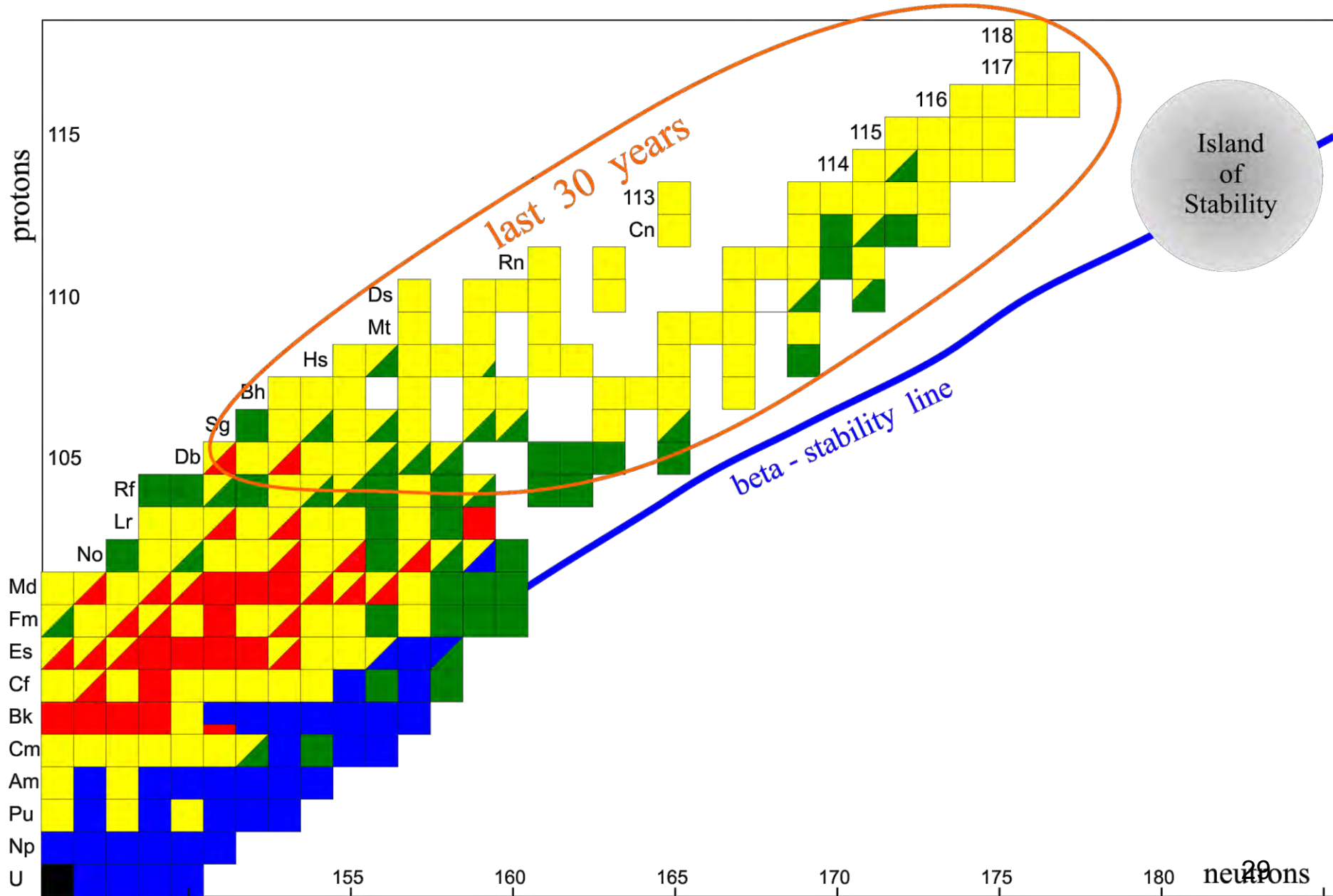


Cross sections of the “cold” and “hot” synthesis of SHE



Cross sections for formation of SHE with $Z=112-118$ have been predicted to be nearly constant owing to increasing values of the fission barriers of formed CN

Great progress in synthesis of superheavy nuclei. What's the next?



New elements 119 and 120 are coming !

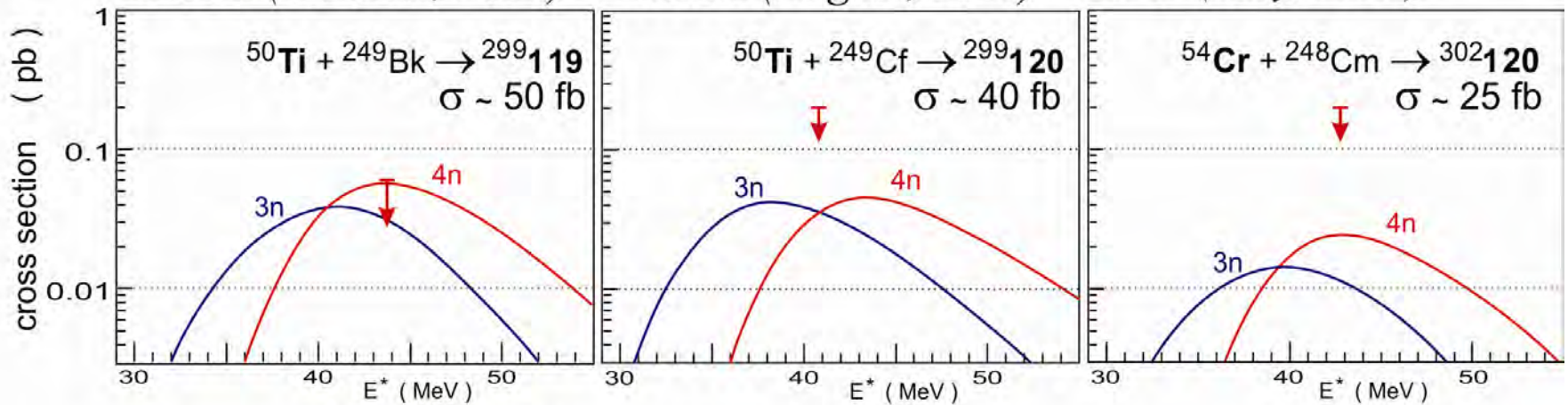
Ti beam:

TASCA (October, 2012)

TASCA (August, 2011)

Cr beam:

SHIP (May, 2011)

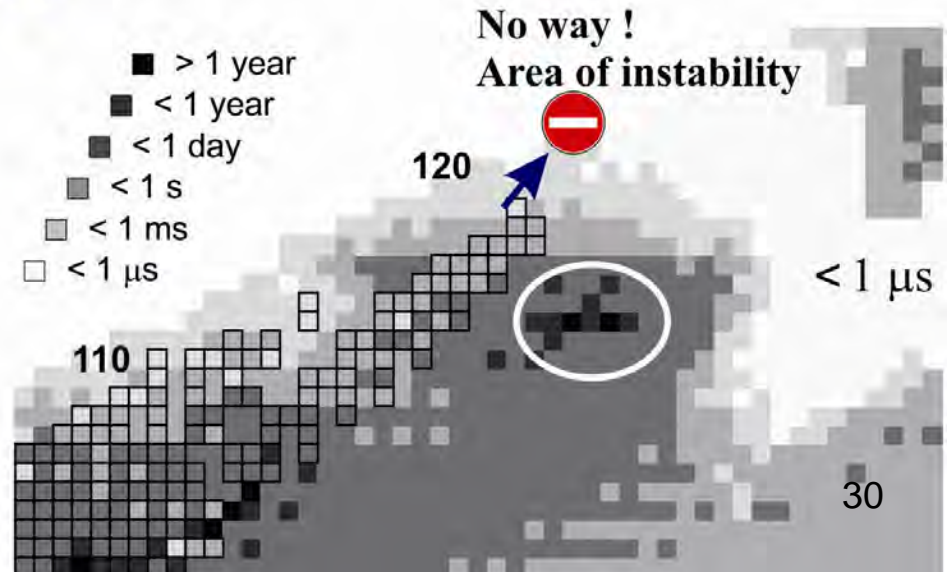


our predictions (PRC 2008):

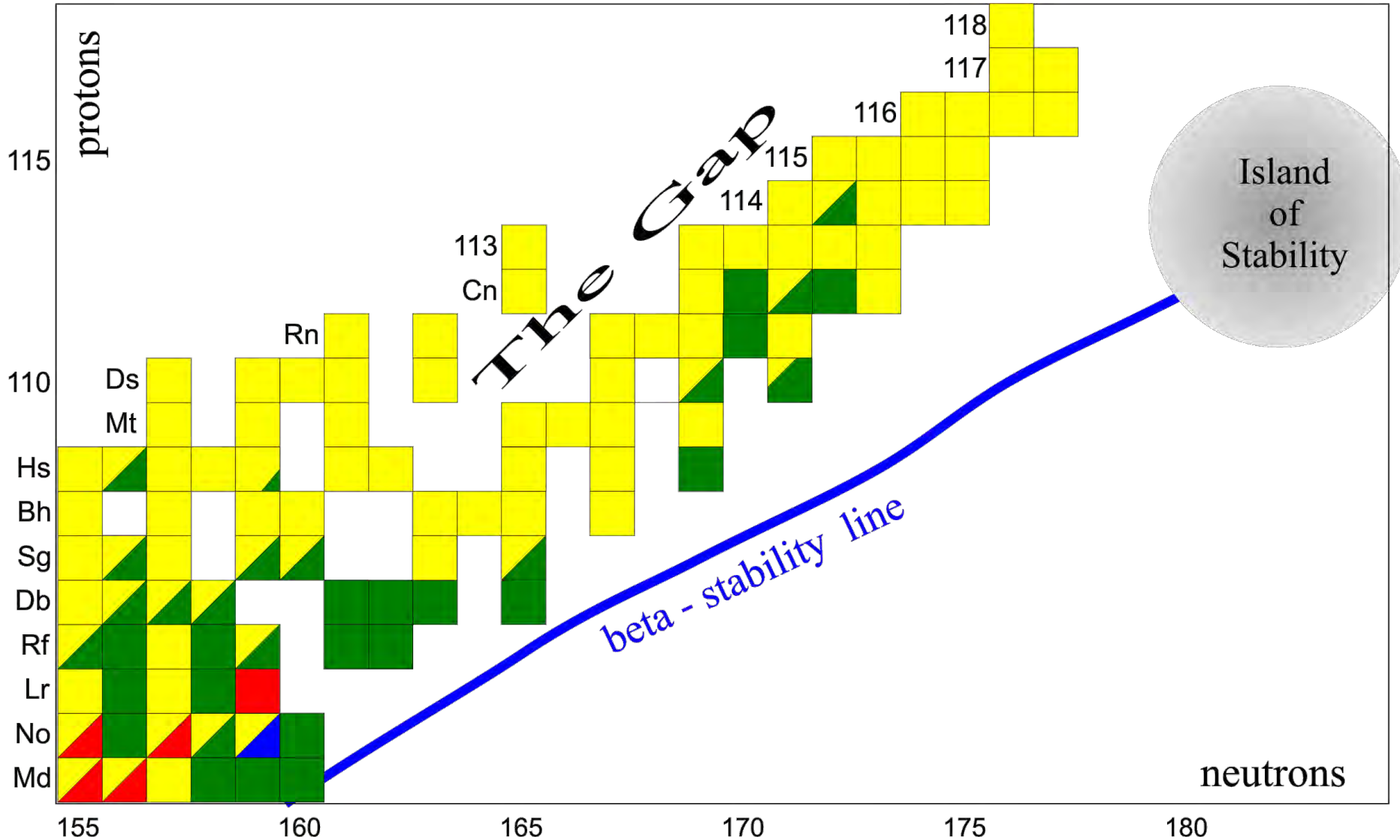
factor $\frac{1}{20}$ as compared to ^{48}Ca

Approaching the area of instability:

