Search for rare processes with DAMA/LXe experiment at Gran Sasso

R. Bernabei¹, P. Belli¹, F. Montecchia¹, F. Nozzoli¹, F. Cappella², A. Incicchitti², D. Prosperi², R. Cerulli^{3,a}, C.J. Dai⁴, V.Yu. Denisov⁵, and V.I. Tretyak⁵

¹ Dipartimento di Fisica, Università di Roma "Tor Vergata" and INFN, Sezione di Roma II, I-00133, Roma, Italy

² Dipartimento di Fisica, Università di Roma "La Sapienza" and INFN, Sezione di Roma, I-00185, Roma, Italy

³ INFN - Laboratori Nazionali del Gran Sasso, I-67010 Assergi (AQ), Italy

 $^4\,$ IHEP, Chinese Academy, P.O. Box 918/3, Beijing 100039, PRC $\,$

⁵ Institute for Nuclear Research, MSP 03680 Kiev, Ukraine

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Abstract. After a short introduction on the low background liquid xenon DAMA set-up (DAMA/LXe) and its main previous results, we discuss the search for the nucleon, di-nucleon and tri-nucleon decays into invisible channels (disappearance or decay to neutrinos, Majorons, etc.) in the ¹³⁶Xe isotope. The obtained limits (90% C.L.) on the lifetimes are: $\tau_n > 3.3 \cdot 10^{23}$ yr, $\tau_p > 4.5 \cdot 10^{23}$ yr, $\tau_{np} > 3.2 \cdot 10^{23}$ yr, $\tau_{pp} > 1.9 \cdot 10^{24}$ yr, $\tau_{nnp} > 1.4 \cdot 10^{22}$ yr, $\tau_{npp} > 2.7 \cdot 10^{22}$ yr and $\tau_{ppp} > 3.6 \cdot 10^{22}$ yr. In particular, the tri-nucleon decay into invisible channels is investigated here for the first time.

PACS. 29.40.Mc Scintillation detectors – 95.35.+d Dark matter (stellar, interstellar, galactic, and cosmological) – 11.30.Fs Global symmetries (*e.g.*, baryon number, lepton number)

1 Introduction on previous results

The DAMA/LXe experiment has realized several prototype detectors and, then, has preliminarily put in measurement the set-up used in the data taking of refs. [1,2]; this set-up was firstly upgraded in fall 1995 as mentioned in refs. [3,4,5,6,7]. In particular, it has an inner vessel filled by $\simeq 6.5$ kg (*i.e.* $\simeq 21$) of liquid xenon. Firstly it used Kr-free xenon enriched in 129 Xe at 99.5% [8]; then, in 2000 the set-up was deeply modified reaching the configuration in fig. 5 of ref. [8] to handle also Kr-free xenon enriched in 136 Xe at 68.8% [9,10]. In this latter case, the interest has mainly been focused on the higher energy region. The main features of the set-up, details on the data acquisition, on the cryogenic and vacuum systems and on the running parameters control are described in [6, 7, 8, 10, 11]. In particular, the energy scale in the high energy region is determined with the help of external standard gamma sources, having identified the possibility to perform high energy calibrations with external sources near a flange located on the top of the detector, opening a limited upper part of the external shield. The energy resolution for $^{208}\mathrm{Tl}$ line $(E_{\gamma} = 2614 \,\text{keV})$ is $\sigma \simeq 220 \,\text{keV}$, while for ¹³⁷Cs line $(E_{\gamma} = 662 \,\mathrm{keV})$, which is the source usually exploited for the routine calibrations, σ is $\simeq 70 \text{ keV} [10, 11, 7]$.

We pointed out the interest in using liquid xenon as target-detector for particle Dark Matter (DM) investigation deep underground since long time [12]. After preliminary measurements both on elastic and inelastic DM particles-¹²⁹Xe scattering [1,3], the recoil/electron light ratio and pulse shape discrimination capability in a similar pure LXe scintillator have been measured both with Am-B neutron source and with 14 MeV ENEA-Frascati neutron generator [6]. Moreover, in 2000/2001 further measurements on the recoil/electron light ratio with 2.5 MeV ENEA-Frascati neutron generator have also been carried out; see ref. [5] for details and comparisons.

After upgrading of the LXe set-up, new results on the DM particles investigation have been obtained [4,6]. In particular, in ref. [6] pulse shape discrimination between recoils and electromagnetic background in the developed pure LXe scintillators has been exploited. Afterwards the inelastic excitation of 129 Xe by DM particles with spin-dependent coupling has further been investigated in ref. [4].