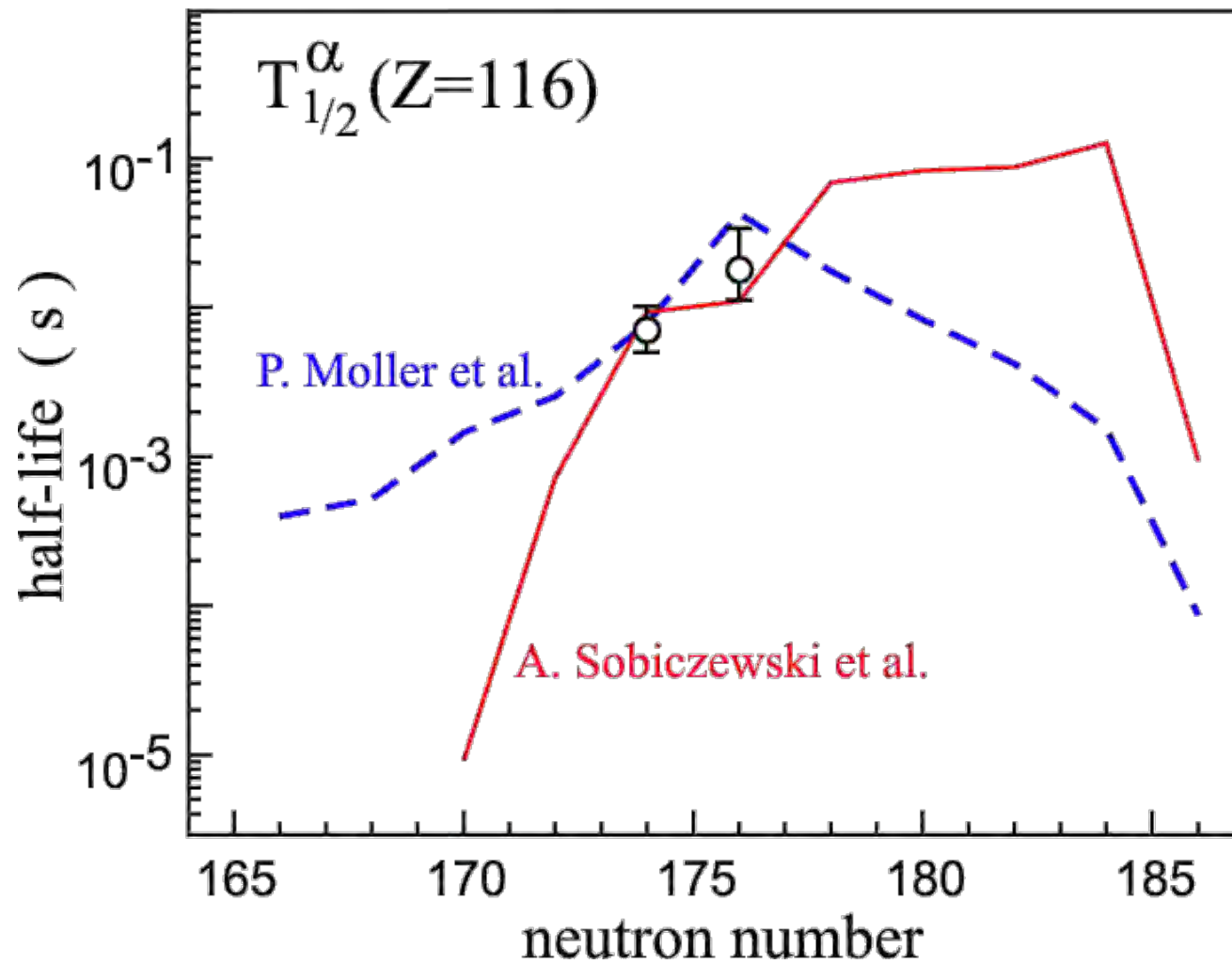
Probably, these elements are the last ones which will be synthesized in the nearest future



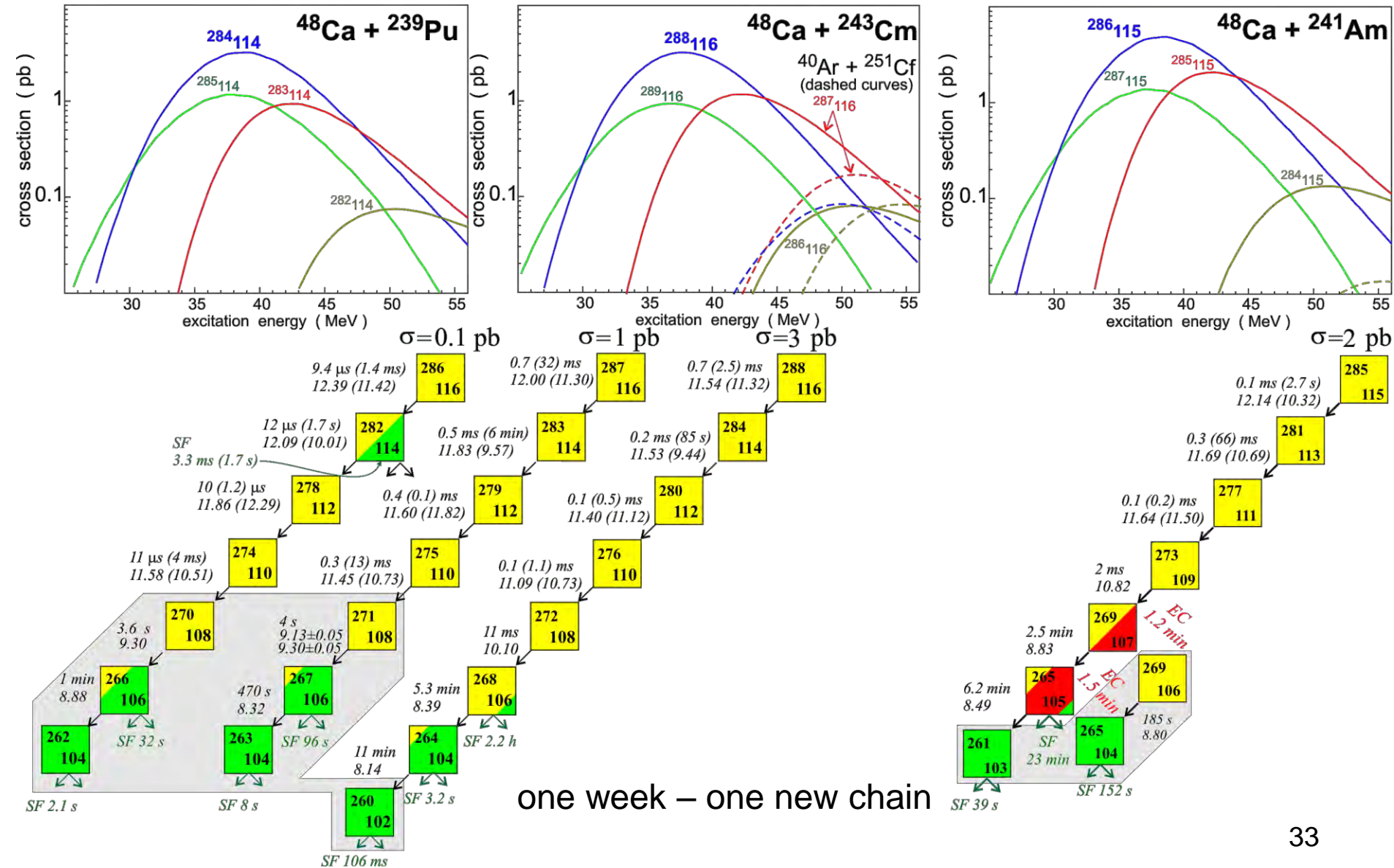
The gap in SH mass area must be filled somehow



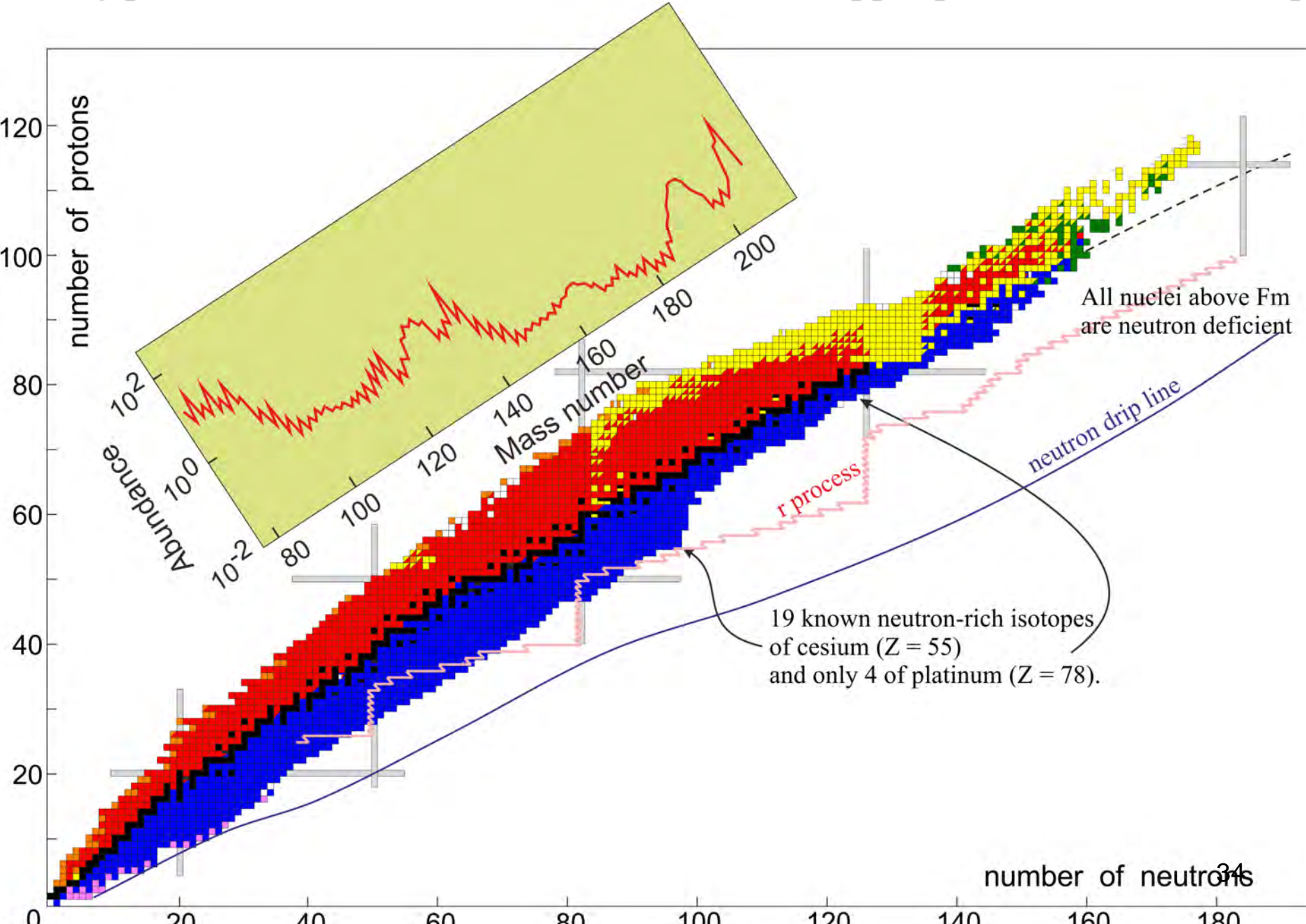
Our ability of predictions in superheavy mass area