Several other rare processes have also been searched for by means of the detector filled with the Kr-free xenon gas enriched in ¹²⁹Xe. In particular, as regards the electron stability, limits on the lifetime of the electron decay in both the disappearance and the $\nu_e + \gamma$ channels were set in ref. [2]. The latter has been more recently improved up to $2.0(3.4) \cdot 10^{26}$ yr at 90% (68%) C.L. [13]. Furthermore,

^a e-mail: riccardo.cerulli@lngs.infn.it

new lifetime limits on the charge non-conserving electron capture with excitation of $^{129}\mathrm{Xe}$ nuclear levels have also been established to be in the range $(1\text{-}4)\cdot10^{24}\,\mathrm{yr}$ at 90% C.L. for the different excited levels of $^{129}\mathrm{Xe}$ [14]. The stringent restrictions on the relative strengths of charge non-conserving (CNC) processes have been consequently derived: $\epsilon_W^2 < 2.2\cdot10^{-26}$ and $\epsilon_\gamma^2 < 1.3\cdot10^{-42}$ at 90% C.L. [14].

Moreover, we have searched for the nucleon and dinucleon decay into invisible channels in the 129 Xe isotope by exploiting a new approach [15]. It consisted in a search (in the real-time experiment) for radioactive decay of unstable daughter nuclei which were created in result of the N or NN disappearance in parent nucleus. If the half-life of the daughter nucleus is of the order of 1s or greater, such a decay will be time-resolved from prompt products if they were emitted and observed in a detector. This approach has the advantage of a branching ratio close to 1 and —if the parent and daughter nuclei are located in the detector itself— also of an efficiency close to 1. The obtained limits at 90% C.L. are: $\tau(p \rightarrow invisible \ channels) >$ $1.9 \cdot 10^{24}$ yr; $\tau(pp \rightarrow invisible \ channels) > 5.5 \cdot 10^{23}$ yr and $\tau(nn \rightarrow invisible \ channels) > 1.2 \cdot 10^{25} \, \text{yr}.$ These limits are similar to or better than those previously available; moreover, they are valid for every possible disappearance channel [15] and the limits for the di-nucleon decay in $\nu_{\tau}\bar{\nu}_{\tau}$ are set there for the first time.

Afterwards, measurements have been carried out by using the Kr-free xenon gas containing 17.1% of $^{134}\rm Xe$ and 68.8% of $^{136}\rm Xe$ to investigate the $^{134}\rm Xe$ and $^{136}\rm Xe$ double beta decay modes. After the preliminary results of refs. [9, 16] a joint analysis of the $0\nu\beta\beta$ decay mode in ¹³⁴Xe and in 136 Xe (as suggested in ref. [17]) has been carried out. New lower limits on various $\beta\beta$ decay modes have been obtained: for the $0\nu\beta\beta(0^+ \rightarrow 0^+)$ decay mode in ¹³⁴Xe and in ¹³⁶Xe the limits at 90% C.L. are $T_{1/2} = 5.8 \cdot 10^{22} \,\mathrm{yr}$ and $T_{1/2} = 1.2 \cdot 10^{24}$ yr, respectively; they correspond to a limit value on effective light Majorana neutrino mass ranging from 1.1 eV to 2.9 eV (90% C.L.), depending on the adopted theoretical model. For the neutrinoless double beta decay with Majoron (M) in the ¹³⁶Xe isotope the limit is $T_{1/2} = 5.0 \cdot 10^{23} \text{ yr}$ (90% C.L.); for the $2\nu\beta\beta(0^+ \rightarrow 0^+)$ and the $2\nu\beta\beta(0^+ \rightarrow 2^+)$ decay modes in ¹³⁶Xe the limits at 90% C.L. are $1.0 \cdot 10^{22} \text{ yr}$ and $9.4\cdot 10^{21}\,{\rm yr},$ respectively. It is worthy to note that the experimental limit on the $2\nu\beta\beta(0^+ \rightarrow 0^+)$ decay mode is in the range of the theoretical estimate by [18] $(2.11 \cdot 10^{22} \text{ yr})$ and about a factor 5 higher than that of ref. $[19]^1$.

A search for the charge non-conserving decay of ¹³⁶Xe into ¹³⁶Cs has also been performed for the first time [7], using the data collected during 8823.54 h and already published in ref. [10]. The used approach has been the investigation of the CNC processes by the search for the possible CNC decay firstly considered in [20]: if in a β decay $(A, Z) \rightarrow (A, Z + 1) + e^- + \overline{\nu}_e$ some massless uncharged particle would be emitted instead of the electron $(e.g., \nu_e \text{ or } \gamma \text{ or Majoron})$, an additional 511 keV energy release would occur. Thus, usually forbidden decays to the ground state or to the excited levels of the daughter nuclei would become energetically possible. The presence of the (A, Z + 1) isotope or of its daughter products in a sample, initially free from them, would indicate the existence of the CNC decay searched for. Large advantages arise when the so-called "active-source" technique (source = detector) is considered as in the case described here. In particular, after the possible 136 Xe CNC decay, the daughter nucleus $^{136}\mathrm{Cs}$ will be created. It is β unstable $(T_{1/2} = 13.16 \,\mathrm{d})$ with quite high energy release $(Q_{\beta} = 2.548 \,\mathrm{MeV})$ [7]. Comparing the experimental energy distribution with the expected response function, no evidence for the effect searched for has been found. Thus, the lifetime limit is: $\tau_{CNC}(^{136}\text{Xe} \rightarrow ^{136}\text{Cs}) > 1.3 \cdot 10^{23} \,\text{yr}$ at 90% C.L. This limit is one of the highest available limits for similar processes [7]; however, the bound on the charge non-conserving admixture in the weak interactions, which can be derived according to ref. [21], is modest mainly due to the big change in nuclear spin in the considered CNC transition $(\Delta J^{\Delta \pi} = 5^+).$