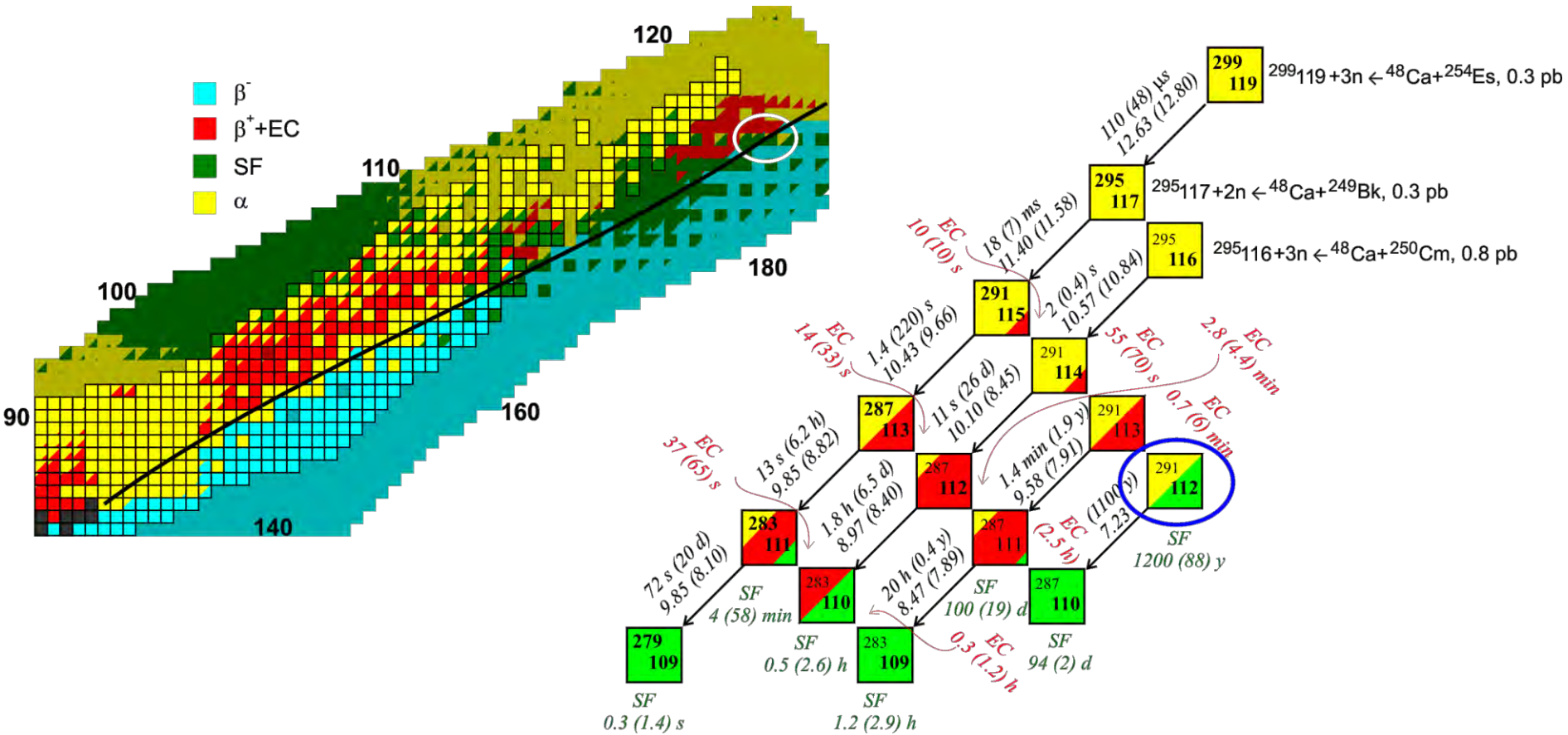
Predicted cross sections are high enough to perform experiments at available facilities just now



Mostly proton-rich nuclei were studied so far in the upper part of the nuclear map



Narrow pathway to the island of stability just by fusion reactions !



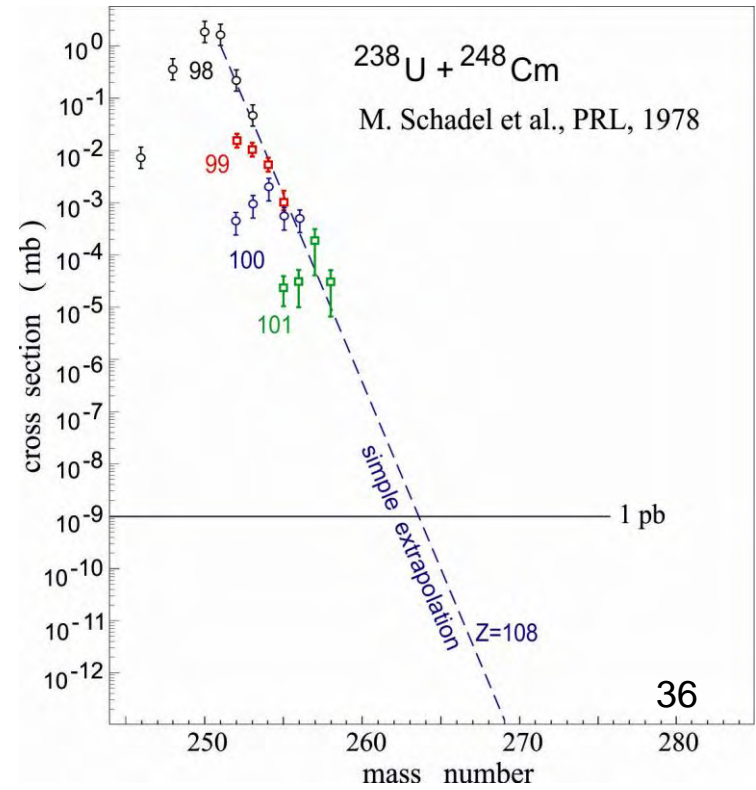
Synthesis of SH nuclei in transfer reactions

- [1] E. K. Hulet *et al.*, Phys. Rev. Lett. **39**, 385 (1977).
- [2] M. Schaedel *et al.*, Phys. Rev. Lett. **41**, 469 (1978).
- [3] H. Essel, K. Hartel, W. Henning, P. Kienle, H. J. Koerner, K. E. Rehm, P. Sperr, W. Wagner, and H. Spieler, Z. Phys. A **289**, 265 (1979).
- [4] H. Freiesleben, K. D. Hildenbrand, F. Puhlhofer, W. F. W. Schneider, R. Bock, D. V. Harrach, and H. J. Specht, Z. Phys. A **292**, 171 (1979).
- [5] H. Gaeggeler *et al.*, Phys. Rev. Lett. **45**, 1824 (1980).
- [6] M. Schaedel *et al.*, Phys. Rev. Lett. **48**, 852 (1982).
- [7] K. J. Moody, D. Lee, R. B. Welch, K. E. Gregorich, G. T. Seaborg, R. W. Lougheed, and E. K. Hulet, Phys. Rev. C **33**, 1315 (1986).
- [8] R. B. Welch, K. J. Moody, K. E. Gregorich, D. Lee, and G. T. Seaborg, Phys. Rev. C **35**, 204 (1987).

...

... a long history.

Isotopes of Fm and Md were synthesized 30 years ago.



Theoretical models of transfer reactions

Multi-nucleon transfers in damped collisions

Master equation

L.G. Moretto and J.S. Sventek, Phys. Lett. B **58**, 26 (1975)

Fokker-Plank equation

W. Norenberg, Phys. Lett. B **52**, 289 (1974)

Langevin equations

P. Frobrich and S.Y. Xu, Nucl. Phys. **A477**, 143 (1988)

Semi-classical approaches

E. Vigezzi and A. Winther, Ann. Phys. (N.Y.) **192**, 432 (1989).

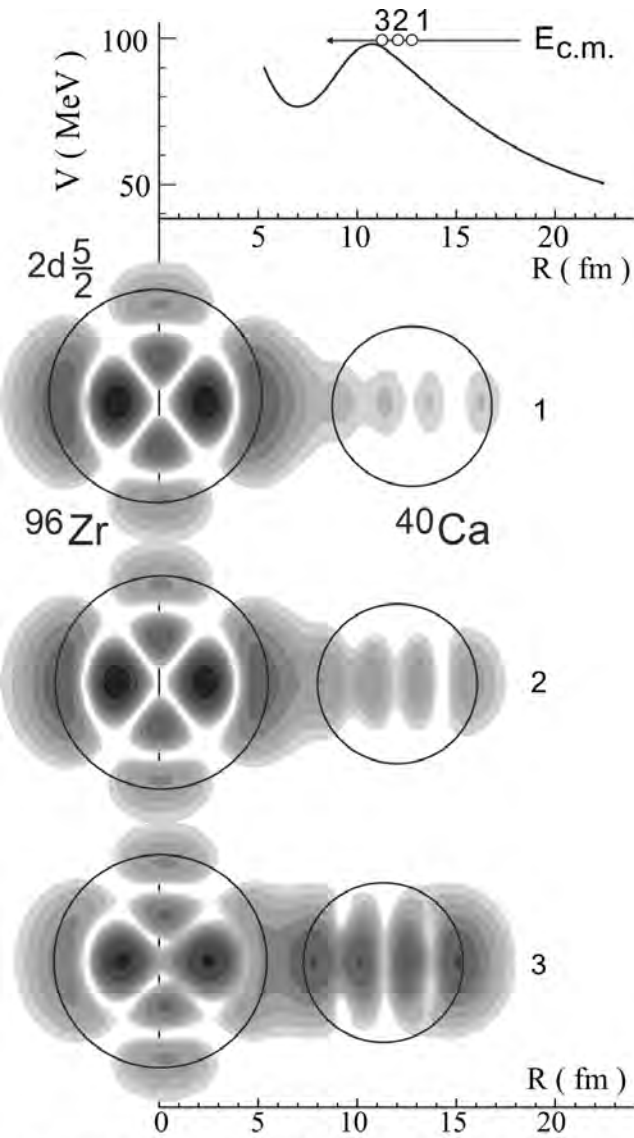
V.I. Zagrebaev, Ann. Phys. (N.Y.) **197**, 33 (1990).

Few-nucleon transfers (GRAZING)

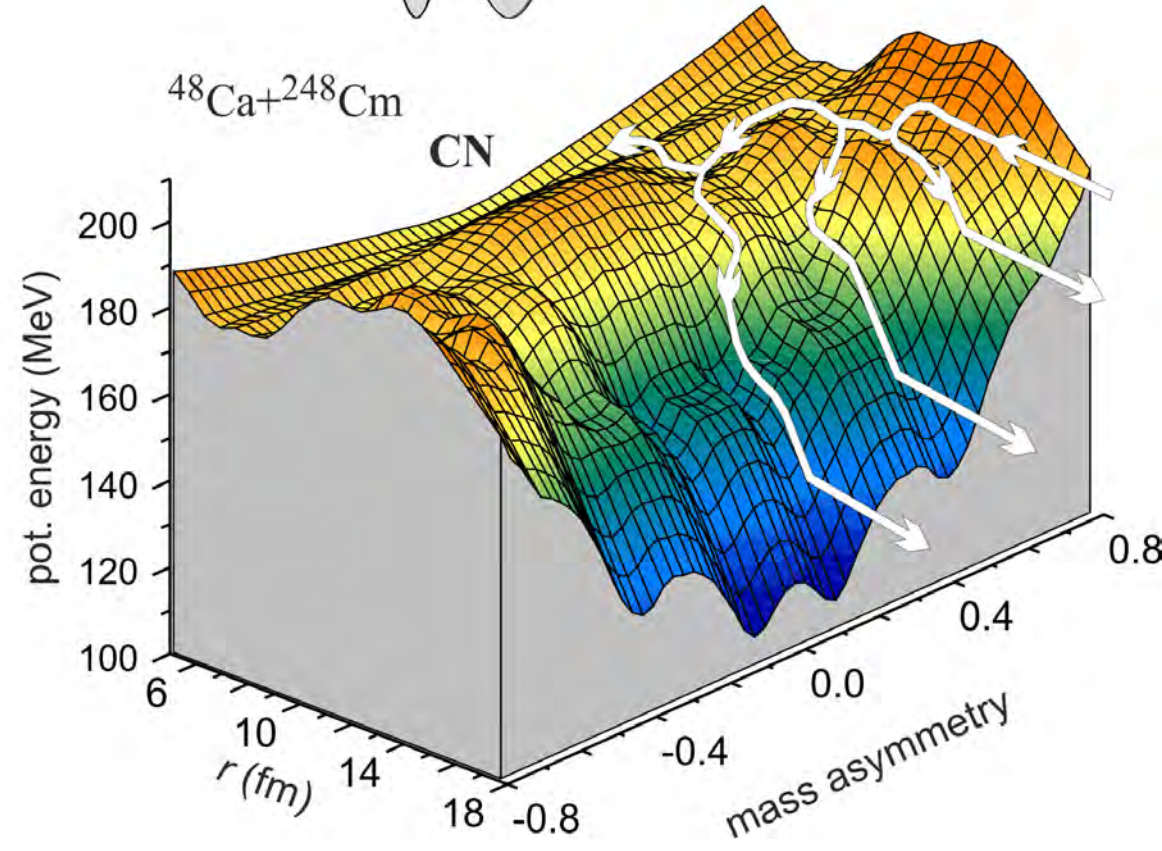
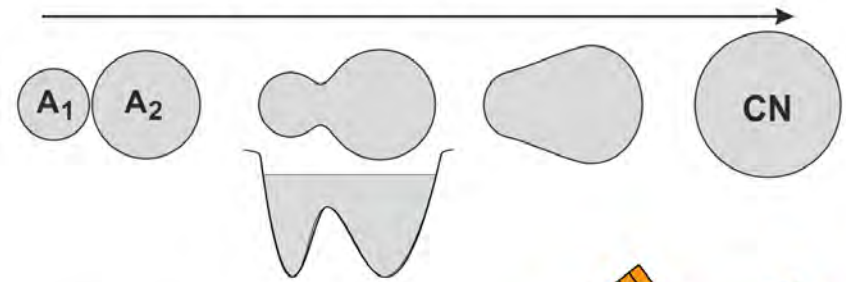
A. Winther, Nucl. Phys. **A594**, 203 (1995)

<http://personalpages.to.infn.it/nanni/grazing>

Adiabatic dynamics of low-energy heavy ion collisions and nucleon transfers



- overlapped mean fields
- two-center shell model
- adiabatic potential energy



time-dependent Schrödinger equation for single particle wave functions
 (Zagrebaev, Samarin, Greiner, 2007);

System of coupled Langevin type Equations of Motion

$$\frac{dR}{dt} = \frac{p_R}{\mu_R}$$

Variables: $\{R, \theta, \varphi_1, \varphi_2, \beta_1, \beta_2, \eta_Z, \eta_N\}$

$$\frac{d\vartheta}{dt} = \frac{\ell}{\mu_R R^2}$$

$$\frac{d\varphi_1}{dt} = \frac{L_1}{\mathfrak{I}_1}, \quad \frac{d\varphi_2}{dt} = \frac{L_2}{\mathfrak{I}_2}$$

$$\frac{d\beta_1}{dt} = \frac{p_{\beta_1}}{\mu_{\beta_1}}$$

$$\frac{d\beta_2}{dt} = \frac{p_{\beta_2}}{\mu_{\beta_2}}$$

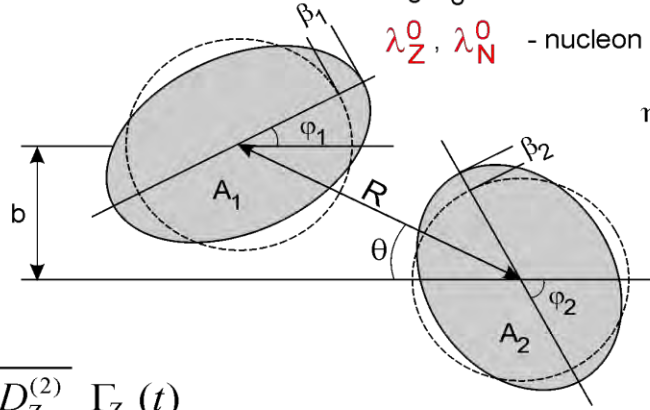
$$\frac{d\eta_Z}{dt} = \frac{2}{Z_{CN}} D_Z^{(1)} + \frac{2}{Z_{CN}} \sqrt{D_Z^{(2)}} \Gamma_Z(t)$$

$$\frac{d\eta_N}{dt} = \frac{2}{N_{CN}} D_N^{(1)} + \frac{2}{N_{CN}} \sqrt{D_N^{(2)}} \Gamma_N(t)$$

Most uncertain parameters:

μ_0, γ_0 - nuclear viscosity and friction,

λ_Z^0, λ_N^0 - nucleon transfer rate



$$\eta = \frac{A_1 - A_2}{A_1 + A_2}$$

$$\eta_Z = \frac{Z_1 - Z_2}{Z_1 + Z_2}$$

$$\eta_N = \frac{N_1 - N_2}{N_1 + N_2}$$

$$\lambda_Z^0 = \lambda_N^0 = \frac{\lambda^0}{2}$$

$$\frac{dp_R}{dt} = -\frac{\partial V}{\partial R} + \frac{\ell^2}{\mu_R R^3} + \left(\frac{\ell^2}{2\mu_R^2 R^2} + \frac{p_R^2}{2\mu_R^2} \right) \frac{\partial \mu_R}{\partial R} + \frac{p_{\beta_1}^2}{2\mu_{\beta_1}^2} \frac{\partial \mu_{\beta_1}}{\partial R} + \frac{p_{\beta_2}^2}{2\mu_{\beta_2}^2} \frac{\partial \mu_{\beta_2}}{\partial R} - \gamma_R \frac{p_R}{\mu_R} + \sqrt{\gamma_R} T \Gamma_R(t)$$

$$\frac{d\ell}{dt} = -\frac{\partial V}{\partial \vartheta} - \gamma_{\text{tang}} \left(\frac{\ell}{\mu_R R} - \frac{L_1}{\mathfrak{I}_1} a_1 - \frac{L_2}{\mathfrak{I}_2} a_2 \right) R + \sqrt{\gamma_{\text{tang}}} T \Gamma_{\text{tang}}(t)$$

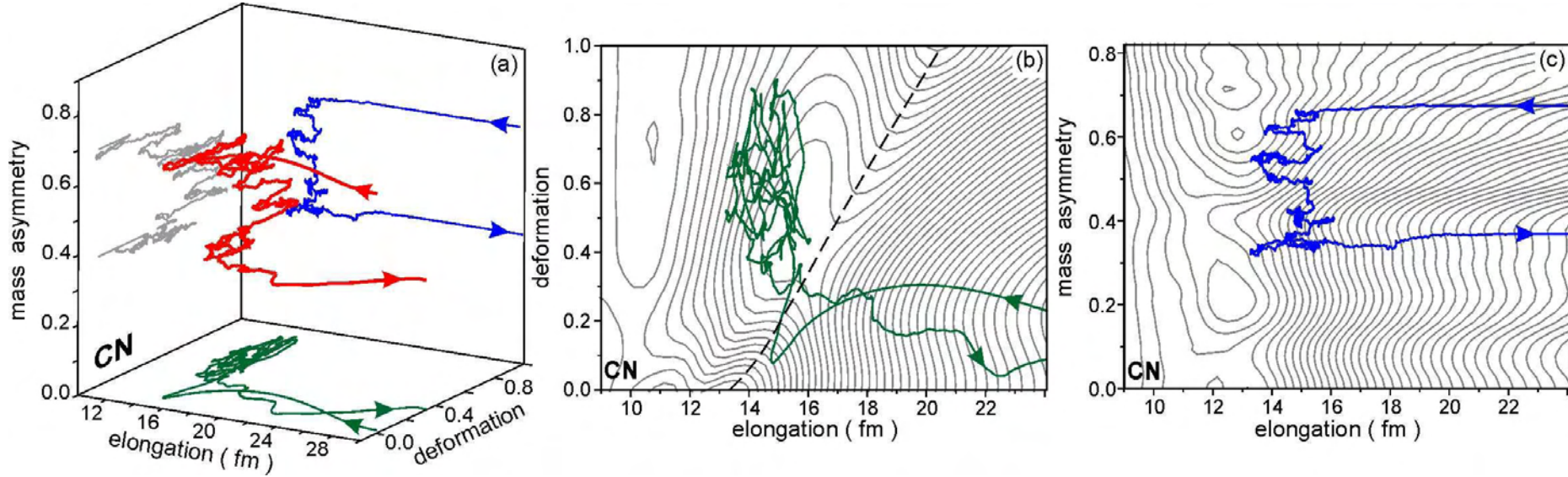
$$\frac{dL_1}{dt} = -\frac{\partial V}{\partial \varphi_1} + \gamma_{\text{tang}} \left(\frac{\ell}{\mu_R R} - \frac{L_1}{\mathfrak{I}_1} a_1 - \frac{L_2}{\mathfrak{I}_2} a_2 \right) a_1 - \frac{a_1}{R} \sqrt{\gamma_{\text{tang}}} T \Gamma_{\text{tang}}(t)$$

$$\frac{dL_2}{dt} = -\frac{\partial V}{\partial \varphi_2} + \gamma_{\text{tan}} \left(\frac{\ell}{\mu_R R} - \frac{L_1}{\mathfrak{I}_1} a_1 - \frac{L_2}{\mathfrak{I}_2} a_2 \right) a_2 - \frac{a_2}{R} \sqrt{\gamma_{\text{tang}}} T \Gamma_{\text{tang}}(t)$$

$$\frac{dp_{\beta_1}}{dt} = -\frac{\partial V}{\partial \beta_1} + \frac{p_{\beta_1}^2}{2\mu_{\beta_1}^2} \frac{\partial \mu_{\beta_1}}{\partial \beta_1} + \frac{p_{\beta_2}^2}{2\mu_{\beta_2}^2} \frac{\partial \mu_{\beta_2}}{\partial \beta_1} + \left(\frac{\ell^2}{2\mu_R^2 R^2} + \frac{p_R^2}{2\mu_R^2} \right) \frac{\partial \mu_R}{\partial \beta_1} - \gamma_{\beta} \frac{p_{\beta_1}}{\mu_{\beta_1}} + \sqrt{\gamma_{\beta_1}} T \Gamma_{\beta_1}(t)$$

$$\frac{dp_{\beta_2}}{dt} = -\frac{\partial V}{\partial \beta_2} + \frac{p_{\beta_1}^2}{2\mu_{\beta_1}^2} \frac{\partial \mu_{\beta_1}}{\partial \beta_2} + \frac{p_{\beta_2}^2}{2\mu_{\beta_2}^2} \frac{\partial \mu_{\beta_2}}{\partial \beta_2} + \left(\frac{\ell^2}{2\mu_R^2 R^2} + \frac{p_R^2}{2\mu_R^2} \right) \frac{\partial \mu_R}{\partial \beta_2} - \gamma_{\beta} \frac{p_{\beta_2}}{\mu_{\beta_2}} + \sqrt{\gamma_{\beta_2}} T \Gamma_{\beta_2}(t)$$