2 The search for nucleon instabilities into invisible channels in the ¹³⁶Xe isotope

The baryon number (B) conservation was introduced by Stuckelberg [22] and Wigner [23] more than fifty years ago and remains an accidental symmetry of the Standard Model, when this is seen as a renormalizable theory, not explained by deeper theoretical understanding. Modern theories of particle physics (GUTs, SUSY), unifying quarks and leptons into the same multiplets and predicting new interactions which transform quarks into leptons, naturally lead to the decay of the protons and the otherwise stable bound neutrons [24]. Many decay mechanisms, which violate B on one or two units, have been discussed [24, 25]. In a recent work [26], a new process was examined in which two neutrons simultaneously decay to bulk Majoron, $nn \to M$, with typical lifetime 10^{32} - 10^{39} yr; due to weak coupling of Majorons to normal matter they are not detected in an experiment, and such a process looks as a disappearance of two neutrons from a nucleus. Also mechanisms for the tri-nucleon decay have been proposed in the literature; in particular, very recently in ref. [27] also tri-nucleon decay processes with $\Delta B = 3$ have been considered. Moreover, disappearance of particles (electrons, e^- , or nucleons, N) are expected also in theories with extra dimensions [28, 29, 30]. No processes with baryon number violation were detected to date. We refer to our previous article [15] for a review of various ideas used in the search for the N and NN decays into invisible channels.

Here a search for the N, NN and NNN instabilities in the ¹³⁶Xe isotope is described, looking for decays of the unstable daughter nuclei. The experimental energy distribution collected during 8823.54 h by the LXe scintillator (enriched in ¹³⁶Xe at 68.8%) in the energy region 550– 3550 keV (the same as in refs. [10] and [7]) is shown later.

¹ On the other hand, similar theoretical estimates suffer from the large uncertainties typically associated to the calculations of the nuclear matrix elements.

Table 1. Daughter nuclei produced in N, NN and NNN decays in ¹³⁶Xe when the de-excitation of the daughter nucleus occurs by γ emission. The half-life times of the isotopes involved in the decay chains vary from 2.5 m (¹³³Sb) to 5.243 d (¹³³Xe) assuring that the chains are in equilibrium and that subsequent decays are well separated in time.

Decay	Daughter nucleus	Subsequent decays
n	$^{135}\mathrm{Xe}$	135 Xe $\xrightarrow{\beta^-} ^{135}$ Cs *
p	135 I	${}^{135}\mathrm{I} \xrightarrow{\beta^-} {}^{135}\mathrm{Xe} \xrightarrow{\beta^-} {}^{135}\mathrm{Cs} \ ^*$
nn	$^{134}\mathrm{Xe}$	Stable
np	134 I	$^{134}\text{I} \xrightarrow{\beta^{-}} ^{134}\text{Xe}$
pp	$^{134}\mathrm{Te}$	$^{134}\text{Te} \xrightarrow{\beta^-} {}^{134}\text{I} \xrightarrow{\beta^-} {}^{134}\text{Xe}$
nnn	$^{133}\mathrm{Xe}$	133 Xe $\xrightarrow{\beta^-}$ 133 Cs
nnp	133 I	${}^{133}\mathrm{I} \xrightarrow{\beta^-} {}^{133}\mathrm{Xe} \xrightarrow{\beta^-} {}^{133}\mathrm{Cs}$
npp	$^{133}\mathrm{Te}$	${}^{133}\text{Te} \xrightarrow{\beta^-} {}^{133}\text{I} \xrightarrow{\beta^-} {}^{133}\text{Xe} \xrightarrow{\beta^-} {}^{133}\text{Cs}$
ppp	$^{133}\mathrm{Sb}$	${}^{133}\mathrm{Sb} \xrightarrow{\beta^-} {}^{133}\mathrm{Te} \xrightarrow{\beta^-} {}^{133}\mathrm{I} \xrightarrow{\beta^-} {}^{133}\mathrm{Xe} \xrightarrow{\beta^-} {}^{133}\mathrm{Cs} **$

* ¹³⁵Cs is not stable, but has $T_{1/2} = 2.3 \cdot 10^6$ yr and breaks the decay chain.

** Given here only the main part of the chain.

We remind that a cut has been applied to reject events whose charge is dominant in one PMT, that is to reject —as much as possible— background contribution from outside the inner vessel. As a consequence, a correction (which depends on the used hardware cut in high energy measurements) estimated by Monte Carlo program, properly considering the real geometry and the features of the detector, has been applied to the rate.