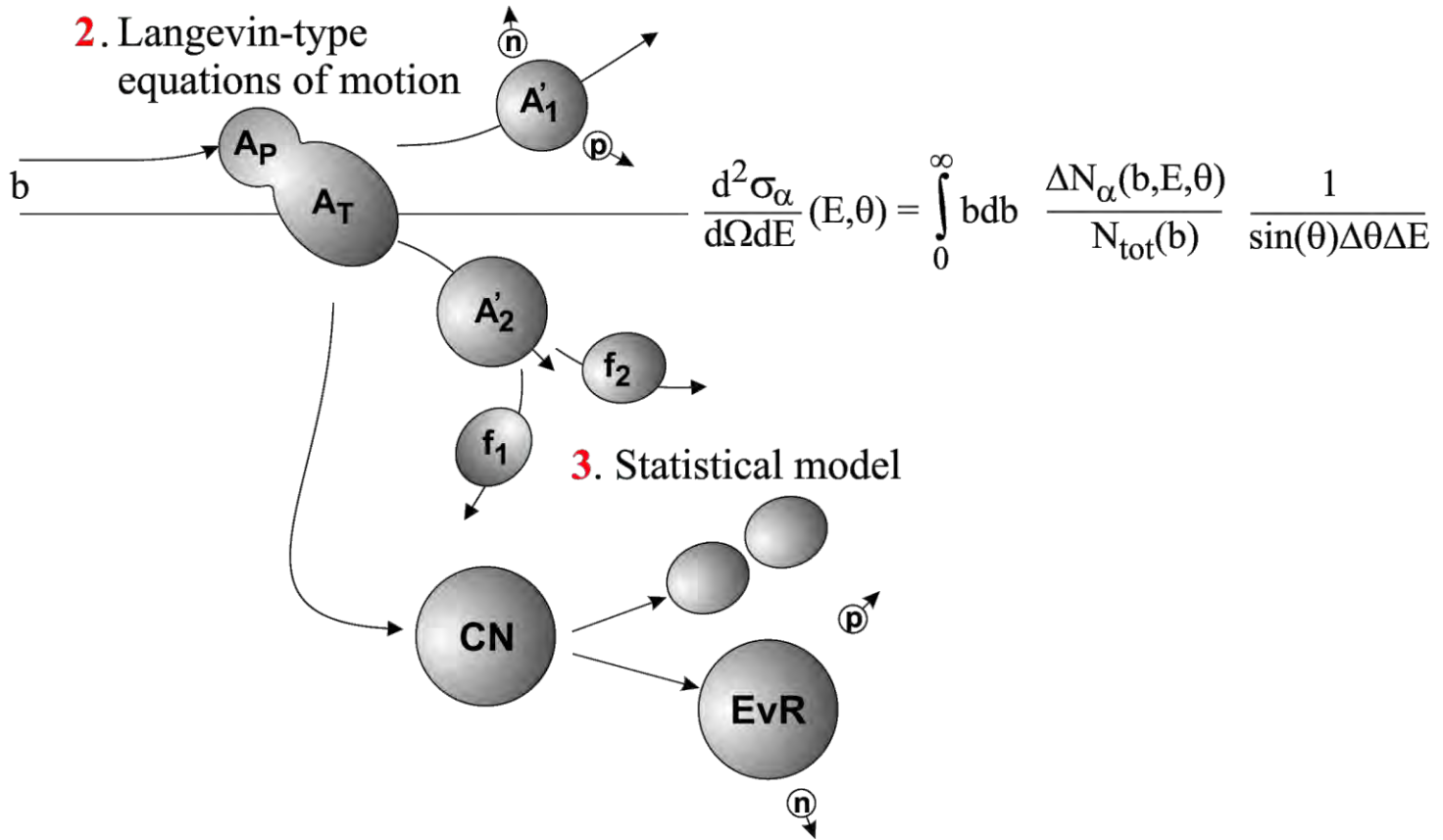
Typical trajectory in the “distance-deformation-mass asymmetry” space ($48\text{Ca} + 248\text{Cm}$, $E=210$ MeV)



Simulation of experiment. Cross sections

1. Time-dependent driving potential $V(r, \xi; t)$:
Folding \rightarrow Adiabatic Two-Center Shell Model

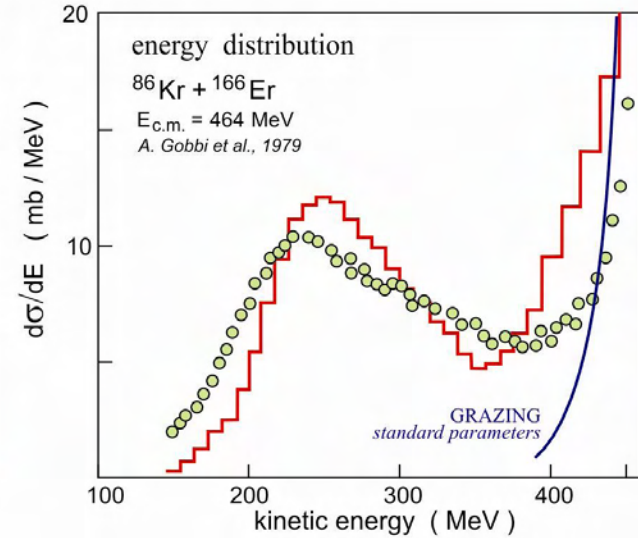
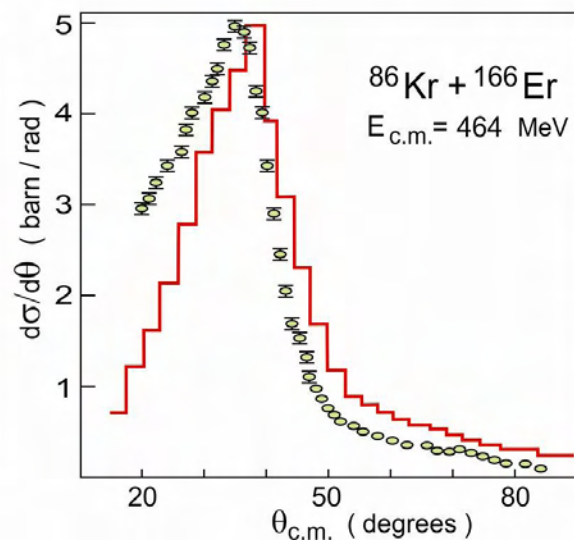
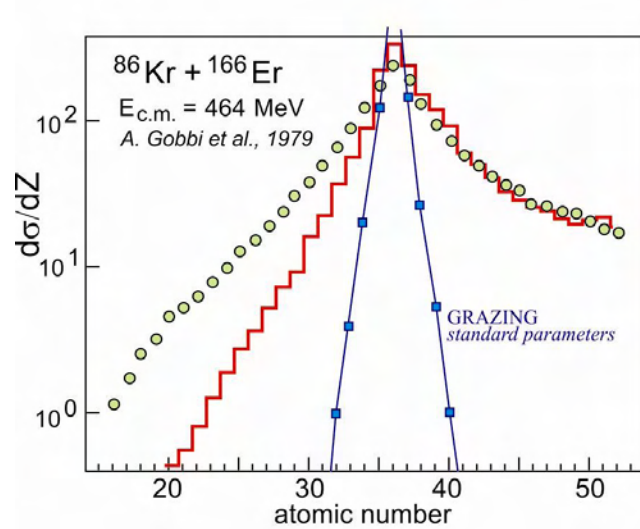
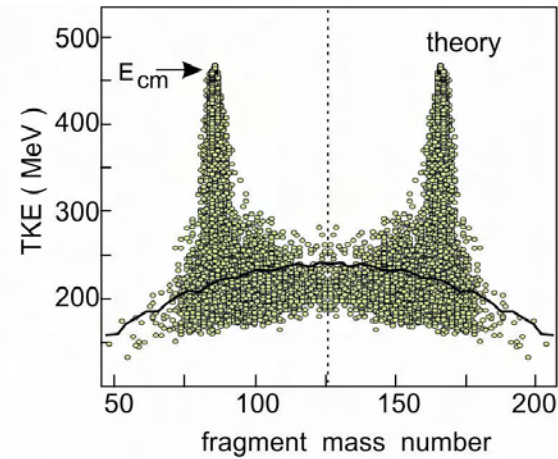
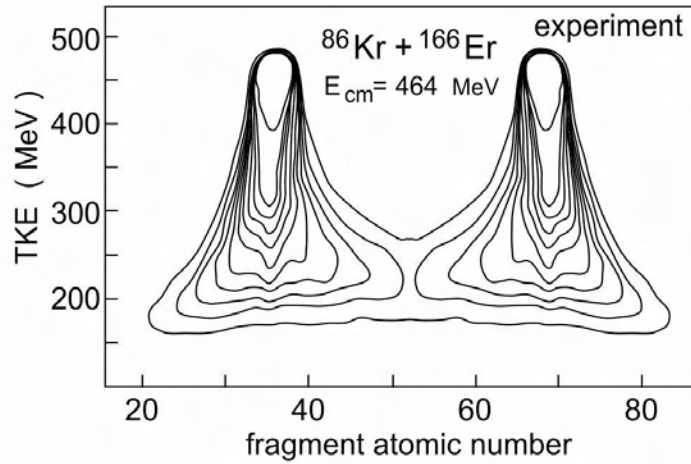
2. Langevin-type equations of motion



3. Statistical model

Dynamics: 10^6 tested events (trajectories),
 Statistical model: 10^{-6} ($3n$), 10^{-7} ($4n$) survival probability
 cross sections up to **0.1 pb** can be calculated

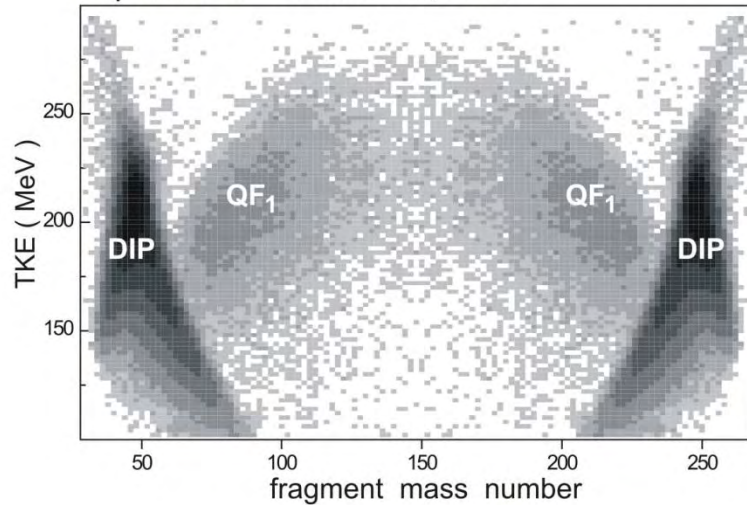
Quite satisfactory agreement with experiments on DI scattering