The isotopes given in table 1 are produced [31] after the disappearance of one, two or three nucleons in the parent $^{136}_{54}$ Xe nucleus, when the daughter de-excitation occurs only by γ emission. In general, the created daughter nucleus —with one, two or three holes in nuclear shells due to disappeared nucleons— will be in an excited state, unless the nucleons were on the outermost shell. The holes will be filled in the subsequent de-excitation process in which different particles could be emitted. If the excitation energy is lower than the binding energy of the least bound nucleon, only γ quanta will be emitted; otherwise, heavy particles such as p, n or α will be ejected leading to daughter nuclei with lower atomic masses and numbers. In that follows we will take into account the N, NN and NNN disappearance from a few outermost shells in the parent nucleus, when only γ 's would be emitted, avoiding in this way uncertainty in created daughter nuclide.

The lifetimes for the N, NN and NNN disappearance in 136 Xe can be calculated using the formula:

$$\tau = \frac{\epsilon_{\Delta E} \cdot N_{\text{nucl}} \cdot N_{\text{obj}}^{\text{eff}} \cdot T}{S_{\Delta E}}, \qquad (1)$$

where $N_{\text{nucl}} = 2.00 \cdot 10^{25}$ is the number of ¹³⁶Xe nuclei; T = 8823.54 h is the time of measurement; $N_{\text{obj}}^{\text{eff}}$ is the "effective" number of objects (n, p, NN) pairs or NNN

Table 2. Values of $N_{\rm obj}^{\rm eff}$ used in the present data analysis to search for N, NN and NNN disappearance in ¹³⁶Xe. See ref. [15] and text.

Decay	n	p	np	pp	nnn	nnp	npp	ppp
$N_{\rm obj}^{\rm eff}$	32	26	2	7	1	1	1	1

groups) whose disappearance in the parent nucleus will result in the creation of the specific daughter nuclide; $S_{\Delta E}$ is the number of events which can be ascribed to the decay process searched for in the considered energy window, ΔE , while $\epsilon_{\Delta E}$ is the related detection efficiency.

Equation (1) requires the knowledge of the number of objects, $N_{\rm obj}^{\rm eff}$, that was calculated following the method already used in ref. [15]. As regards the NNN decays into invisible channels, there is no available information from previous search and, in particular, there are no reference criteria to evaluate $N_{\rm obj}^{\rm eff}$ in these cases. Thus, in the search we have carried out, we cautiously assumed $N_{\rm obj}^{\rm eff} = 1$ for all the NNN processes (nnn, nnp, npp and ppp), as done in the search for the NN decays into invisible channels in ref. [32]. The used values of $N_{\rm obj}^{\rm eff}$ are summarized in table 2. Referring to table 1, we note that, except for the $^{134}\mathrm{Xe}$ nucleus which is stable and does not allow us to search for the nn disappearance, all other daughter nuclei are radioactive. In the subsequent decays (third column of table 1), if we exclude the 135 Cs which has $T_{1/2} = 2.3 \cdot 10^6$ yr and breaks the decay chain, the half-lives of the involved nuclides are relatively small. This ensures equilibrium in chains and thus equal number of decays for 135 I and 135 Xe (in case of the *p* disappearance), for 134 Te and 134 I (pp decay), etc. The expected response functions of the LXe detector for the β^- decays of the nuclei involved in the decay chains were simulated with the help of the EGS4 package [33]. The whole schemes of the decays, using the information from ref. [34] for the A = 133isotopes, from ref. [35] for the A = 134 isotopes and from ref. [36] for the A = 135 isotopes, were implemented in an event generator which described the initial kinematics of the events. The response functions for the N, NN and NNN disappearance are given by a linear combination of the response functions obtained for the single decays of the generated decay chain. Comparison of the experimental spectrum with the calculated response functions gives no strong indication of the signals searched for. Thus, we limit ourselves to extract only the limits on the probability of these processes. We will see that, except for the np and pp channel, the calculated response function do not present any distinctive structure, as a peak, at energies above the experimental energy threshold $(550 \, \text{keV})$ but their shapes are similar to some extent to that of the measured spectrum. As a consequence, the limit on the amplitude of the expected response function, which can be hidden in the experimental data, was determined in a very cautious and simple way (used also in the investigation of other rare processes). In fact, it has been required that —in no energy interval, ΔE — the number of events which could be ascribed to the investigated process, $S_{\Delta E}$,



Fig. 1. Comparison between the experimental spectrum measured during 8823.54 h (thick histogram) and the expected signal (colored histogram) for: left) the *n* disappearance with $\tau_n = 3.3 \cdot 10^{23}$ yr excluded at 90% C.L.; right) the *p* disappearance with $\tau_p = 4.5 \cdot 10^{23}$ yr excluded at 90% C.L.

can exceed the number of measured events plus $m \times \sigma$ (where *m* determines the C.L.). It is evident that the derived limits are very conservative because at least some events in the experimental spectrum can be induced by residual radioactive contaminations.

2.1 Results for N decays into invisible channel

A possible *n* disappearance in the ¹³⁶Xe nucleus is followed by the β^- decay ¹³⁵Xe \rightarrow ¹³⁵Cs ($T_{1/2} = 9.14$ h, $Q_{\beta^-} = 1151$ keV) in the sensitive volume of DAMA/LXe. The expected energy distribution is mainly determined by the β^- decay with end-point ~ 900 keV (96% of the cases) followed by a 250 keV de-excitation γ (the β^- energy distribution is moved towards a higher energy of 250 keV) [11]. In the data analysis we considered the 550–700 keV energy region of the experimental spectrum (see fig. 1 left), where 369 events were accumulated. It gives $S_{\Delta E} < 369 + 1.29 \cdot \sqrt{369} = 393.6$ at 90% C.L. The related efficiency, calculated with the EGS4, is $\varepsilon_{\Delta E} = 20.2\%$. Substituting these values into eq. (1) together with the effective number of neutrons which should be taken into account ($N_n^{\rm eff} = 32$), we obtain the following restriction on the neutron lifetime:

$$\tau_n > 3.3 \cdot 10^{23} \,\mathrm{yr} \,\mathrm{at} \,90\% \,\mathrm{C.L.}$$
 (2)

The p disappearance in the ¹³⁶Xe nucleus is instead followed by a chain of two β^- decays: ¹³⁵I \rightarrow ¹³⁵Xe $(T_{1/2} = 6.57 \text{ h}, Q_{\beta^-} = 2648 \text{ keV})$ and ¹³⁵Xe \rightarrow ¹³⁵Cs (see above). The expected energy distribution for the ¹³⁵I $\beta^$ decay is characterized by a peak due to the \sim 527 keV metastable state of ¹³⁵Xe $(T_{1/2} = 15.29 \text{ m})$, but below the energy threshold considered in this data analysis; the response function for the p disappearance is given by the ¹³⁵I + ¹³⁵Xe distribution [11]. The most sensitive energy region of the experimental spectrum is 550–600 keV (see fig. 1 right). For 131 accumulated events, the 90% C.L. limit is $S_{\Delta E} < 145.7$; taking into account the calculated efficiency $\varepsilon_{\Delta E} = 12.7\%$ to detect the ¹³⁵I + ¹³⁵Xe decays in the 550–600 keV interval and the effective number of protons $N_p^{\text{eff}} = 26$, the corresponding limit for the proton lifetime is:

$$\tau_p > 4.5 \cdot 10^{23} \,\mathrm{yr} \,\mathrm{at} \,90\% \,\mathrm{C.L.}$$
 (3)



Fig. 2. Comparison between the experimental spectrum measured during 8823.54 h (thick histogram) and the expected signal (colored histogram) for: left) the np disappearance with $\tau_{np} = 4.0 \cdot 10^{22}$ yr excluded at 90% C.L.; right) the pp disappearance with $\tau_{pp} = 2.1 \cdot 10^{23}$ yr excluded at 90% C.L. In the insets the residuals between the experimental spectrum and the background fits are shown together with excluded distributions for np ($\tau_{np} = 3.2 \cdot 10^{23}$ yr) and pp ($\tau_{pp} = 1.9 \cdot 10^{24}$ yr) disappearances, respectively. See text.

2.2 Results for NN decays into invisible channel

A possible np disappearance in the ¹³⁶Xe nucleus can be investigated by searching for the following β^- decay ¹³⁴I \rightarrow ¹³⁴Xe ($T_{1/2} = 52.5 \text{ m}, Q_{\beta^-} = 4175 \text{ keV}$) in the sensitive volume of DAMA/LXe. Considering the expected energy distribution calculated for this decay, the most sensitive energy region of the experimental spectrum (see fig. 2 left) is 2250–2300 keV interval with 14 events collected: $S_{\Delta E} < 18.8$ events at 90% C.L. With the related efficiency $\varepsilon_{\Delta E} = 1.86\%$ and the effective number of pairs $N_{np}^{\text{eff}} = 2$, we obtain the restriction on the np lifetime: $\tau_{np} > 4.0 \cdot 10^{22} \text{ yr (90\% C.L.)}.$

The result of a pp disappearance in the ¹³⁶Xe nucleus is instead the creation of a ¹³⁴Te isotope and the subsequent β^- decays chain: ¹³⁴Te \rightarrow ¹³⁴I ($T_{1/2} = 41.8 \text{ m}$, $Q_{\beta^-} =$ 1550 keV) and ¹³⁴I \rightarrow ¹³⁴Xe (see above). The expected energy distribution for these processes in DAMA/LXe sensitive volume gives in particular for the ¹³⁴Te β^- decay a bump at ~ 1000 keV due to events where the de-excitation γ 's, following the β^- ray with end-point ~ 730 keV (42% of the cases), ~ 610 keV (44%) and ~ 450 keV (14%), are fully contained in the detector, and a structure with energy below ~ 800 keV associated with events where one or more γ 's escape the detector. For the pp disappearance the expected signal is given by the ¹³⁴Te + ¹³⁴I distribution and the considered region of the experimental spectrum (see fig. 2 right) is 1100–1150 keV with 35 events. With the $S_{\Delta E} < 42.6$, $\varepsilon_{\Delta E} = 6.47\%$, and $N_{pp}^{ep} = 7$, it results in the value $\tau_{pp} > 2.1 \cdot 10^{23}$ yr at 90% C.L.