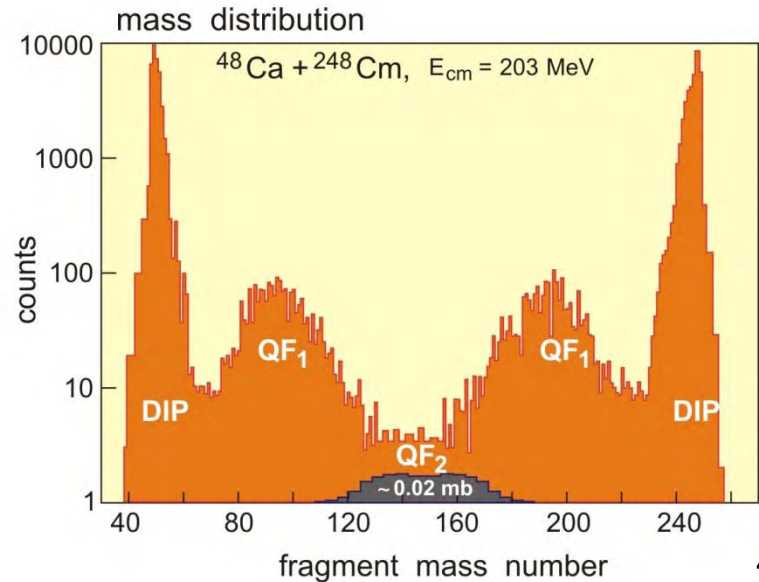
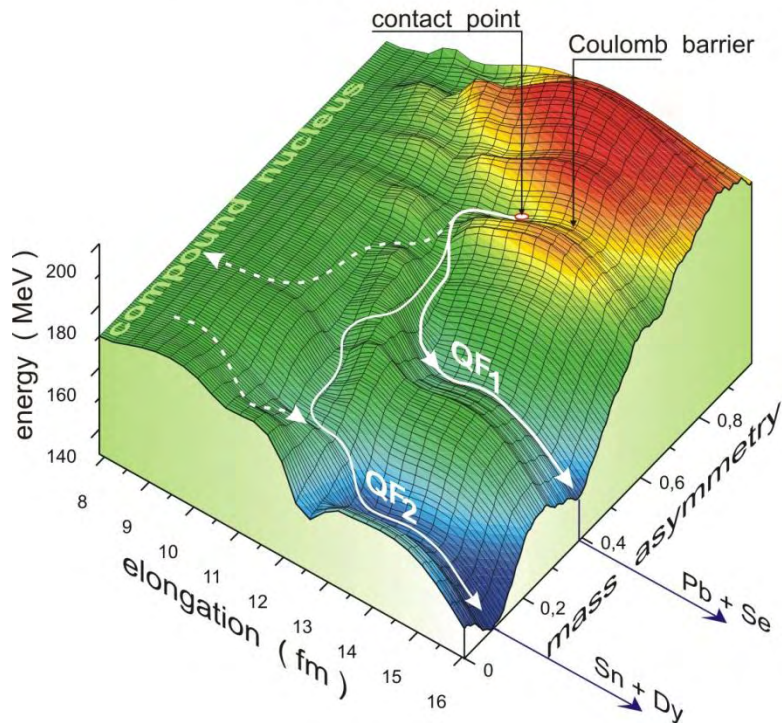
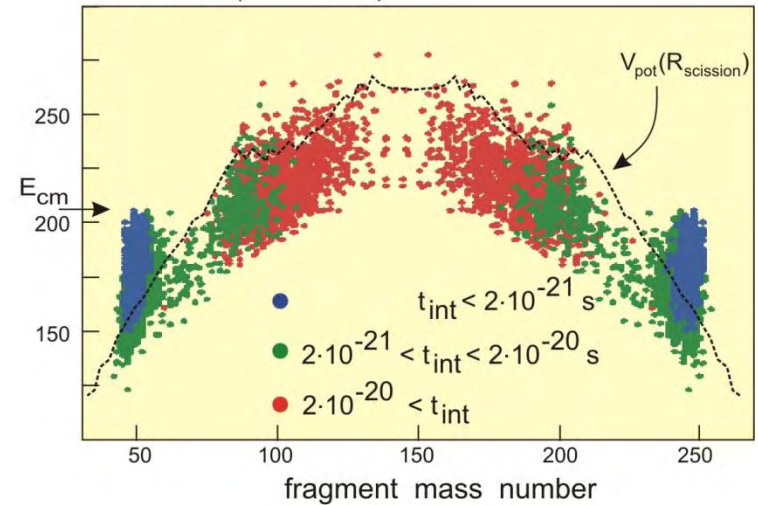
Quasi-Fission process is understood quite well

(example: $^{48}\text{Ca} + ^{248}\text{Cm}$)

experiment: M. Itkis et al., 2000

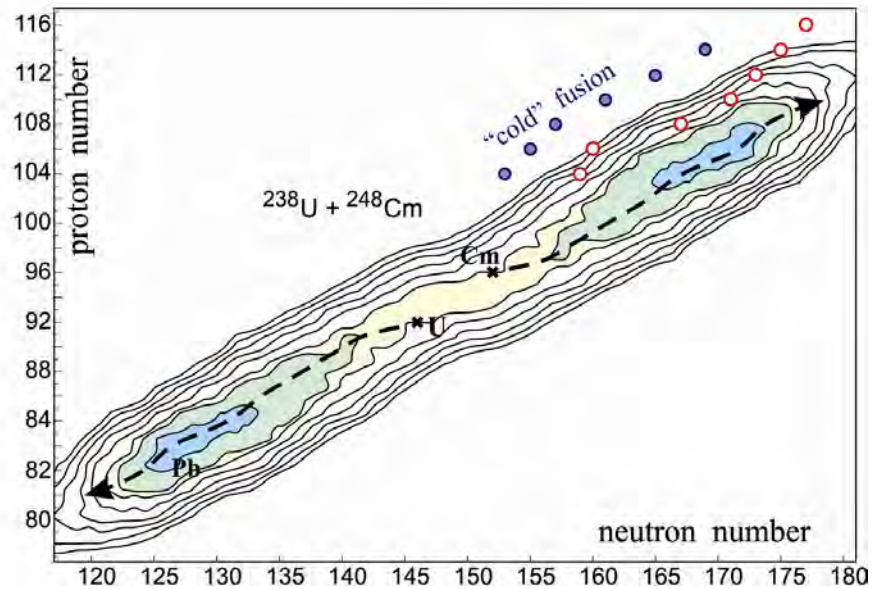
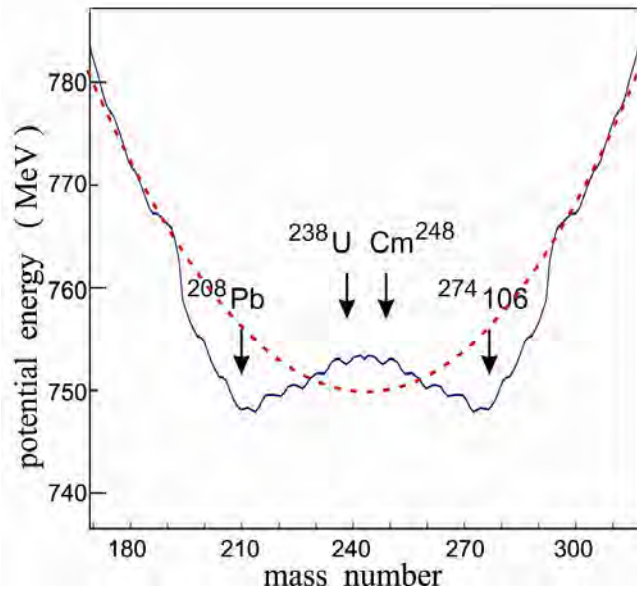
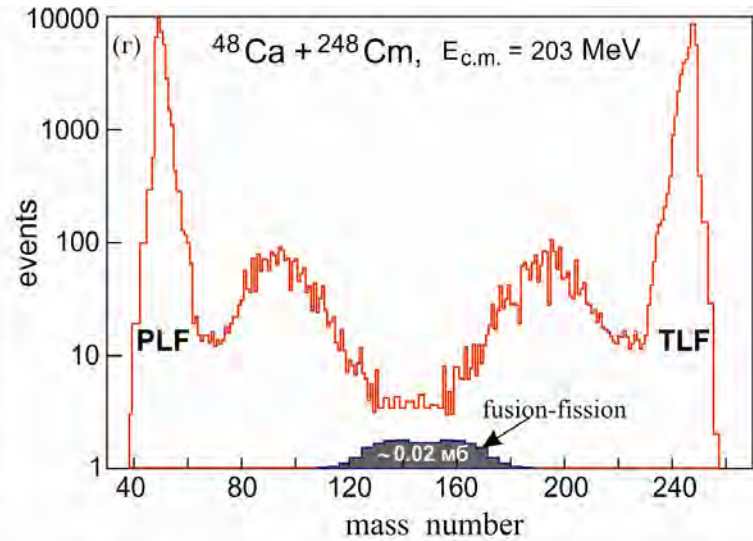
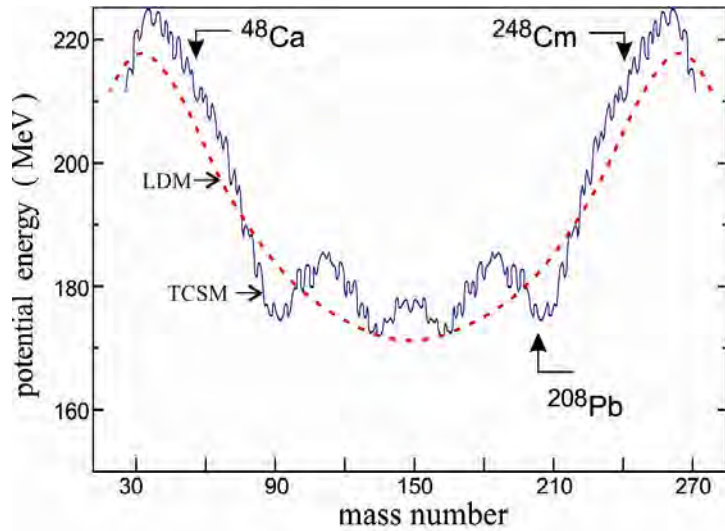


calculation (10^5 events)