However, when the experimental data have a smooth behaviour and the expected response function of the effect has some peculiarities, as the peak-like structure near 1 MeV for the $^{134}\text{Te} + ^{134}\text{I}$ decay (*pp* disappearance), it is justified to use an approach that allows to take into account the background contribution when estimating the limit on the lifetime. In particular, the experimental spectrum can be fitted by some appropriate background model together with the effect's response function with parameters of the model and the amplitude of the effect being the free parameters of the fit. The fit of the experimental



Fig. 3. Comparison between the experimental spectrum measured during 8823.54 h (thick histogram) and the expected signal (colored histogram) for: left) the *nnp* disappearance with $\tau_{nnp} = 1.4 \cdot 10^{22}$ yr excluded at 90% C.L.; center) the *npp* disappearance with $\tau_{npp} = 2.7 \cdot 10^{22}$ yr excluded at 90% C.L.; right) the *ppp* disappearance with $\tau_{ppp} = 3.6 \cdot 10^{22}$ yr excluded at 90% C.L.

spectrum in the 550–3200 keV region by the sum of an exponent and a straight line, as background model², and the ¹³⁴Te + ¹³⁴I decay response function gave for its area the value $S = 0.0 \pm 44.2$, providing no evidence for the effect being searched; the corresponding $\chi^2/\text{n.d.f.}$ value was equal to 1.1. According to the Feldman-Cousins procedure [37], the 90% C.L. limit on the number of observed ¹³⁴Te + ¹³⁴I events is S < 72.5. This value, together with the effective number of the *pp* pairs $N_{pp}^{\text{eff}} = 7$, results in the *pp* lifetime limit:

$$\tau_{pp} > 1.9 \cdot 10^{24} \,\mathrm{yr} \,\mathrm{at} \,90\% \,\mathrm{C.L.}$$
 (4)

In the same way the limit on the number of events under the 134 I distribution (*np* disappearance) was determined. The fit (χ^2 /n.d.f. = 1.1) gave for the *S* value $S = 19.3 \pm 65.3$ that resulted in the bound S < 126.4 at 90% C.L. Taking into account that $N_{np}^{\text{eff}} = 2$, we obtain the *np* lifetime limit:

$$\tau_{np} > 3.2 \cdot 10^{23} \,\mathrm{yr} \,\mathrm{at} \,90\% \,\mathrm{C.L.}$$
 (5)

The distributions corresponding to a lifetime equal to these limit values for the np and pp disappearances in ¹³⁶Xe are shown together with the residuals between the experimental spectrum and the background model in the insets of fig. 2 left and right, respectively.

2.3 Results for NNN decays into invisible channel

The result of a $nnn \rightarrow invisible \ channels$ decay in the ¹³⁶Xe nucleus is the creation of the ¹³³Xe isotope, which is unstable and β^- decays to ¹³³Cs (stable) with $T_{1/2} = 5.243$ d and $Q_{\beta^-} = 427.4$ keV. As evident, the maximum energy released in this process is below the 550 keV energy threshold of the present measurements; thus, this decay process cannot be investigated here.

As a consequence of a *nnp* disappearance in the ¹³⁶Xe nucleus, a ¹³³I nucleus is instead created in the sensitive volume of DAMA/LXe. This isotope is the parent of the β^- decay chain ¹³³I \rightarrow ¹³³Xe \rightarrow ¹³³Cs. The energy distribution expected for the ¹³³I β^- decay ($T_{1/2} = 20.8$ h and $Q_{\beta^-} = 1770 \text{ keV}$) in our liquid xenon set-up is mainly determined by the β^- decay with end-point $\sim 1240 \text{ keV}$ (83% of the cases) followed by a 530 keV de-excitation γ

(the β^- energy distribution is moved towards higher energy of 530 keV); a 233 keV peak is instead due to the deexcitation of the ¹³³Xe metastable state ($T_{1/2} = 2.19 \,\mathrm{d}$), which is reached by the ~ 3% of the decays [11]. The signal to be searched for in case of a *nnp* disappearance in ¹³⁶Xe is given by the ¹³³I+¹³³Xe β^- decay distributions but the last one in the analysis does not play any role since it falls under the experimental energy threshold (550 keV).

The more selective region in the present case is $\Delta E = 1100-1150 \text{ keV}$ (see fig. 3 left); it contains 35 events, which gives rise to the upper limit (90% C.L.) $S_{\Delta E} < 42.6$ events and being $\epsilon_{\Delta E} = 3.0\%$ gives:

$$\tau_{nnp} > 1.4 \cdot 10^{22} \,\mathrm{yr} \,\mathrm{at} \,90\% \,\mathrm{C.L.}$$
 (6)

In the case of the process $npp \rightarrow invisible \ channels$ in the ¹³⁶Xe nucleus, a ¹³³Te nucleus is created. This isotope $(T_{1/2} = 12.5 \text{ m and } Q_{\beta^-} = 2920 \text{ keV})$ produces a β^- decay to ¹³³I followed by the decay chain ¹³³I \rightarrow ¹³³Xe \rightarrow ¹³³Cs already described above for the nnp channel. The simulation result for a ¹³³Te decay in the DAMA/LXe detector gives peculiar structures in the expected distribution determined by the γ 's emitted in the ¹³³I de-excitation, whose energies are summed to the initial β^- of ¹³³Te. The more probable emitted γ 's are those with energies 312 keV (62%) and 408 keV (27%), from the first two excited levels of $^{133}\mathrm{I.}$ In fact, a bump at \sim 300 keV is due to the γ at $312\,{\rm keV}$ which, in most cases, releases all its energy in the detector, while a *bump* at $\sim 700 \text{ keV}$ is due to the sum of the energies released by the two γ 's emitted in cascade. Each possible npp decay in 136 Xe is associated to a signal given by the sum of the energy distributions expected for the β^- decays of the ¹³³Te, of the ¹³³I and of the 133 Xe (this latter, as already mentioned, gives here a signal under the experimental energy threshold and, therefore, is not considered) [11]. Comparing the experimental spectrum measured (see fig. 3 center) and the expected signal, one gets that the more selective energy window is $\varDelta E = 1100\text{--}1150\,\mathrm{keV}$ which contains 35 events, giving rise to the upper limit: $S_{\Delta E} < 42.6$ events (90% C.L.). Taking into account the calculated efficiency $\epsilon_{\Delta E} = 5.7\%$, one can derive

$$\tau_{npp} > 2.7 \cdot 10^{22} \,\mathrm{yr} \,\mathrm{at} \,90\% \,\mathrm{C.L.}$$
 (7)

Finally, a possible $ppp \rightarrow invisible \ channels$ decay in ¹³⁶Xe will create in the liquid xenon the ¹³³Sb nucleus which β^- decays with $T_{1/2} = 2.5 \,\mathrm{m}$ and $Q_{\beta^-} = 4003 \,\mathrm{keV}$. As can be derived from the decay scheme of the ¹³³Sb, this

² Other parameterizations gave similar results.



Fig. 4. The full decay chain generated by the 133 Sb.

process gives rise to the production of ¹³³Te nuclei in the ground state (82.4%) and in the ¹³³mTe metastable state at 334 keV (17.6%). The full decay chain generated by the ¹³³Sb is summarized in fig. 4. With simple calculation one can get that a possible *ppp* disappearance in ¹³⁶Xe is followed by a sequence of β^- decays weighted according to the following formula:

$$1 \times {}^{133}\text{Sb} + 0.855 \times {}^{133}\text{Te} + 0.176 \times {}^{133\text{m}}\text{Te} + 1 \times {}^{133}\text{I} + 1 \times {}^{133}\text{Xe} .$$
(8)

The energy distributions expected for the 133 Te, 133 I and ¹³³Xe isotopes have been already described above. As regards the response function of DAMA/LXe for the β^{-} decay of ¹³³Sb, it is more difficult to identify the structure due to the γ 's emitted in the daughter nucleus deexcitation. In fact, the more probable γ 's have energies of the order of MeV and, therefore, they can escape the detector releasing only a part of their energy. The last process in (8) is the 133m Te decay ($T_{1/2} = 55.4$ m). The expected energy distribution for the $^{133\text{m}}$ Te decay process gives a 334 keV peak due to the de-excitation γ , which in the 17.5% of the cases move the 133m Te into the ground state of 133 Te; the remaining part of the expected signal (branching ratio = 82.5%) is instead given by the ^{133m}Te β^- decay ($Q_{\beta^-} = 2920 \text{ keV}$) to ¹³³I. In particular, the presence of the ¹³³I metastable state at 1634 keV with $T_{1/2} = 9 \,\mathrm{s}$ induces other peaks in the expected energy distribution (a 74 keV peak which can be well distinguished, but also peaks at 721 keV, 987 keV and 1634 keV) [11]. The signal expected for the $ppp \rightarrow invisible \ channels$ in ¹³⁶Xe can be obtained by summing according to (8) the obtained energy distributions [11] (we remind that the signal associated to the ¹³³Xe decay is below the experimental threshold 550 keV). Considering the expected energy distribution (see fig. 3 right), the most sensitive energy window is $\Delta E = 1100\text{--}1150 \text{ keV}$ which contains 35 events, giving rise to the upper limit: $S_{\Delta E} < 42.6$ events (90% C.L.); the detection efficiency is: $\epsilon_{\Delta E} = 7.6\%$. One obtains

$$\tau_{ppp} > 3.6 \cdot 10^{22} \,\mathrm{yr} \,\mathrm{at} \,90\% \,\mathrm{C.L.}$$
 (9)

3 Conclusion

The DAMA/LXe set-up, deeply improved several times, has allowed to achieve competitive results in the search for various rare processes as summarized in the first section of this paper.