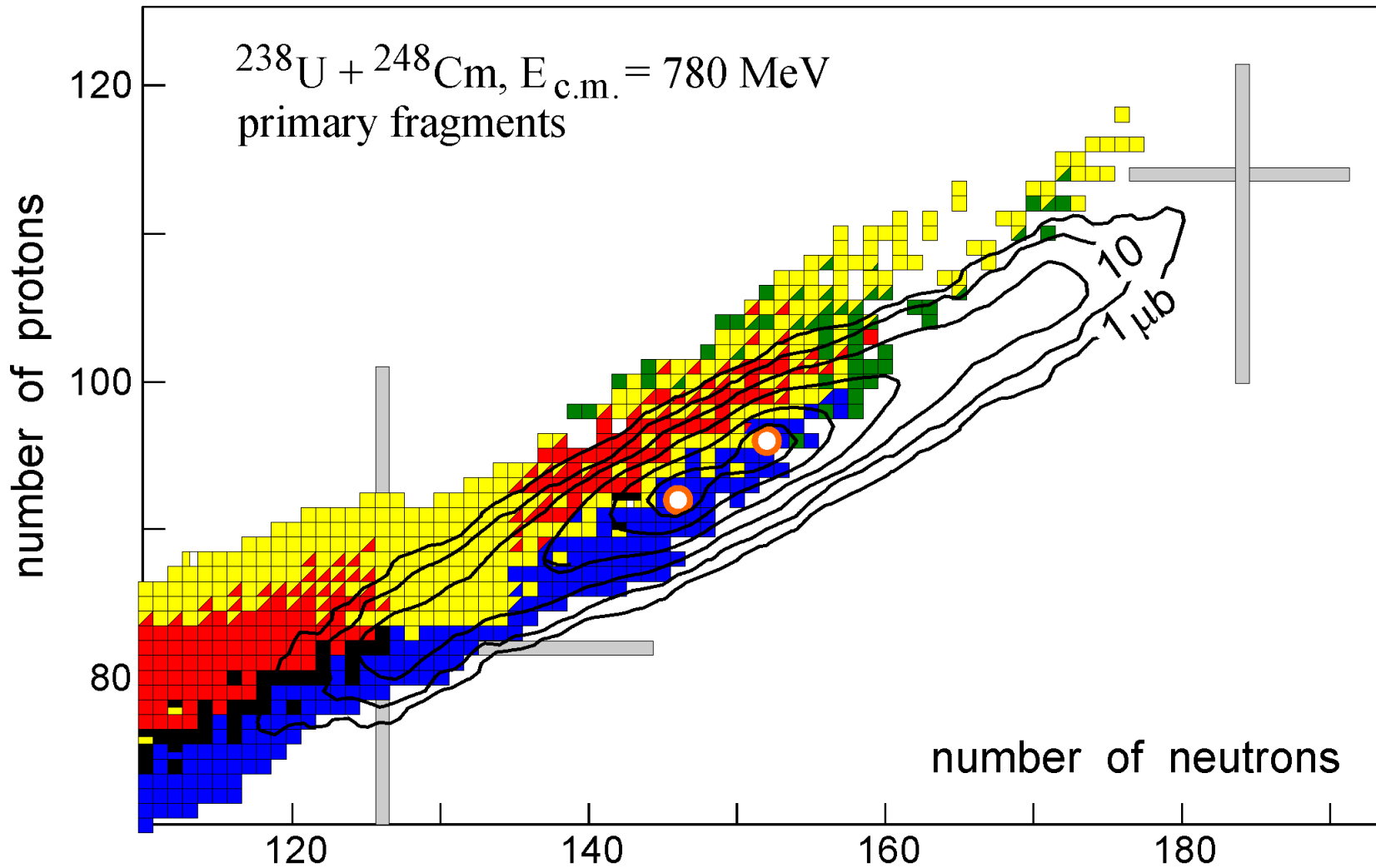


Shell effects: lead valley

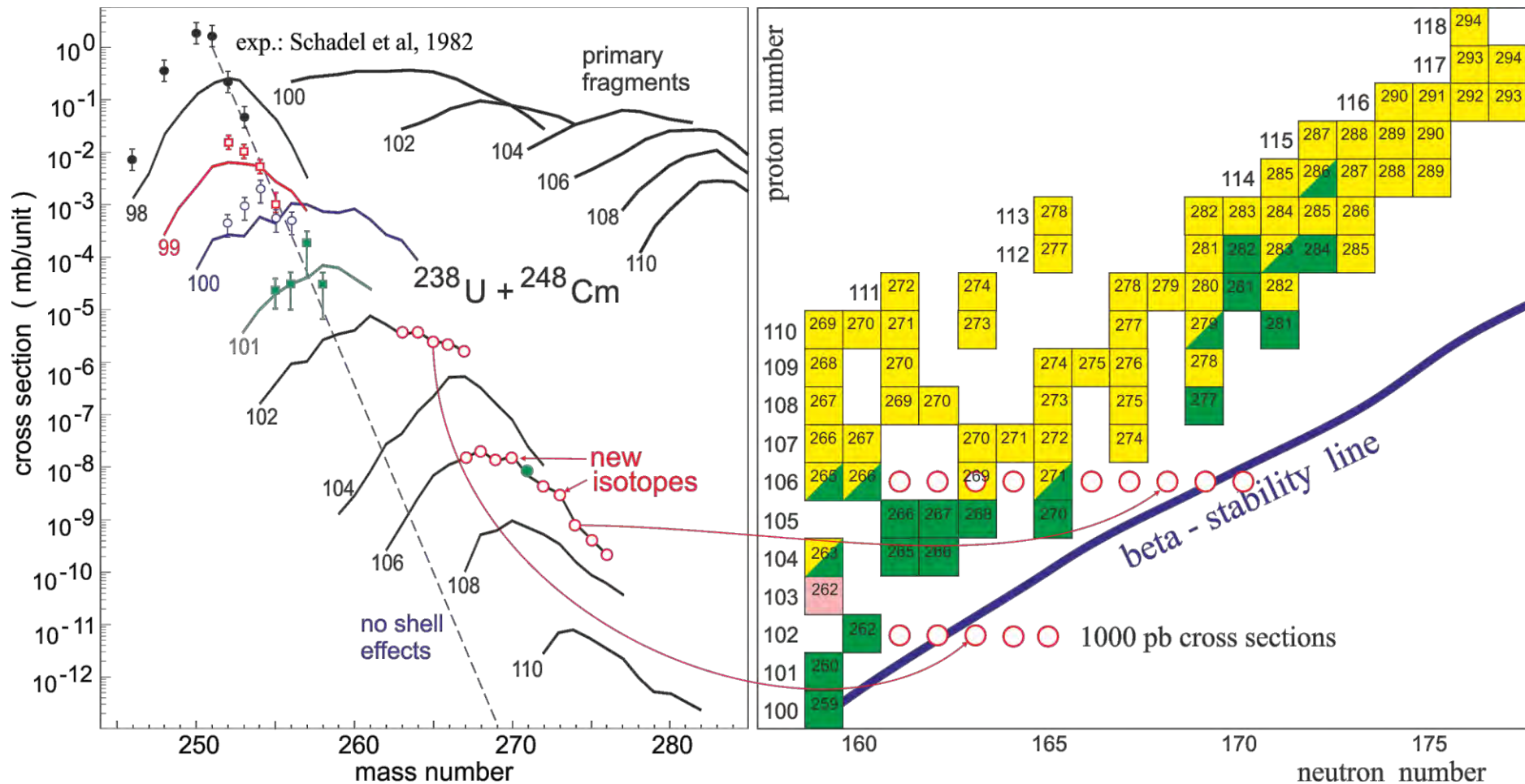
normal (symmetrizing) and inverse (anti-symmetrizing) quasi-fission



238U + 248Cm. Primary fragments

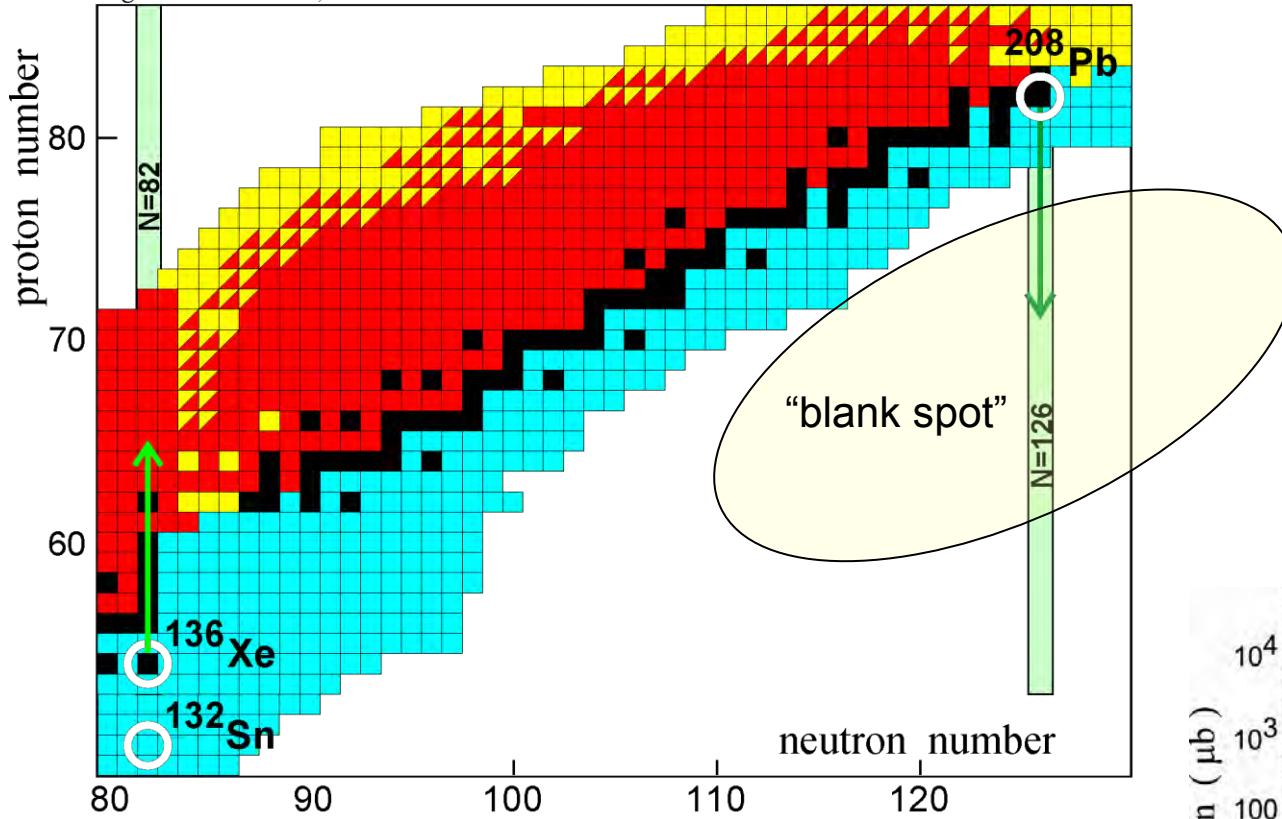


Production of transfermium nuclei along the line of stability looks quite possible in multi-nucleon transfer reactions

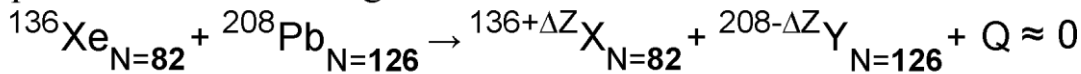


Production of new heavy nuclei in the region of N=126

Zagrebaev & Greiner, PRL 2008

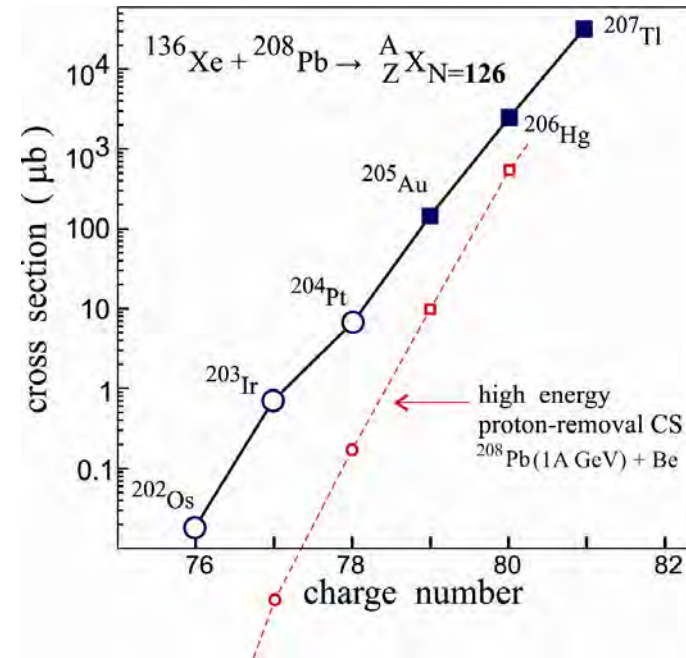


proton transfer along the neutron closed shells:

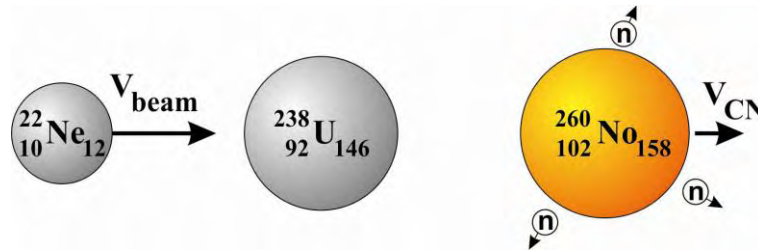


Reactions with $Q \approx 0$ are very favorable for proton transfer

The use of ${}^{132}\text{Sn}$ is even better !

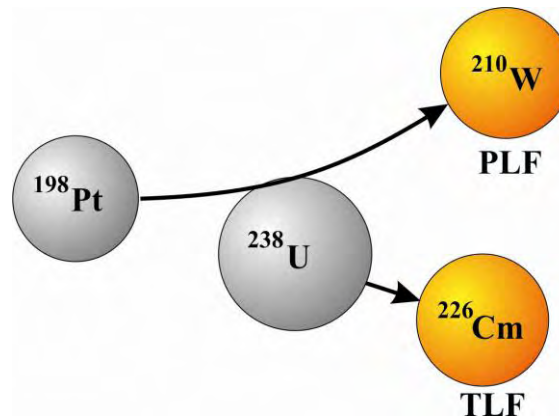


Fusion reactions



How to separate a given nucleus from all the other transfer reaction products ?

Transfer reactions



Available separators are not applicable !