Here in particular, the most recent results on nucleon decays into invisible channels in the 136 Xe isotope have

Table 3. Experimental limits (90% C.L.) obtained by DAMA/LXe on the lifetime values of N, NN and NNN decays into invisible channels in the ¹³⁶Xe isotope.

Decay	τ_{limit} years (90% C.L.)
n	$3.3\cdot10^{23}$
p	$4.5\cdot10^{23}$
nn	—
np	$3.2\cdot 10^{23}$
pp	$1.9\cdot 10^{24}$
nnn	-
nnp	$1.4 \cdot 10^{22}$
npp	$2.7 \cdot 10^{22}$
ppp	$3.6\cdot 10^{22}$

been discussed. The considered experimental approach assures a high detection efficiency and a branching ratio ~ 1 with respect to other different approaches that used very large mass installations to compensate for the much lower values for those quantities. In particular, *NNN* decays into invisible channels have been investigated here for the first time. The obtained results are summarized in table 3. All the limits achieved here are valid for every invisible decay channel, including disappearance in extra-dimensions or decay into particles which weakly interact with matter.

Further data taking is foreseen.

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References

- 1. P. Belli et al., Nuovo Cimento C 19, 537 (1996).
- 2. P. Belli et al., Astropart. Phys. 5, 217 (1996).
- P. Belli *et al.*, Phys. Lett. B **387**, 222 (1996); **389**, 783 (1996)(E).
- 4. R. Bernabei et al., New J. Phys. 2, 15.1 (2000).
- 5. R. Bernabei et al., EPJdirect C 11, 1 (2001).
- 6. R. Bernabei et al., Phys. Lett. B 436, 379 (1998).
- R. Bernabei *et al.*, in *Beyond the Desert 2003* (Springer, 2003) p. 365.
- R. Bernabei *et al.*, Nucl. Instrum. Methods A **482**, 728 (2002).
- 9. R. Bernabei et al., Phys. Lett. B 527, 182 (2002).
- 10. R. Bernabei *et al.*, Phys. Lett. B **546**, 23 (2002).
- F. Cappella, PhD Thesis, Università di Roma "Tor Vergata" (2005).
- 12. P. Belli et al., Nuovo Cimento A 103, 767 (1990).
- 13. P. Belli *et al.*, Phys. Rev. D **61**, 117301 (2000).
- 14. P. Belli et al., Phys. Lett. B 465, 315 (1999).
- 15. R. Bernabei et al., Phys. Lett. B 493, 12 (2000).
- R. Bernabei et al., in Technique and Application of Xenon Detectors (World Scientific Publ., 2002) p. 50.
- 17. F. Simkovic, P. Domin, A. Faessler, hep-ph/0204278.
- 18. A. Staudt et al., Europhys. Lett. 13, 31 (1990).
- 19. E. Caurier et al., Nucl. Phys. A 654, 973 (1999).
- G. Feinberg, M. Goldhaber, Proc. Natl. Acad. Sci. U.S.A. 45, 1301 (1959).

- J.N. Bahcall, Rev. Mod. Phys. 50, 881 (1978); Neutrino Astrophysics (Cambridge University Press, 1989) p. 359.
- 22. E.C.G. Stuckelberg, Helv. Phys. Acta 11, 225 (1938).
- 23. E.P. Wigner, Proc. Am. Philos. Soc. **93**, 521 (1949).
- 24. P. Langacker, Phys. Rep. **71**, 185 (1981).
- 25. C.E. Carlson, C.D. Carone, Phys. Lett. B 512, 121 (2001).
- 26. R.N. Mohapatra et al., Phys. Lett. B 491, 143 (2000).
- 27. K.S. Babu et al., Phys. Lett. B 570, 32 (2003).
- F.J. Yndurain, Phys. Lett. B **256**, 15 (1991);
 N. Arkani-Hamed *et al.*, Phys. Lett. B **429**, 263 (1998);
 N. Arkani-Hamed *et al.*, Phys. Today **55**, February issue, 35 (2002).
- S.L. Dubovsky *et al.*, Phys. Rev. D **62**, 105011 (2000);
 S.L. Dubovsky *et al.*, JHEP **08**, 041 (2000); V.A. Rubakov,
 Phys. Usp. **44**, 871 (2001).
- 30. S.L. Dubovsky, JHEP **01**, 012 (2002).
- R.B. Firestone, V.S. Shirley et al. (Editors), Table of Isotopes, 8th ed. (John Wiley & Sons, New York, 1996).
- 32. C. Berger et al., Phys. Lett. B 269, 227 (1991).
- 33. W.R. Nelson et al., SLAC-Report-265, Stanford, 1985.
- 34. S. Rab, Nucl. Data Sheets **75**, 491 (1995).
- 35. Yu.V. Sergeenkov, Nucl. Data Sheets 71, 557 (1994).
- 36. Yu.V. Sergeenkov et al., Nucl. Data Sheets 84, 115 (1998).
- 37. G.J. Feldman, R.D. Cousins, Phys. Rev. D 57, 3873 (1998).