Selective laser ionization ! (Au & Hg as an example)

Ionization Schemes

Au I

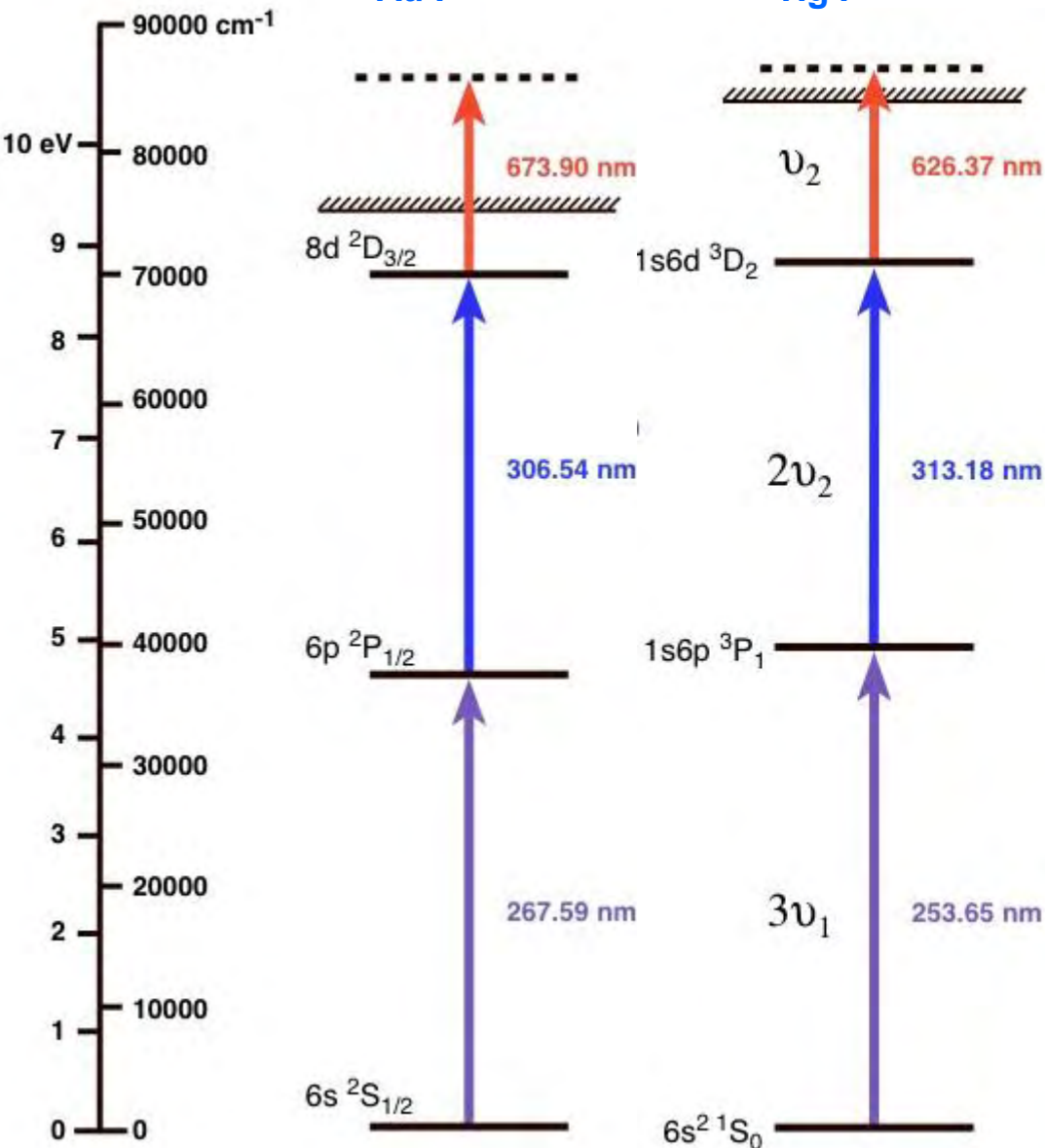
Level Energy, cm ⁻¹	Configuration	Wavelength, nm
E_0 0	$6s^2 2S_{1/2}$	λ_1 267.6
E_1 37358.99	$6p^2 P_{1/2}$	λ_2 306.5
E_2 69971.42	$8d^2 D_{3/2}$	λ_3 673.9

Chemical series: Transition metals
 Group, Period, Block: 11, 6, *d*
 Atomic mass: 196.96655(2) g/mol
 Electron configuration: [Xe] 4f¹⁴ 5d¹⁰ 6s
[Ionization potential](#): 74408.88 cm⁻¹ (9.22553 eV)

Hg I

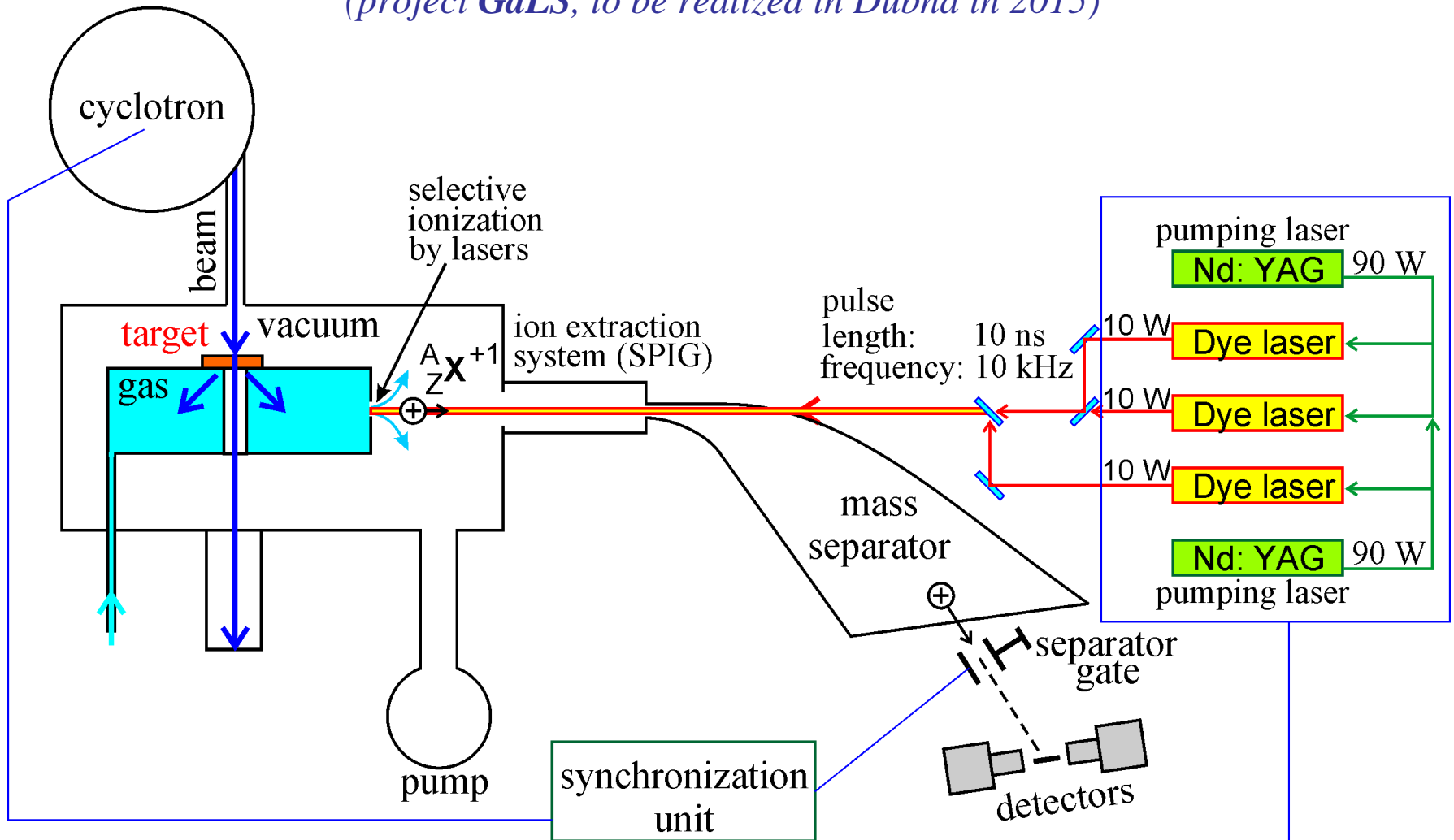
Level Energy, cm ⁻¹	Configuration	Wavelength, nm
E_0 0	$6s^2 1S_0$	λ_1 253.65
E_1 39412.30	$6s 6p^3 P_1$	λ_2 313.18
E_2 71396.22	$6s 6d^3 D_2$	λ_3 626.37

Chemical series: *d*
 Group, Period, Block: 12, 6, *d*
 Atomic mass: 200.59(2) g/mol
 Electron configuration: [Xe] 4f¹⁴ 5d¹⁰ 6s²
[Ionization potential](#): 84184.1 cm⁻¹ (10.4375 eV)

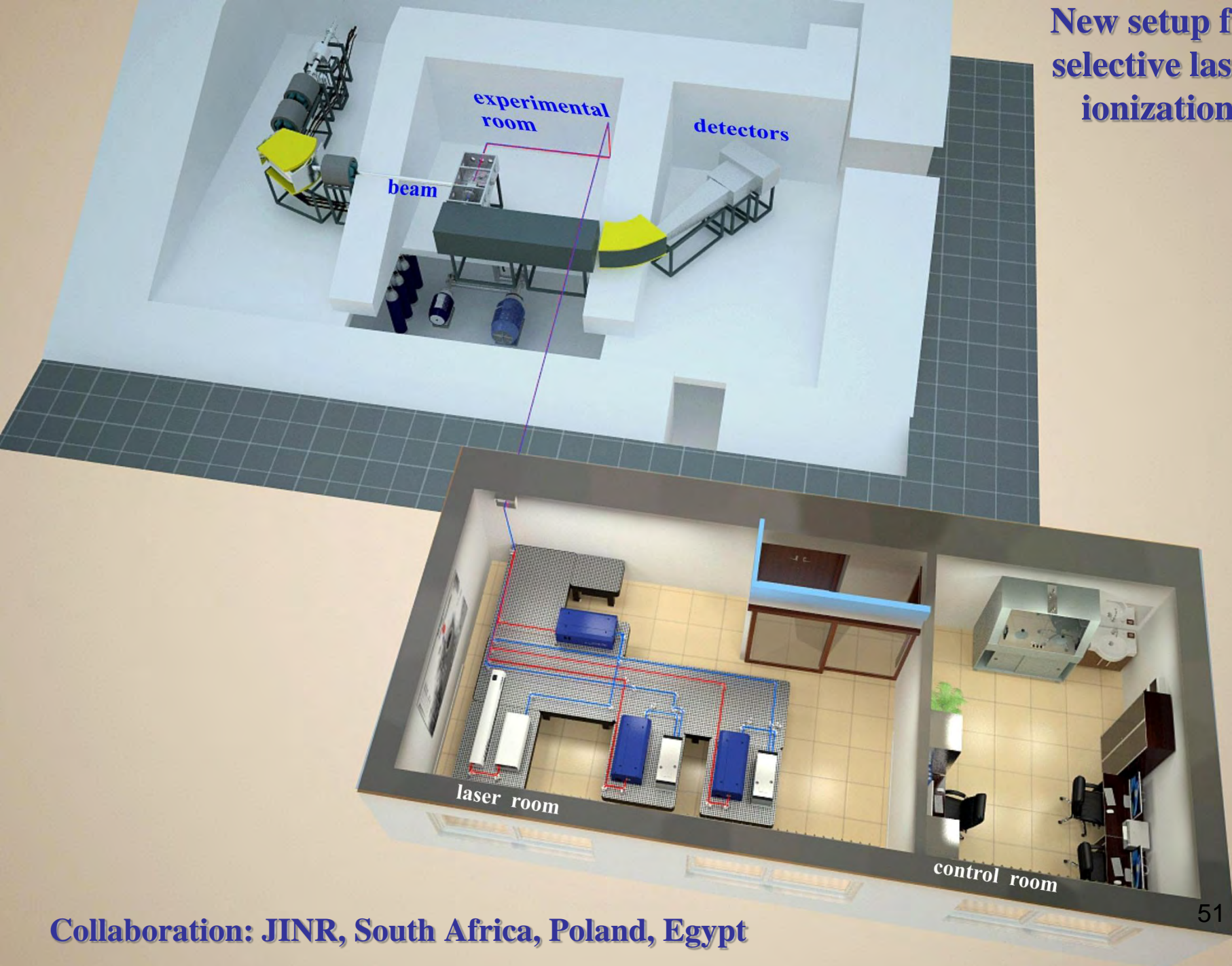


New setup for selective laser ionization of multi-nucleon transfer reaction products stopped in gas

(project *GaLS*, to be realized in Dubna in 2015)



New setup for selective laser ionization



Collaboration: JINR, South Africa, Poland, Egypt